



EXPLORE SCIENCE

NASA Astrophysics Technology Needs: Plans for Factoring in the Decadal Survey's Recommendations

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NASA Astrophysics 2021 Technology Plans

1. How will the recommendations from Astro2020 factor into APD's technology planning?
2. How can the community participate in the technology development and planning?

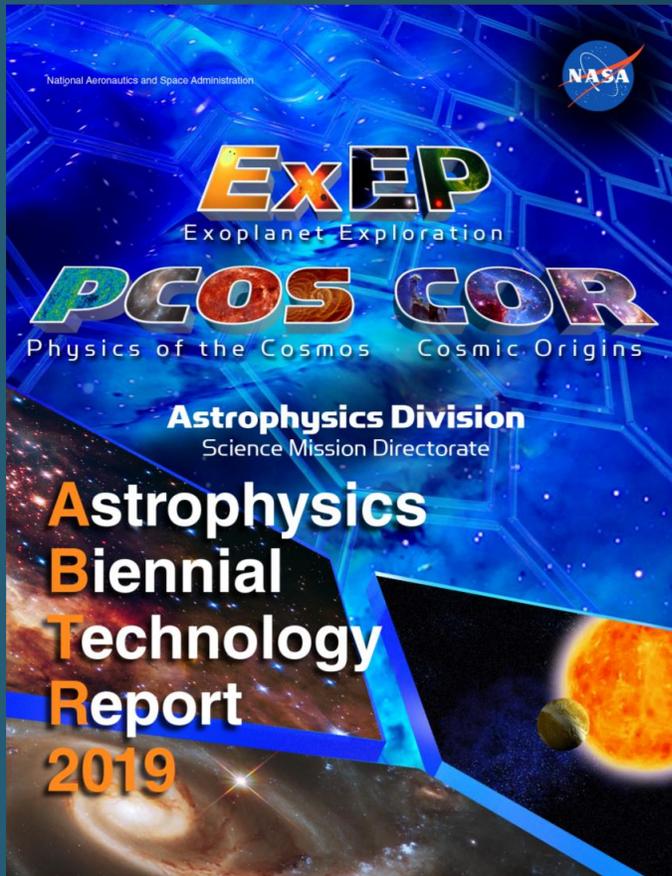
How will the recommendations from Astro2020 factor into APD's technology planning?

- **APD will respond to the Astro2020 Committee publicly on which of the recommendations it will commence first and expected timelines.**
 - ❑ *Preliminary response within a couple months of the released Astro2020 report*
 - ❑ *Completed Implementation Plan targeting Dec 2021*
- **A list of technology needs will then be updated and re-prioritized.**
- **APD will decide which technology development activities will be directed and competed.**
 - ❑ *A community-wide RFI is planned to be released to gather a wide range of information pertaining to the Astro2020 recommendations such as new innovations, approaches, architectures, and qualified sources.*
 - ❑ *The SAT solicitation can be part of a competed approach reflecting space mission priorities and longer-term technology development.*

How can the astrophysics community participate in technology planning and development?

- **Review our current technology gaps and propose new ones during the open solicitation period.**
 - ❑ *Technology Gaps and Technology Gap Form: <https://apd440.gsfc.nasa.gov/technology.html>*
 - ❑ *Jan 8 – June 1 (will defer if Astro2020 report is released later than mid-May)*
- **Respond to the community-wide RFI.**
 - ❑ *Released within two months of the public Astro2020 report*
- **Propose to the SAT element of the ROSES-2021 solicitation (*pending NASA's plans in response to the Astro2020 recommendations*).**
 - ❑ *ROSES-2021 solicitation ~ 2/14/21*
 - ❑ *Amendment to ROSES-2021 expected ~ 9/16/21 will confirm the SAT solicitation*
 - ❑ *If confirmed, NOI's due ~ 10/22/21 and proposals due ~ 12/16/21*

NASA Astrophysics Technology Needs



<https://apd440.gsfc.nasa.gov/technology.html>

Tier 1 Technology Gaps	
Angular Resolution (UV/Vis/NIR)	
Coronagraph Contrast	
Coronagraph Contrast Stability	
Cryogenic Readouts for Large-Format Far-IR Detectors	
Fast, Low-Noise, Megapixel X-Ray Imaging Arrays with Moderate Spectral Resolution	
High-Efficiency X-Ray Grating Arrays for High-Resolution Spectroscopy	
High-Resolution, Large-Area, Lightweight X-Ray Optics	
Large-Format, High-Resolution, UV/Vis Focal Plane Arrays	
Large-Format, High-Spectral-Resolution, Small-Pixel X-Ray Focal-Plane Arrays	
Large-Format, Low-Noise and Ultralow-Noise Far-IR Direct Detectors	
Large-Format, Low-Noise, High-QE Far-UV Detectors	
Next-Generation, Large-Format, Object Selection Technology for Multi-Object Spectrometers for LUVOIR	
Vis/NIR Detection Sensitivity	

Tier 2 Technology Gaps	
Advanced Millimeter-Wave Focal-Plane Arrays for CMB Polarimetry	
Detection Stability in Mid-IR	
Heterodyne FIR Detector Arrays and Related Technologies	
High-Efficiency Object Selection Technology for UV Multi-Object Spectrometers	
High-Performance Spectral Dispersion Component/Device	
High-Reflectivity Broadband FUV-to-NIR Mirror Coatings	
High-Throughput Bandpass Selection for UV/Vis	
Large-Format Object Selection Technology for Multi-Object Spectrometers for HabEx	
Starshade Deployment and Shape Stability	
Starshade Starlight Suppression and Model Validation	
Stellar Reflex Motion Sensitivity – Astrometry	
Stellar Reflex Motion Sensitivity – Extreme Precision Radial Velocity	

Tier 3 Technology Gaps	
Advanced Cryocoolers	
High-Performance, Sub-Kelvin Coolers	
Large Cryogenic Optics for the Mid-IR to Far-IR	
Long-Wavelength-Blocking Filters for X-Ray Micro-Calorimeters	
Low-Noise, High-QE UV Detectors	
Low-Stress, Highly Stable X-Ray Reflective Coatings	
Photon-Counting, Large-Format UV Detectors	
Polarization-Preserving Millimeter-Wave Optical Elements	
UV Coatings	
UV Detection Sensitivity	
UV/Vis/NIR Tunable Narrow-Band Imaging Capability	
Warm Readout Electronics for Large-Format Far-IR Detectors	

Tier 4 Technology Gaps	
Compact, Integrated Spectrometers for 100 to 1000 μm	
Optical-Blocking Filters	
Rapid Readout Electronics for X-Ray Detectors	
Short-Wave UV Coatings	

Tier 5 Technology Gaps	
Advancement of X-Ray Polarimeter Sensitivity	
Far-IR Spatio-Spectral Interferometry	
High-Precision Low-Frequency Radio Spectrometers and Interferometers	
Mid-IR Coronagraph Contrast	
Ultra-High-Resolution Focusing X-Ray Observatory Telescope	
Very-Wide-Field Focusing Instrument for Time-Domain X-Ray Astronomy	
Wide-Bandwidth, High-Spectral-Dynamic-Range Receiving System for Low-Radio-Frequency Observations on the Lunar Far Side	

- Technology gap descriptions and prioritization are published in the ABTR.
- A revised Report will come out ~ October incorporating Astro2020 recommendations.

Green: An STDT identified technology

NASA Astrophysics Technology Needs (an Example)

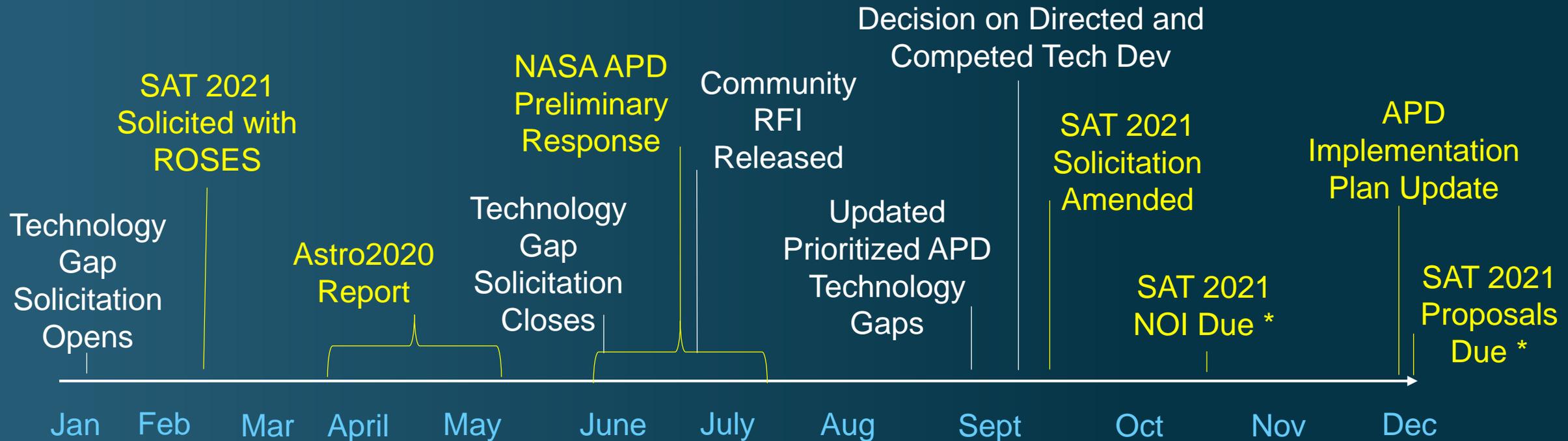
Tier 1 Technology Gaps
Angular Resolution (UV/Vis/NIR)
Coronagraph Contrast
Coronagraph Contrast Stability
Cryogenic Readouts for Large-Format Far-IR Detectors
Fast, Low-Noise, Megapixel X-Ray Imaging Arrays with Moderate Spectral Resolution
High-Efficiency X-Ray Grating Arrays for High-Resolution Spectroscopy
High-Resolution, Large-Area, Lightweight X-Ray Optics
Large-Format, High-Resolution, UV/Vis Focal Plane Arrays
Large-Format, High-Spectral-Resolution, Small-Pixel X-Ray Focal-Plane Arrays
Large-Format, Low-Noise and Ultralow-Noise Far-IR Direct Detectors
Large-Format, Low-Noise, High-QE Far-UV Detectors
Next-Generation, Large-Format, Object Selection Technology for Multi-Object Spectrometers for LUVOIR
Vis/NIR Detection Sensitivity

https://apd440.gsfc.nasa.gov/tech_gap_priorities.html

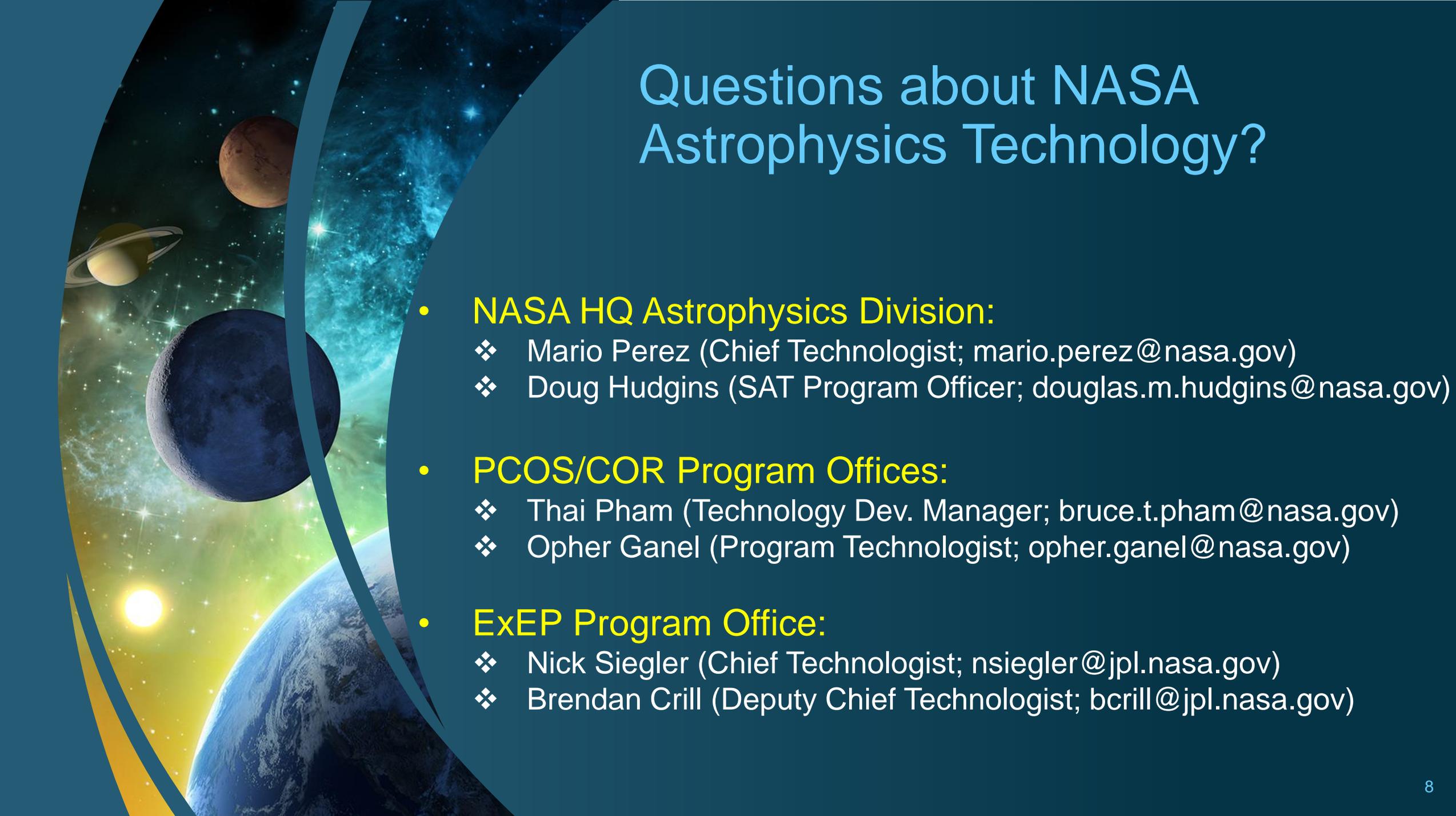
Gap Name	Description	Current State-of-the-Art	Performance Goals and Objectives
Angular Resolution (UV/Vis/NIR)	The capability to resolve the habitable zones of nearby star systems in the UV/Vis/NIR band, with a large space telescope.	Monolith: 3.5-m sintered SiC with < 3 μm SFE (Herschel); 2.4-m ULE with ~10 nm SFE (HST); Depth: Waterjet cutting is TRL 9 to 14", but TRL 3 to >18". Fused core is TRL 3; slumped fused core is TRL 3 (AMTD). Segmented: (no flight SOA): 6.5 m Be with 25 nm SFE (JWST); Non-NASA: 6 DOF, 1-m class SiC and ULE, < 20 nm SFE, and < 5 nm wavefront stability over 4 hr with thermal control	Large (4–16 m) monolith and multi-segmented mirrors for space that meet SFE < 10 nm rms (wavelength coverage 400–2500 nm); Wavefront stability better than 10 pm rms per wavefront control time step; CTE uniformity characterized at the ppb level for a large monolith; Segmented apertures leverage 6 DOF or higher control authority meter-class segments for wavefront control.
Coronagraph Contrast	The capability to suppress starlight with a coronagraph to the level needed to detect and spectrally characterize Earth-like exoplanets in the habitable zones of Sun-like stars.	unobscured pupil: 6×10^{-10} raw contrast at 10% bandwidth, angles of 3-15 λ/D (HLC demo in HCIT); obscured pupil: 1.6×10^{-9} raw contrast at 10% bandwidth across angles of 3-9 λ/D (WFIRST)	Maximized science yield for a direct imaging telescope/mission. $\leq 10^{-10}$ raw contrast, >10% throughput, IWA $\leq 3 \lambda/D$, obscured/segmented pupil

Planned 2021 Technology Timeline

Always consult NSPIRES for actual dates (<https://nspires.nasaprs.com/>)



* Contingent on SAT amendment in Sept 2021

A vibrant space-themed background featuring a large blue and yellow nebula, a bright sun, and several planets including Saturn, Mars, and the Moon. The scene is set against a dark starry sky.

Questions about NASA Astrophysics Technology?

- **NASA HQ Astrophysics Division:**

- ❖ Mario Perez (Chief Technologist; mario.perez@nasa.gov)
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- **ExEP Program Office:**

- ❖ Nick Siegler (Chief Technologist; nsiegler@jpl.nasa.gov)
- ❖ Brendan Crill (Deputy Chief Technologist; bcrill@jpl.nasa.gov)

Estimated Key Technology Dates in 2021

(If Astro2020 is released after mid-May some dates will change;
always consult NSPIRES for actual dates (<https://nspires.nasaprs.com/>))

- Technology gap* submission deadline ~ June 1
- Community-wide technology RFI released ~ June-July
- Updated technology gaps posted ~ September 16
- ABTR** 2021 with new gap priorities published ~ mid-October
- SAT 2021 solicitation (ROSES) ~ February 14
- SAT 2021 solicitation amendment ~ September 16
- Pending SAT 2021 amendment:
 - ❖ NOI's due ~ October 22
 - ❖ Proposals due ~ December 16

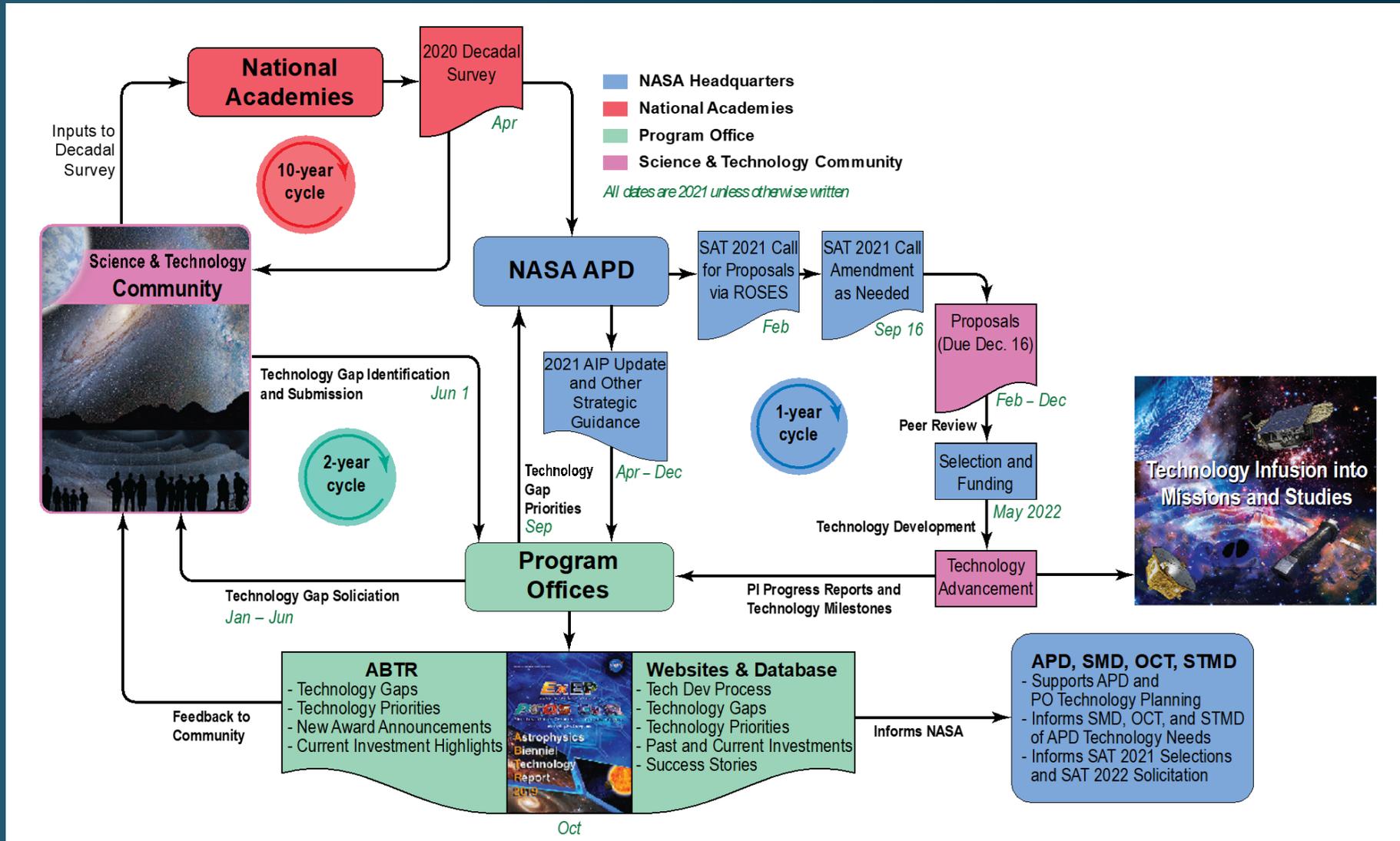
* Current gaps, priorities, and submission form with instructions available at https://apd440.gsfc.nasa.gov/tech_gap_priorities.html

** ABTR: Astrophysics Biennial Technology Report 2019 available at <https://apd440.gsfc.nasa.gov/technology.html>

The background of the slide is a composite of two cosmic images. The top half features a dark blue and black space filled with numerous small, bright stars and a prominent, glowing blue nebula on the right side. The bottom half shows a similar starry field but with a warm, golden-yellow and greenish glow, suggesting a different spectral filter or a different region of space. The text 'Additional Material' is centered in a white, sans-serif font across a dark blue horizontal band that separates the two images.

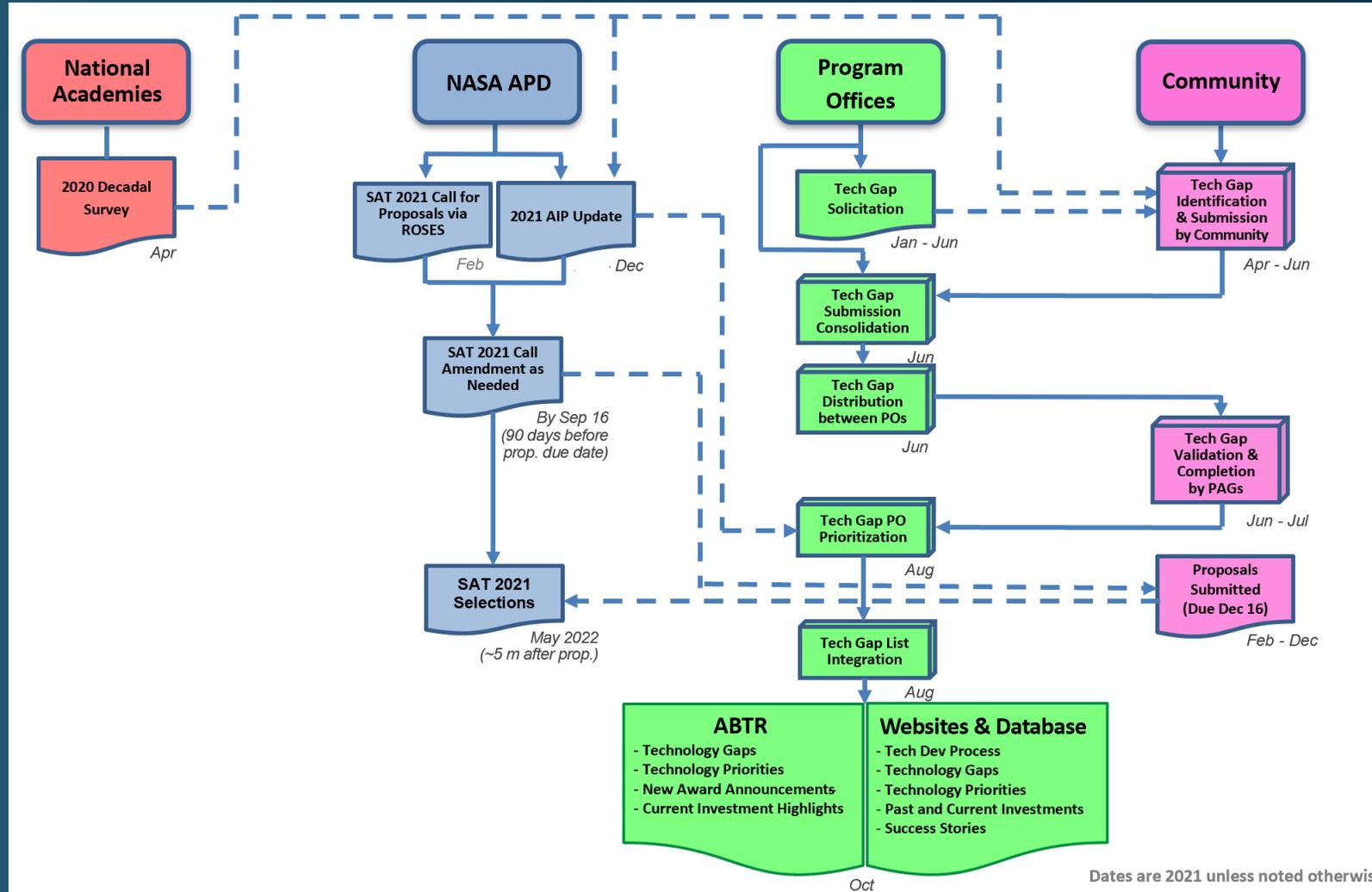
Additional Material

Technology Advancement Flow



Dates are best estimates

Technology Gap Prioritization and SAT Flow



Dates are best estimates

APD Technology Gaps by STDT

Technology Gap Name	STDT
Vis/NIR Detection Sensitivity	HabEx
Starshade Starlight Suppression and Model Validation	HabEx
Starshade Deployment and Shape Stability	HabEx
Large-Format Object Selection Technology for Multi-Object Spectrometers for HabEx	HabEx
Low-Noise, High-QE UV Detectors	HabEx
Photon-Counting, Large-Format UV Detectors	HabEx
Short-Wave UV Coatings	HabEx
Coronagraph Contrast Stability	LUVOIR
Angular Resolution (UV/Vis/NIR)	LUVOIR
Coronagraph Contrast	LUVOIR
Large-Format, Low-Noise, High-QE Far-UV Detectors	LUVOIR
Large-Format, High-Resolution, UV/Vis Focal Plane Arrays	LUVOIR
Next-Generation, Large-Format, Object Selection Technology for Multi-Object Spectrometers for LUVOIR	LUVOIR
High-Reflectivity Broadband FUV-to-NIR Mirror Coatings	LUVOIR
UV Detection Sensitivity	LUVOIR
UV Coatings	LUVOIR
Fast, Low-Noise, Megapixel X-Ray Imaging Arrays with Moderate Spectral Resolution	Lynx
Large-Format, High-Spectral-Resolution, Small-Pixel X-Ray Focal-Plane Arrays	Lynx
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High-Resolution, Large-Area, Lightweight X-Ray Optics	Lynx
Low-stress, Highly Stable X-Ray Reflective Coatings	Lynx
Long-Wavelength-Blocking Filters for X-Ray Micro-Calorimeters	Lynx
Large-Format, Low-Noise and Ultralow-Noise Far-IR Direct Detectors	Origins
Cryogenic Readouts for Large-Format Far-IR Detectors	Origins
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