

# Exoplanet Exploration Program Technology Update

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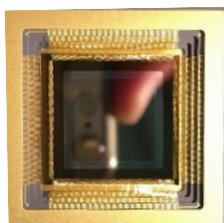
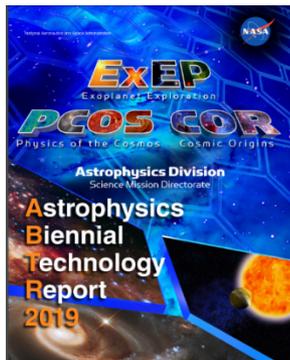
**Nicholas Siegler**  
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Exoplanet Exploration Program  
Jet Propulsion Laboratory / California Institute of Technology

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Deputy Technology Development Manager  
Exoplanet Exploration Program  
Jet Propulsion Laboratory / California Institute of Technology

**ExoPAG 22**  
**19 Jun 2020**

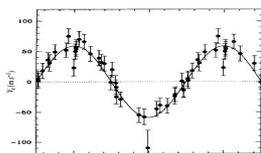
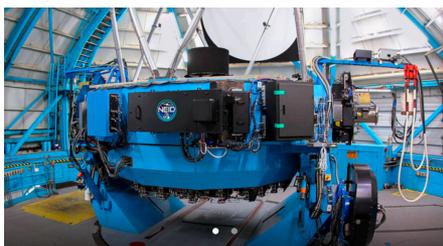
# Technology Activities

## Technology Gaps

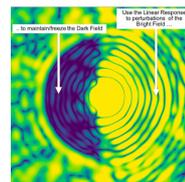


## Deformable Mirror Survey

## Extreme Precision Radial Velocity

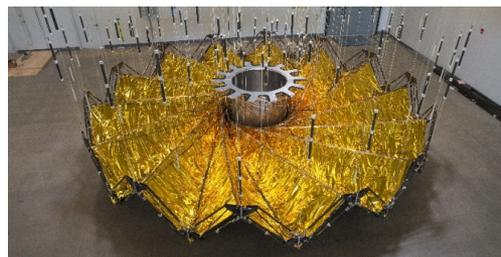


## Strategic Astrophysics Technology Grants

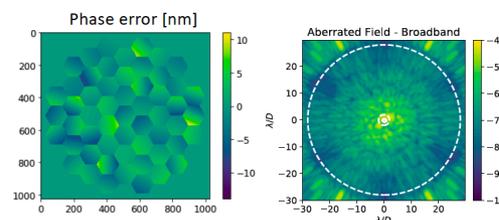


- *Coronagraph architectures: modeling and demonstrations*
- *Wavefront control*
- *Extreme Precision Radial Velocity*
- *Detectors*

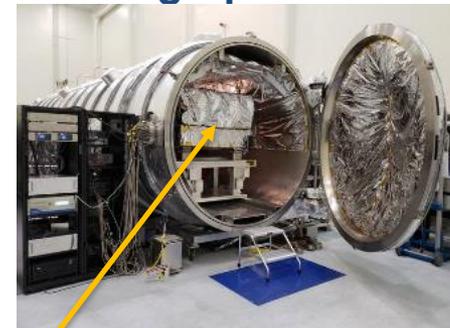
## Starshade Technology Development



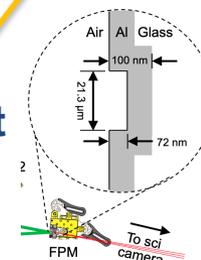
## Segmented Coronagraph Design & Analysis Study



## Ultra-Stable Coronagraph Testbeds



## Zernike Wavefront Sensor



## Roman/CGI



# Strategic Astrophysics Technology (1/2)

- **4 coronagraph masks/architectures**

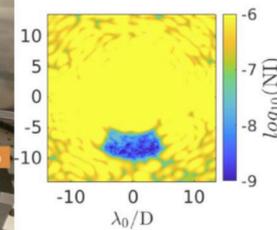
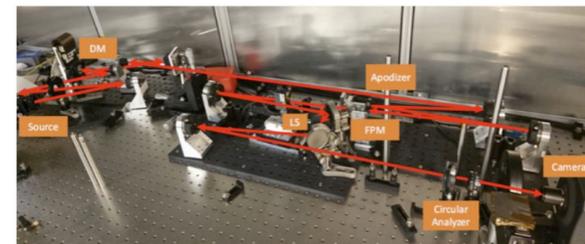
- Vortex Coronagraph (Serabyn/NASA-JPL)
- Phase Induced Amplitude Apodization Complex Mask Coronagraph (Belikov/NASA-ARC)
- Apodized Pupil Lyot Coronagraph (Soummer/STScI)
- Super-Lyot Coronagraph (Trauger/NASA-JPL)

*PIAACMC  
in vacuum  
chamber now*



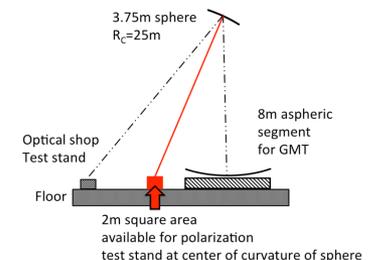
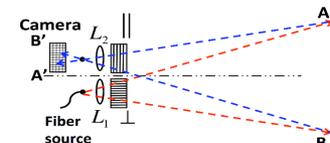
- **4 wavefront-control techniques**

- Single mode fiber and optimization for spectroscopy (Mawet/Caltech)
- Multi-star wavefront control demos (Belikov/NASA-ARC)
- WFC using light outside the dark field (Guyon/UA)
- MEMS deformable mirrors (Bierden/BMC)



*Mawet SAT-2018*

- **Polarization & Coronagraphy (Breckinridge/UA)**

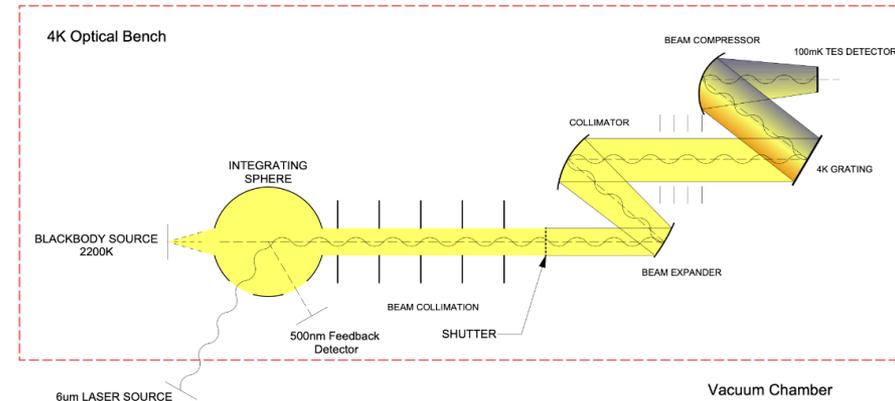


*Breckinridge Final Report under review*

# Strategic Astrophysics Technology (2/2)

## Detectors

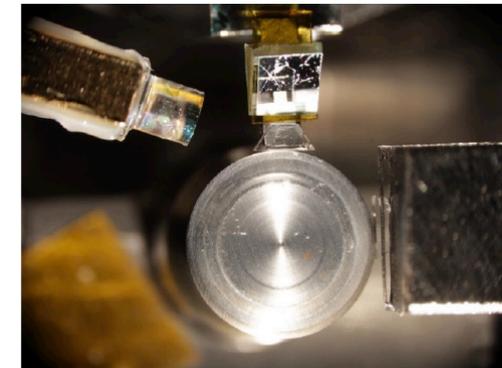
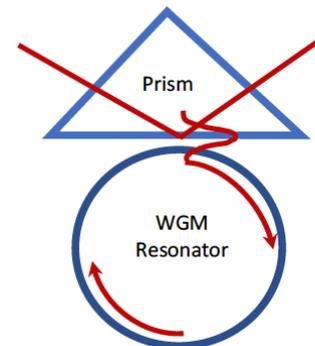
- Vis-band rad-hard photon-counting detectors (Rauscher/NASA-GSFC)
- Ultra-stable mid-IR detector array (Staguhn/JHU)



*Staguhn SAT2018*

## Extreme Precision Radial Velocity

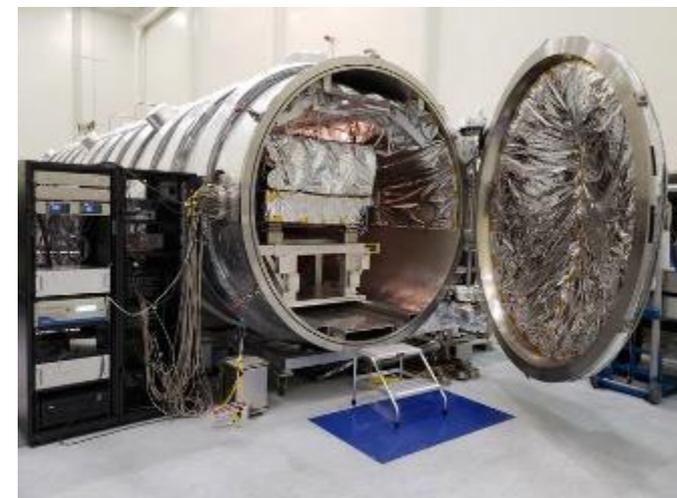
- Micro-resonator optical etalon for radial velocity measurements (Leifer/NASA-JPL)



*Leifer SAT2018*

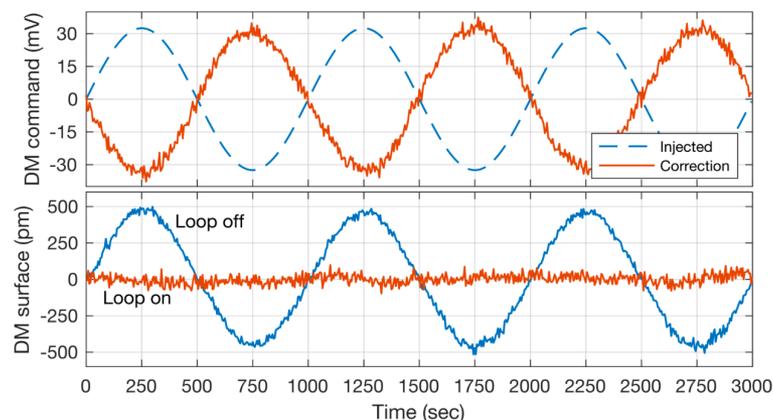
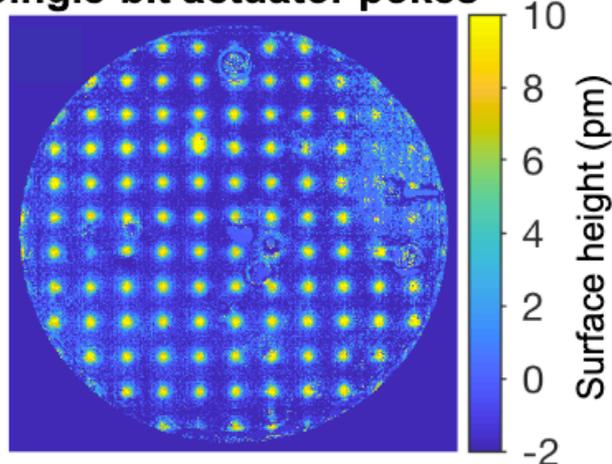
# Decadal Survey Testbed (DST)

- Coronagraph testbed specifically designed for opto-mechanical stability permitting demonstration of  $10^{-10}$  contrast
- Zernike wavefront sensor commissioned in DST, achieving picometer sensitivity, and available to investigators



ExEP's HCIT-2 Vacuum Chamber and the DST at JPL

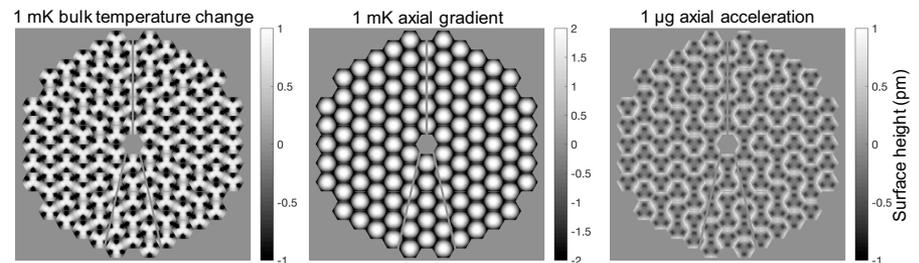
## Single-bit actuator pokes



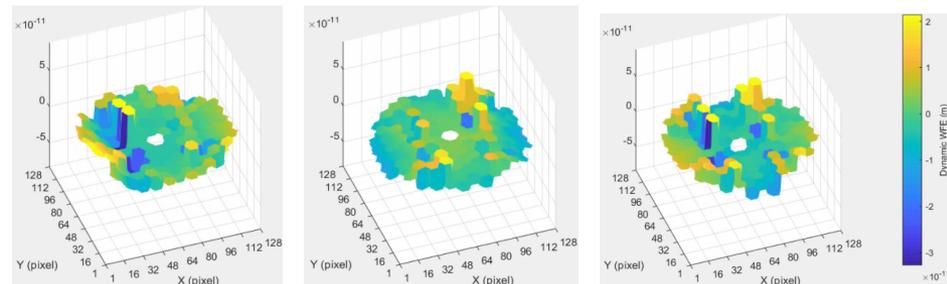
- Segmented pupil (static) will commence later in 2020, targeting large space telescope concepts; also installing EMCCD

# Segmented Coronagraph Design & Analysis

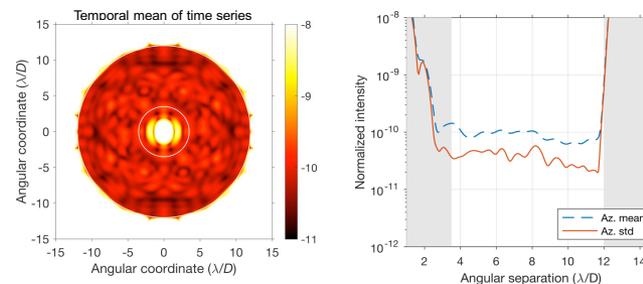
- **2020 Objectives:**
  - Investigate sensitivity of science yield to telescope aberrations/instabilities
  - investigate how coronagraph requirements drive telescope requirements
- **Year-to-Date Accomplishments**
  - Produced end-to-end propagation model
  - Incorporated quasi-static and dynamic telescope aberrations



*Quasi-Static Segment Deformations (Coyle, East)*



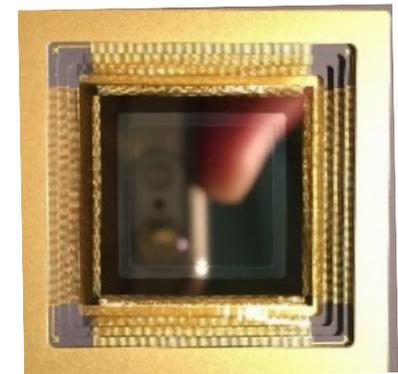
*Snapshots of Dynamic Wavefront Errors (Chopra, Dewell, Nordt)*



*Preliminary Contrast Result, APLC w/ Angular Differential Imaging (Ruane, JPL)*

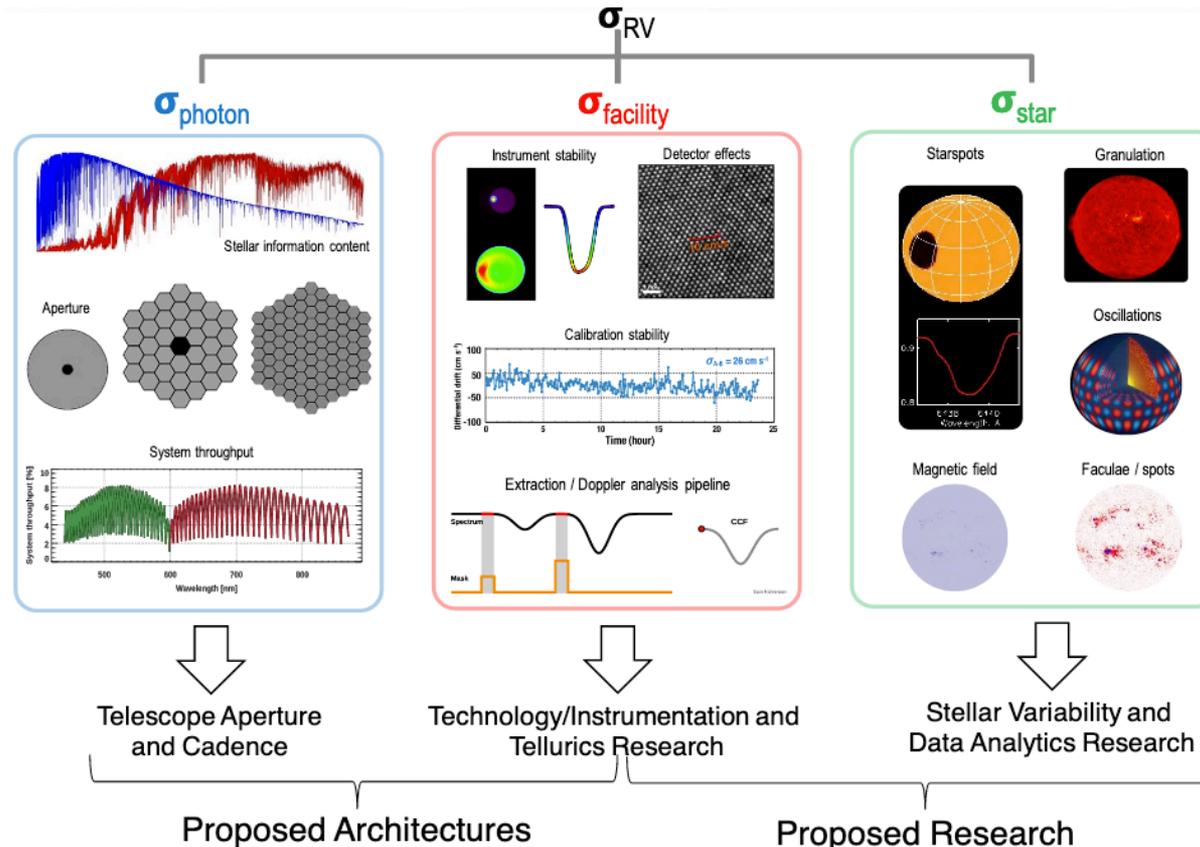
# Deformable Mirror Survey

- *Have we identified all candidate deformable mirrors that might be suitable for a future exoplanet direct imaging mission?*
- **A 1-year survey to assemble a Subject-Matter Expert team, define requirements, perform a global fact-finding effort, and write a report.**
- **Identify novel wavefront control technologies that could mature rapidly.**
- **Current Status: Defined requirements with Subject Matter Experts, fact finding ongoing with vendors.**
- **Report to be made public at the end of the year**



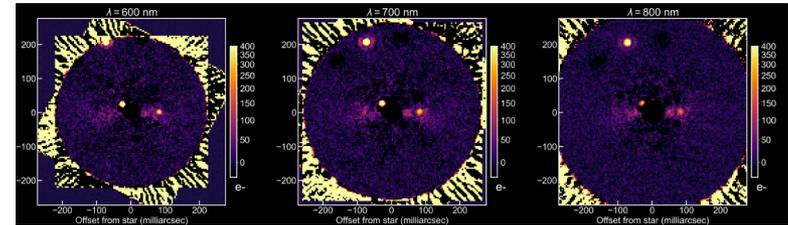
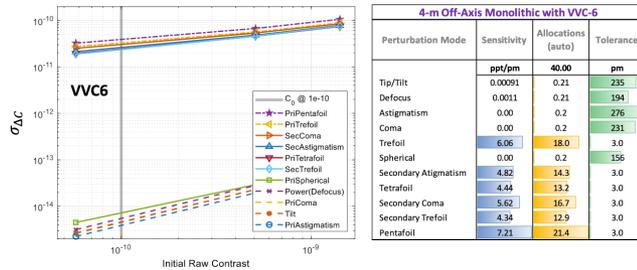
# Extreme Precision Radial Velocity

- Extreme Precision Radial Velocity Initiative Plan delivered to NASA and NSF in March 2020
- See upcoming talk by Scott Gaudi and Jenn Burt for the details!

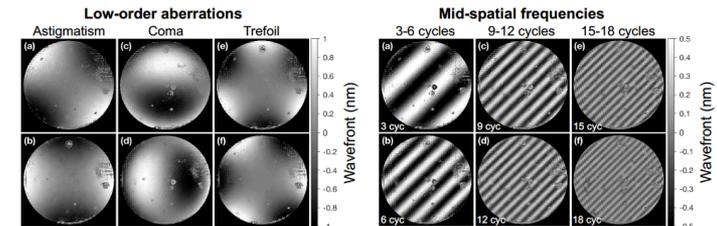
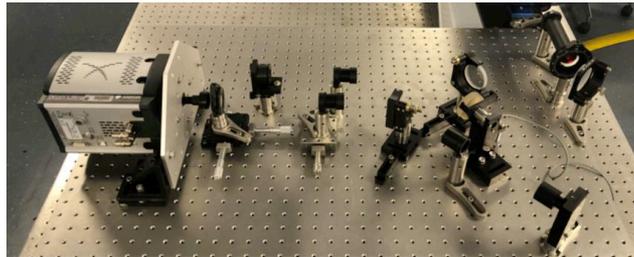


# Exoplanet Exploration Technology Colloquium Series

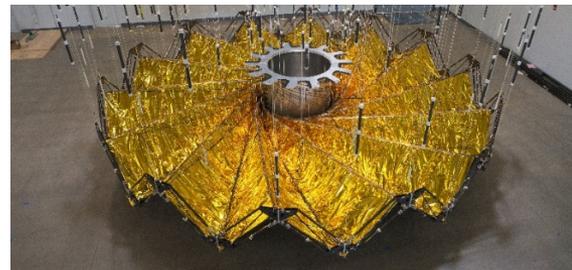
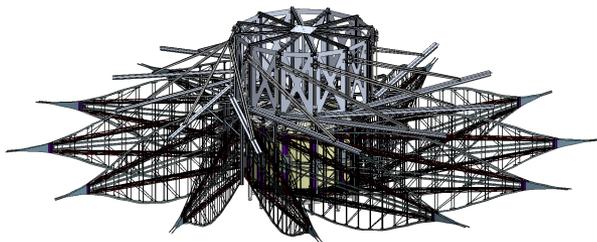
- Telescope Stability + Coronagraph modeling



- Workshop on Advanced Wavefront Sensing for Coronagraphy



- Update on Starshade Technology Development



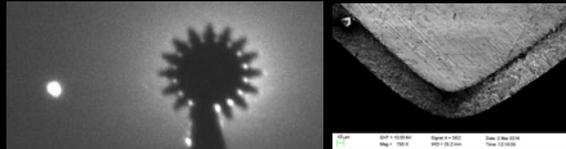
- Recordings and slides available here:

- [https://exoplanets.nasa.gov/exep/technology/tech\\_colloquium/](https://exoplanets.nasa.gov/exep/technology/tech_colloquium/)

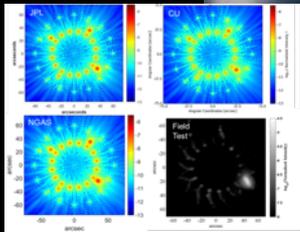
# The Three Starshade Technology Gaps

<https://exoplanets.nasa.gov/exep/technology/starshade/>

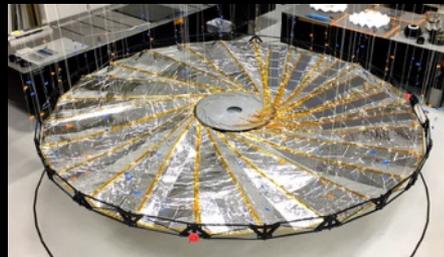
## (1) Starlight Suppression



Suppressing scattered light off petal edges from off-axis Sunlight (S-1)

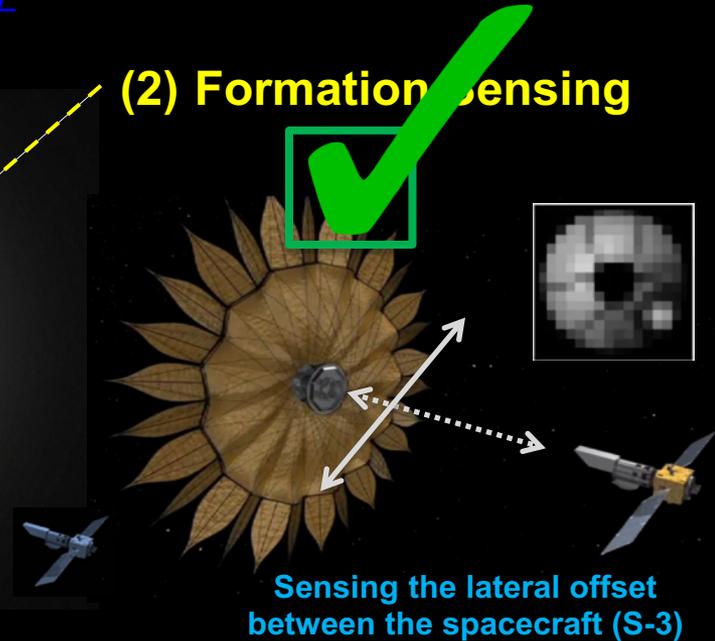


Suppressing diffracted light from on-axis starlight and optical modeling (S-2)



Positioning the petals to high accuracy, blocking on-axis starlight, maintaining overall shape on a highly stable structure (S-5)

## (2) Formation Sensing



Sensing the lateral offset between the spacecraft (S-3)

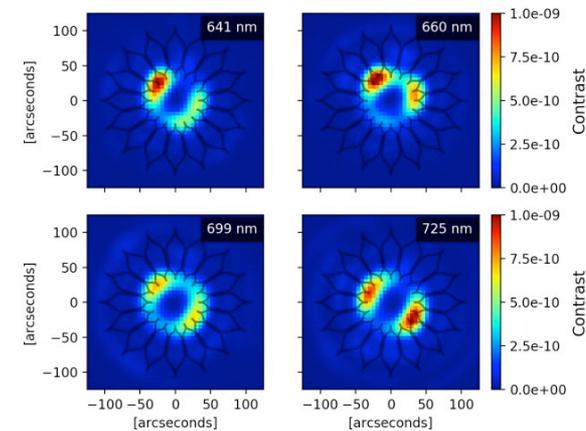
## (3) Deployment Accuracy and Shape Stability



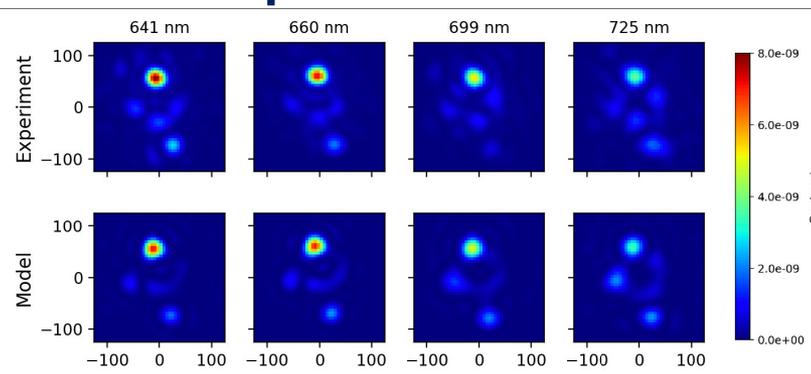
Fabricating the petals to high accuracy (S-4)

# Starlight Suppression and Edge Scatter

- Validate models of starshade performance
- Princeton testbed has been used to demonstrate starlight suppression to  $< 10^{-10}$  over a 10% band of sub-scale starshade

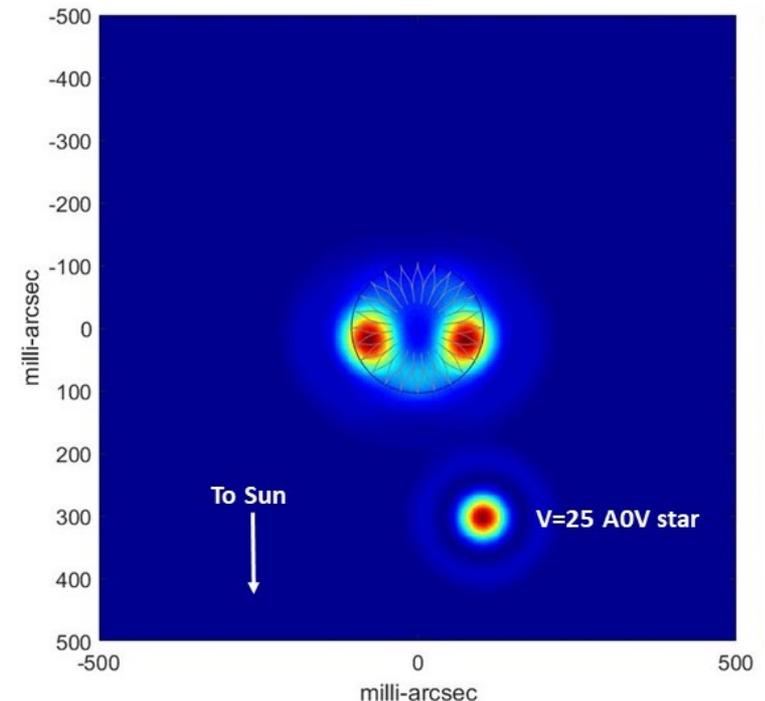
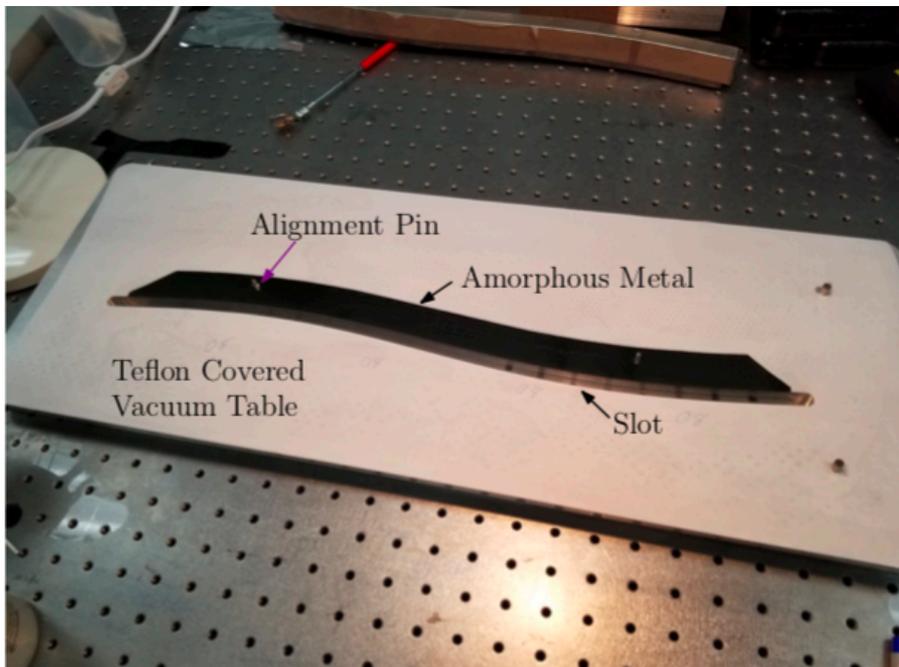


- Last step: measure deliberately misshapen subscale starshades and compare with model predictions



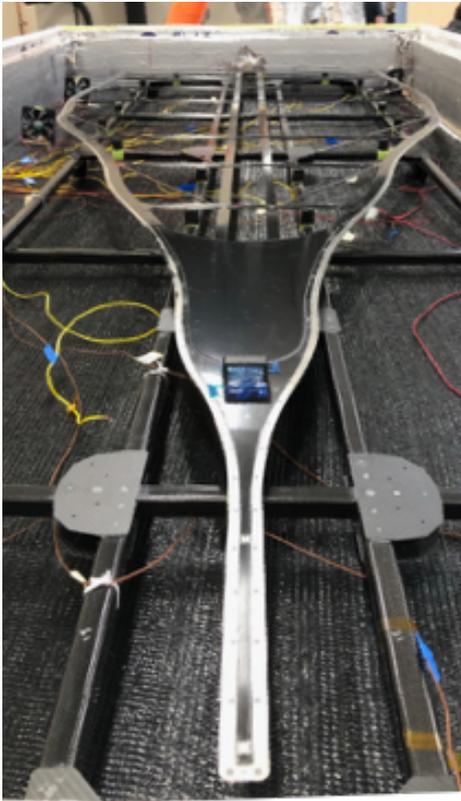
# Starlight Suppression and Edge Scatter

- Demonstrated starshade optical edge limits Solar scatter performance to lobe dimmer than mag 25 and maintains performance after thermal cycling

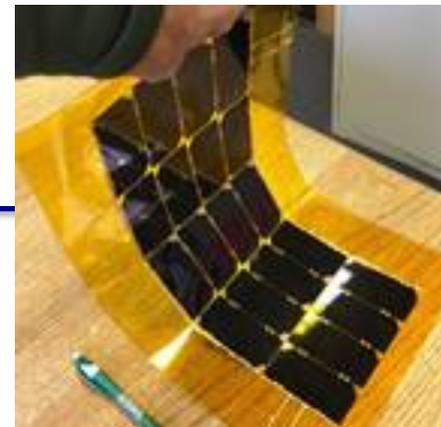


# Shape Accuracy: Critical Features demonstrated

- **Starshade Petal System successfully demonstrated to maintain prelaunch shape within  $\pm 70 \mu\text{m}$  after deploy and thermal cycles**



# Full Scale Inner Disk



Integrated solar cells

Gravity offloading



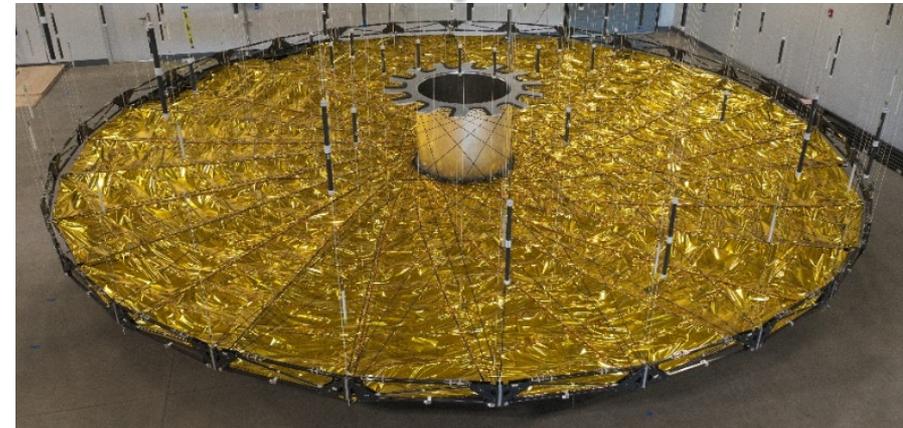
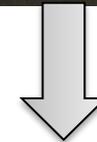
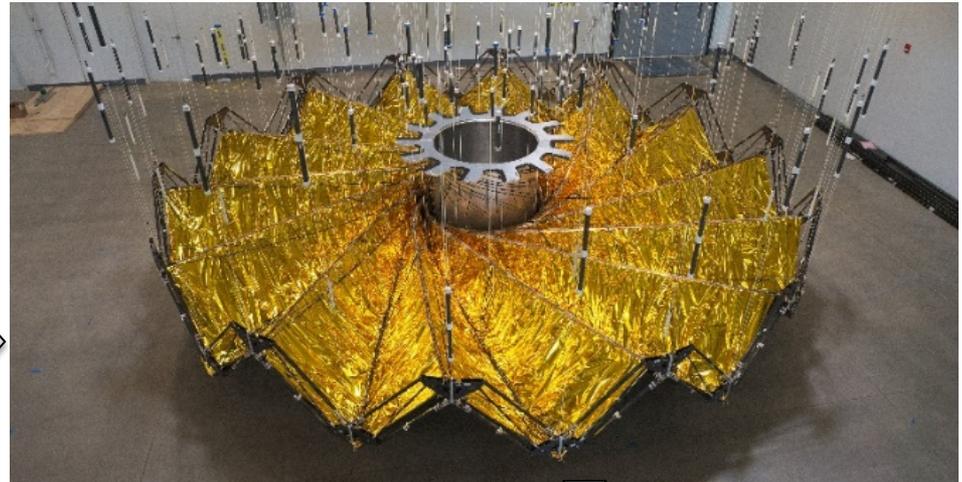
(a) Deployed



(b) Stowed

# Shape Accuracy: Critical Features demonstrated

- Full Scale Prototype Tested

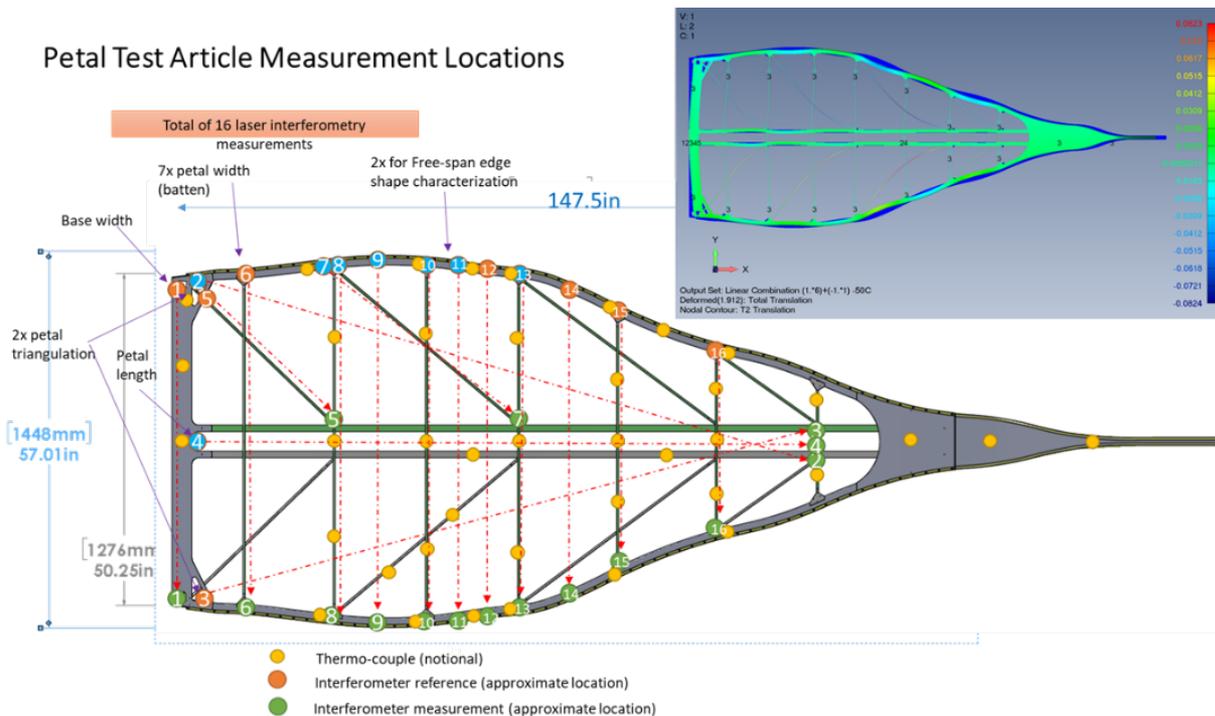


- Inner disk deployment dimensional stability with thermal cycles and storage successfully demonstrated to  $\pm 300 \mu\text{m}$

# Starshade Shape Stability: critical components demonstrated

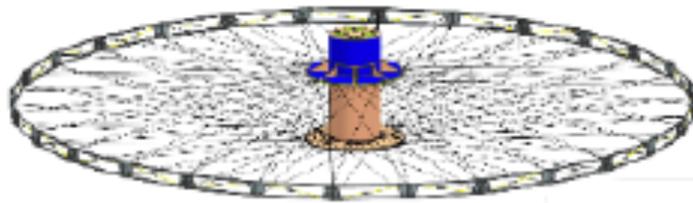
- On-orbit thermal stability on-orbit of petal critical dimensions demonstrated to  $\pm 80 \mu\text{m}$  through measurement and models

Petal Test Article Measurement Locations

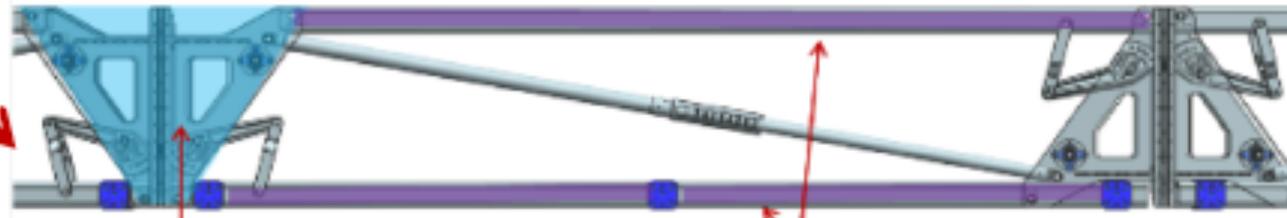
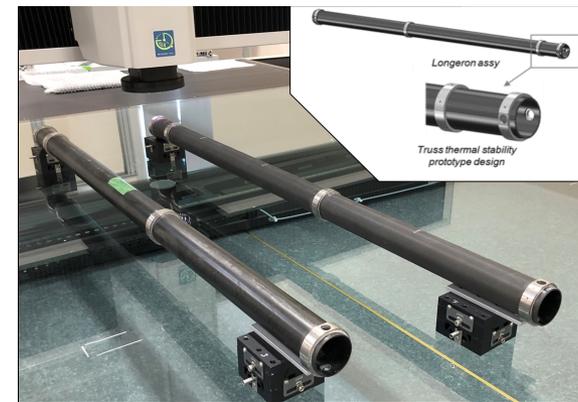


# Starshade Shape Stability: critical components demonstrated

- Thermal stability on-orbit of critical parts of inner disk demonstrated to  $\pm 200 \mu\text{m}$



Perimeter Truss Bay Assembly



Node sub-assy (blue)

Longeron sub-assy (purple)

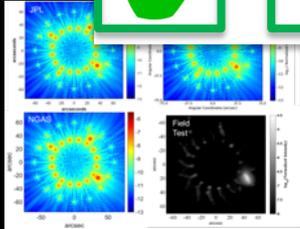
# Starshade Technology Gaps Scorecard

<https://exoplanets.nasa.gov/exep/technology/starshade/>

## (1) Starlight Suppression



Suppressing scattered light off petal edges from off-axis Sunlight (S-1)



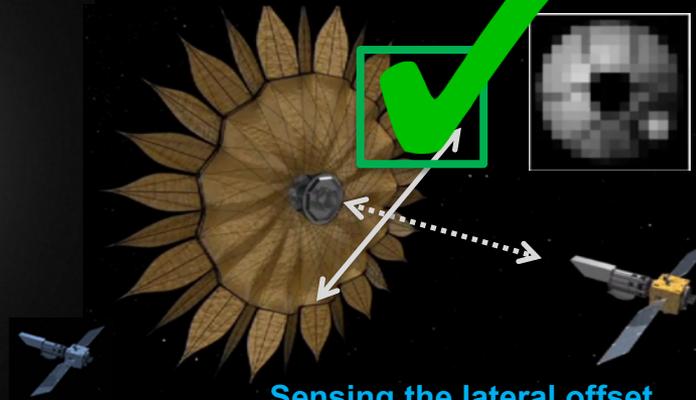
One milestone remaining!



Suppressing diffracted light from on-axis starlight and optical modeling (S-2)

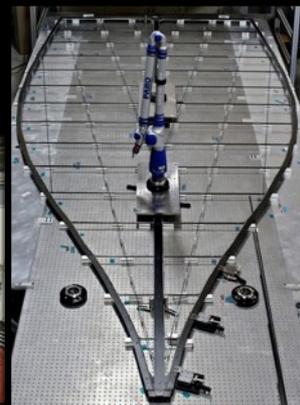


## (2) Formation Sensing



Sensing the lateral offset between the spacecraft (S-3)

## (3) Deployment Accuracy and Shape Stability



Fabricating the petals to high accuracy (S-4)

Positioning the petals to high accuracy, blocking on-axis starlight, maintaining overall shape on a highly stable structure (S-5)

Remaining accuracy and stability milestones complete by 2023

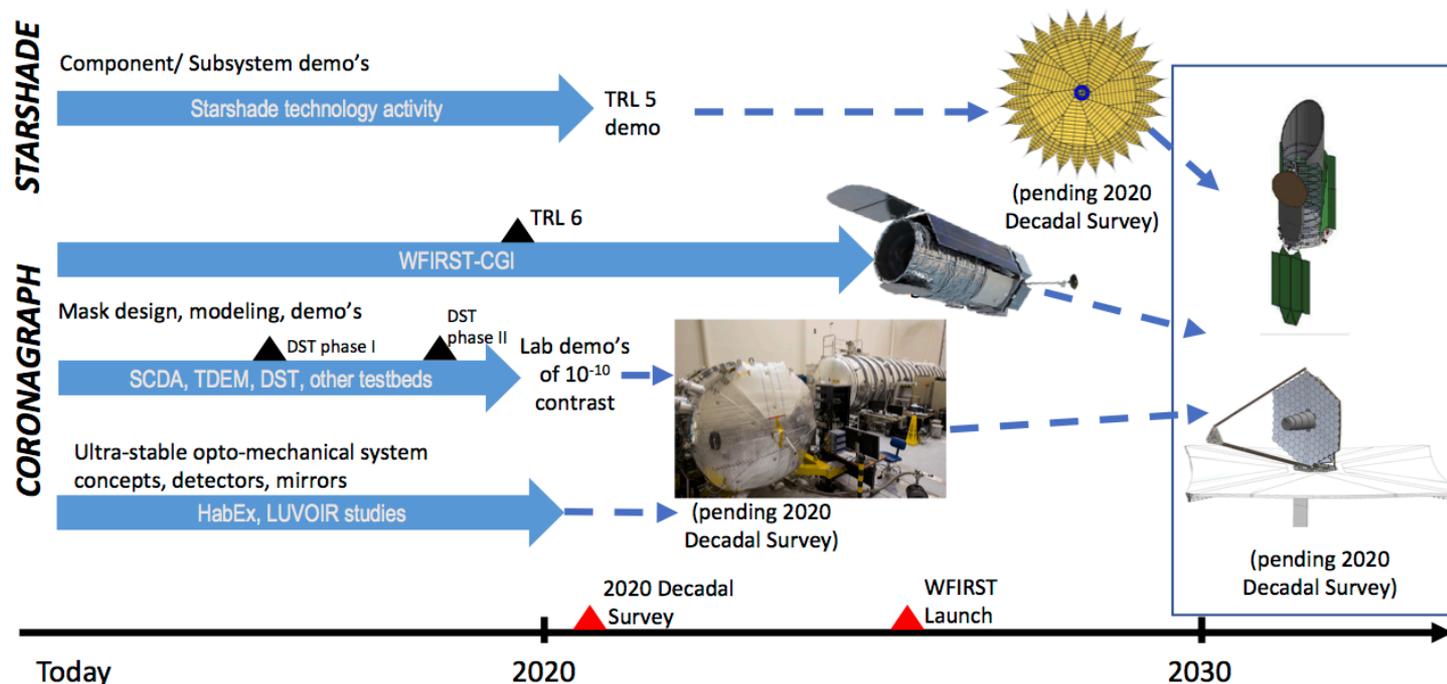
# Looking ahead..

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- **SAT Results coming:**
  - Final Reports on Polarization, MEMS DMs
  - Lab Coronagraph demonstrations of PIAACMC, Vortex, HLC
  - Further upgrades to DST – segmented pupil, EMCCD
- **Starshade work continues: subscale optical demos, deployment shape and stability demos**
- **Segmented Coronagraph Design and Analysis: telescope tolerancing results**
- **Roman/CGI: Zernike Wavefront Sensor and Multi-Star Wavefront Control masks in spare slots**
- **Extreme Precision Radial Velocity solicitation coming in August**
- **Astro2020 coming 2021**

# Want to get involved?

- Please contact me
- in July or August 2020, I plan to offer a 1-hour ExEP Technology Primer briefing: a deeper dive into our technology activities
  - look for an email on ExoPAGannounce



**BACKUP**

# Starshade Technology Activity (S5) Technology Milestones Scorecard



Complete 2020

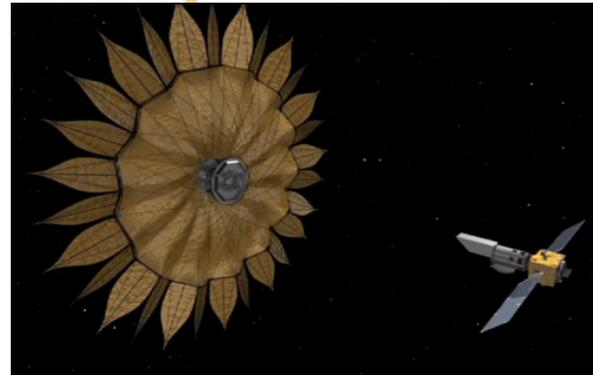


Complete 2023

## Starlight Suppression



## Scattered Sunlight



## Formation Flying



Critical Features

All Features

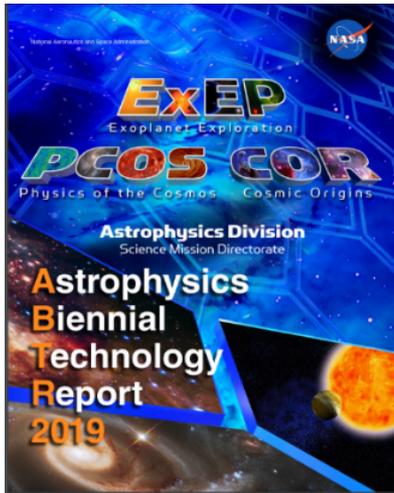
## Shape Accuracy



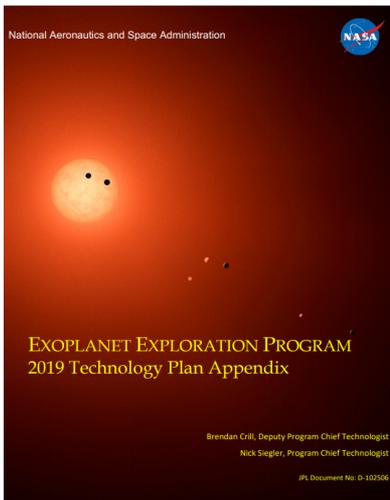
## Shape Stability



# Technology Gap List



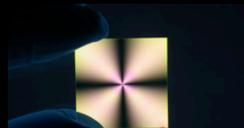
- **Astrophysics Technology Gap List**
  - Technology gaps for all three NASA Astrophysics Division (APD)'s programs
  - Database of technology activities:
    - <http://astrostrategictech.us/>
  - Update coming in 2021, post-decadal



- **Exoplanet Technology Gap List**
  - Subset of APD gap list corresponding to exoplanet science:
    - <https://exoplanets.nasa.gov/exep/technology/gap-lists/>

# V-NIR Coronagraph/Telescope Technology Gaps

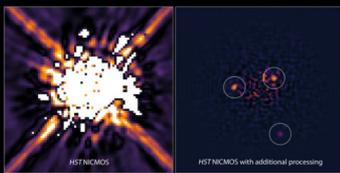
## Contrast



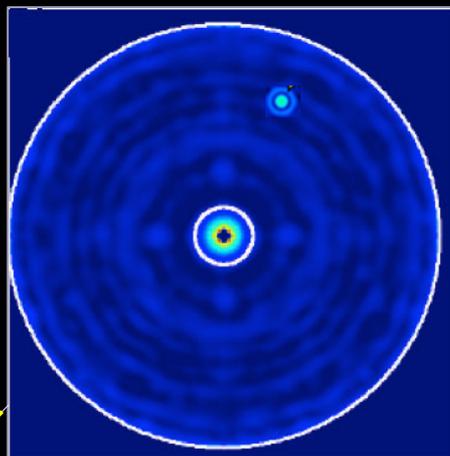
CG-2: Coronagraph Architecture



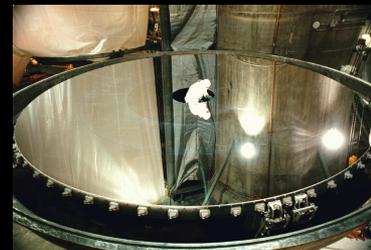
CG-3: Deformable Mirrors



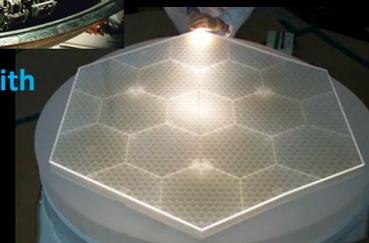
CG-4: Data Post-Processing



## Angular Resolution

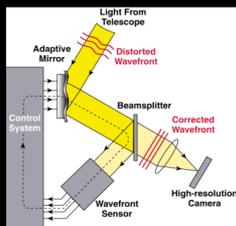


CG-1: Large Monolith Mirrors



CG-1: Segmented Mirrors

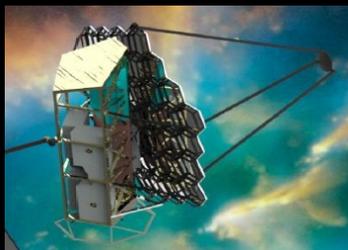
## Contrast Stability



CG-5: Wavefront Sensing and Control

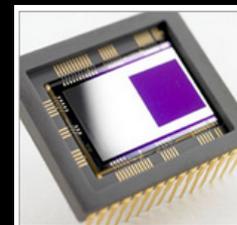


CG-6: Mirror Segment Phasing



CG-7: Telescope Vibration Sensing and Control or Reduction

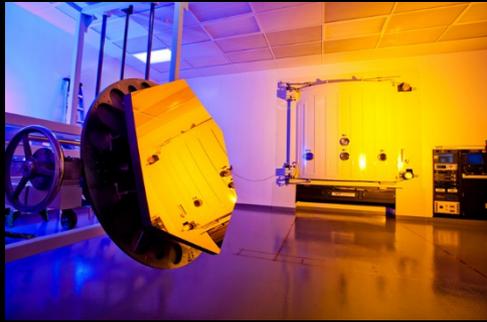
## Detection Sensitivity



Ultra-low Noise Visible (CG-8) and Infrared (CG-9) Detectors

# Other Technology Gaps

## UV Contrast

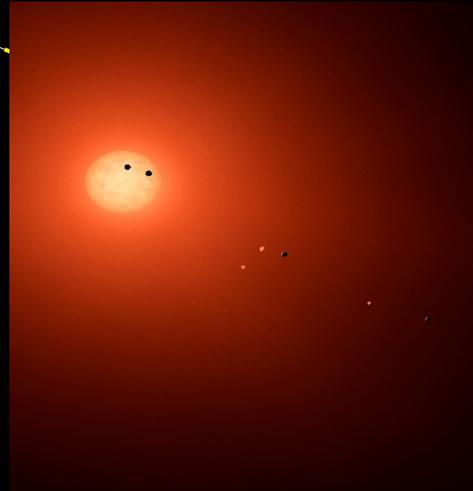
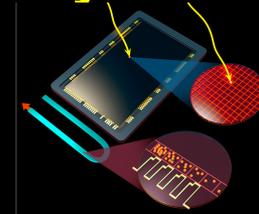


CG-10 UV/V/NIR Mirror Coatings

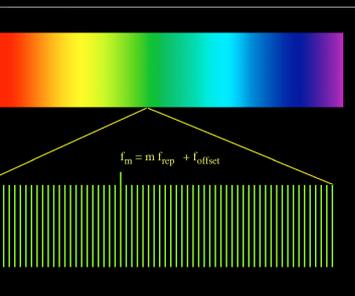
## UV Detection Sensitivity



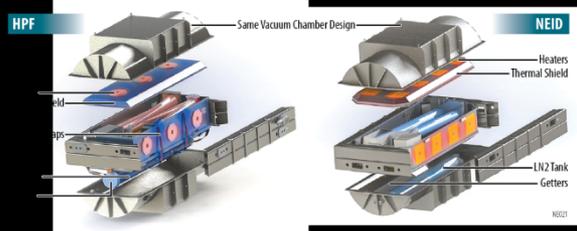
CG-12: Ultra-low Noise UV Detectors



## Stellar Reflex Motion Sensitivity

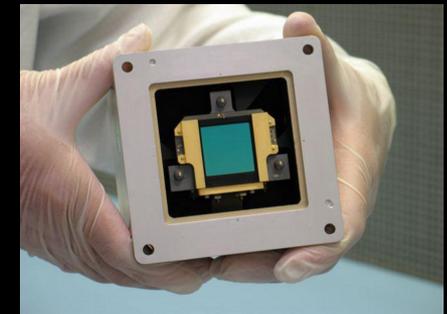


M-2: Laser Frequency Combs for Space-based EPRV

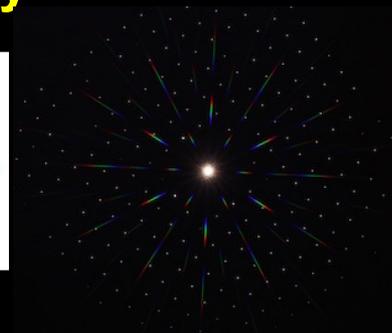


M-1: Ground-based Ultra-high Precision Radial Velocity

## Transit Spectroscopy Sensitivity



M-4: Ultra-stable Mid-IR Detectors for Transit Spectroscopy



M-3: Astrometry