

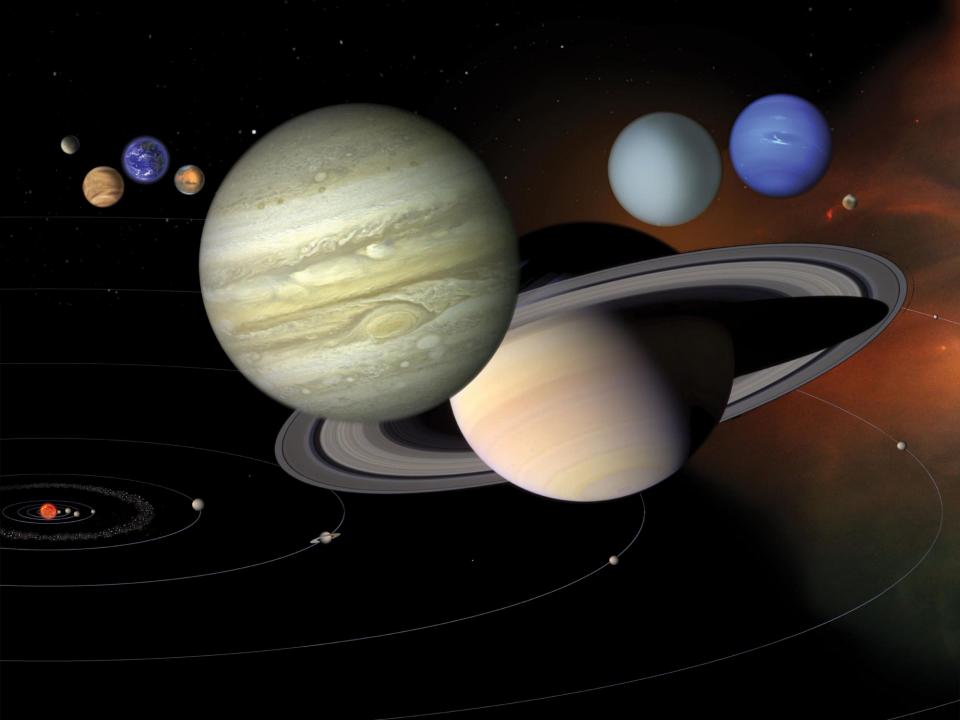
# Cool Planets, New Science Exoplanet Science Enabled by Starshades

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Twice as big in volume as the Earth, HD 40307g straddles the line between "Super-Earth" and "mini-Neptune" and scientists aren't sure if it has a rocky surface or one than'ts buried beneath thick layers of gas and ice. One thing is certain, though: at eight times the Earth's mass, its grantiational pull is much, much stronger.

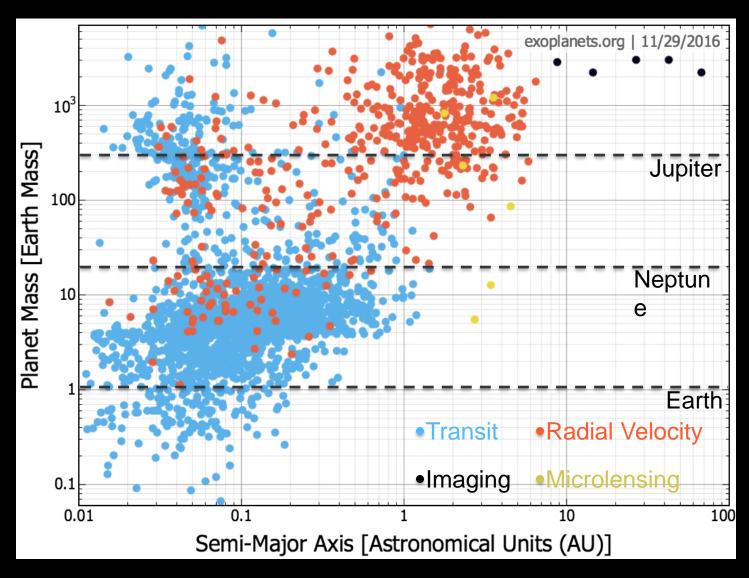
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Kepter-1869 is the first Earth-size planet discovered in the potentially habitable zond around another star, where liquid water could exist on the planet's surface.
Its star is much cooler and redder than our Sun. If plant life does exist on a planet like Kepter-1861, its photosynthesis could have been influenced by the stair's red-wavelength photone, making for a color palede that's very discered than the greens on Earth.
This discovery was made by Kepter, NASAs planet-hunding space lateocops.

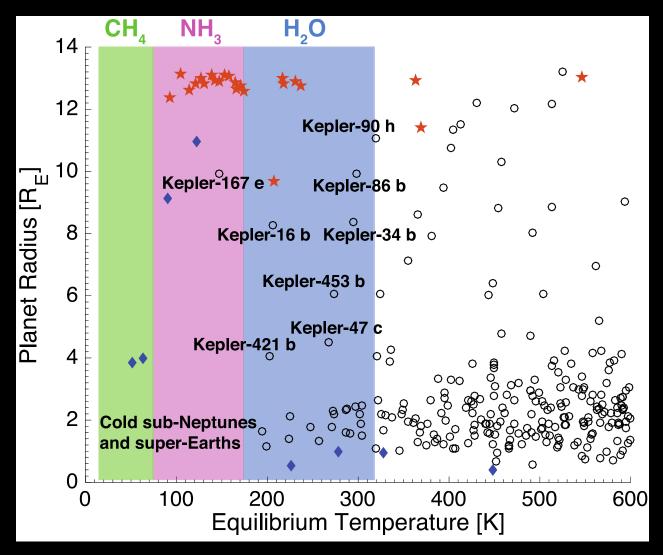
## **Exoplanet Demography**

#### Commonality of Small Exoplanets



### **Exoplanet Demography**

Wide-Separation Exoplanets



Giant exoplanets found by RV surveys

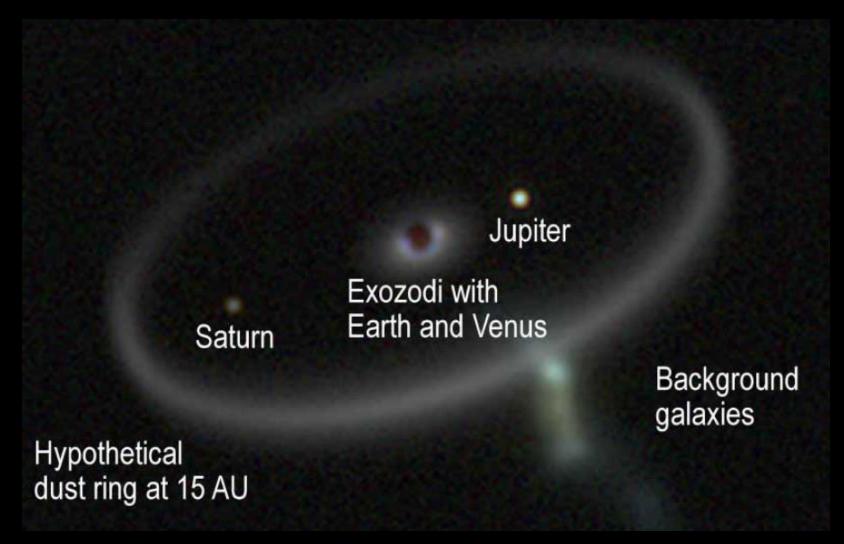
O Transiting exoplanets found by *Kepler* 

Planets in the Solar System

Hu, in prep.

#### Image of a Solar System

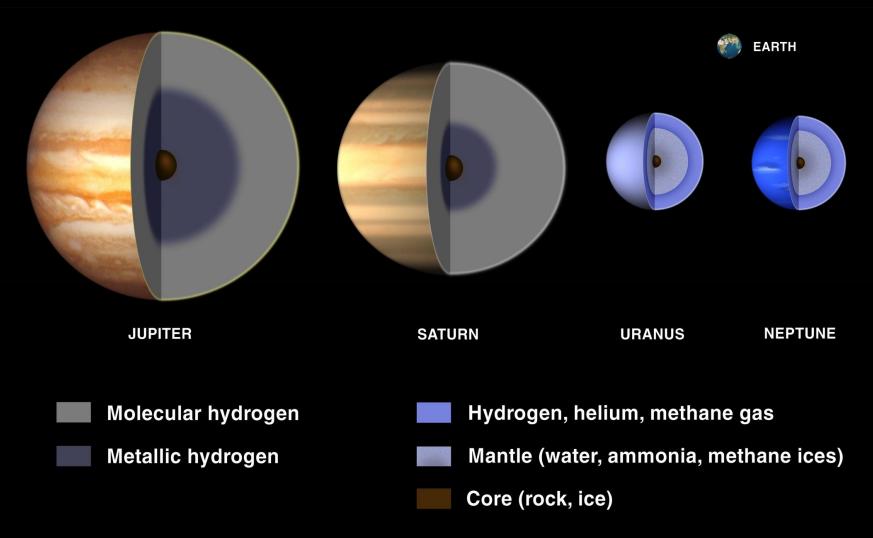
Starshade Rendezvous Mission simulated image of a nearby star



Seager et al. 2015

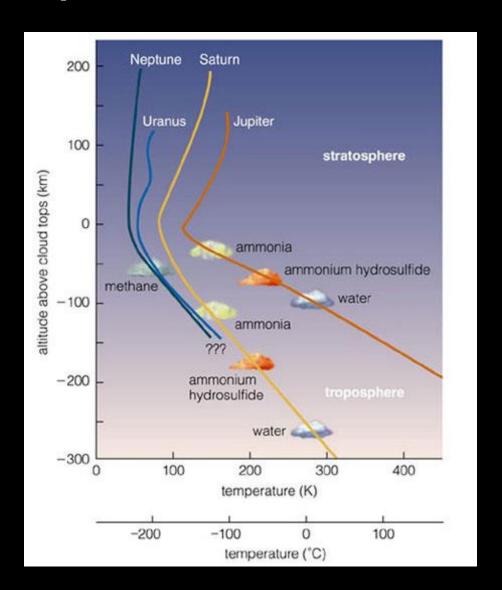
#### **Characterizing Giant Exoplanets**

Solar-System Analogs



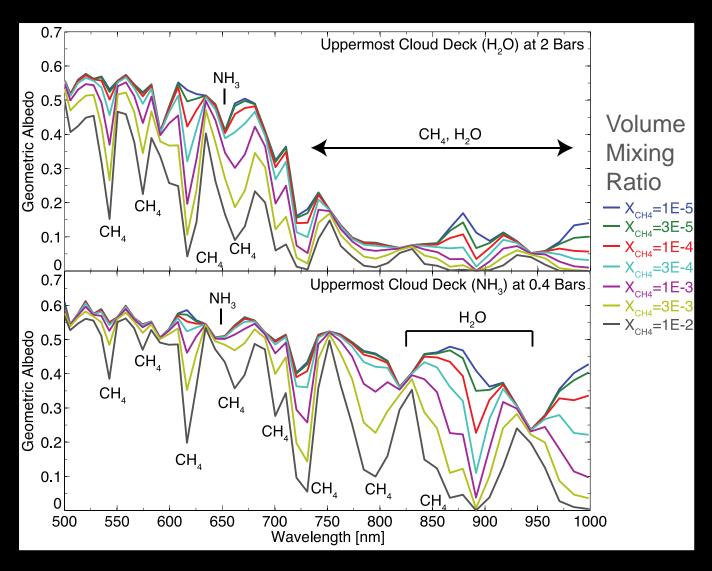
#### **Characterizing Giant Exoplanets**

- Abundances of C, N, O
- Convection and advection
- Thermal evolution
- Role of photochemistry
- Comparative planetology



#### **Characterizing Giant Exoplanets**

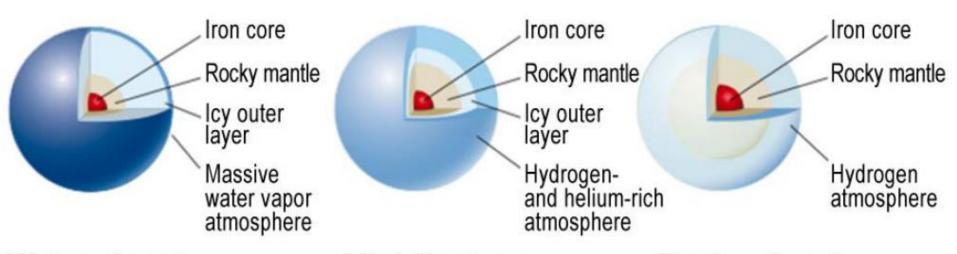
#### Cloud and Gas Abundance



Hu 2014

## Measuring Spectra of Sub-Neptunes and Super Earths

- Is there a surface?
  - Cloud, liquid, or solid
- What is the composition of the atmosphere?
  - H<sub>2</sub>-dominated or non-H<sub>2</sub>-dominated
- What are the formation and evolution pathways?



Water planet

Mini-Neptune

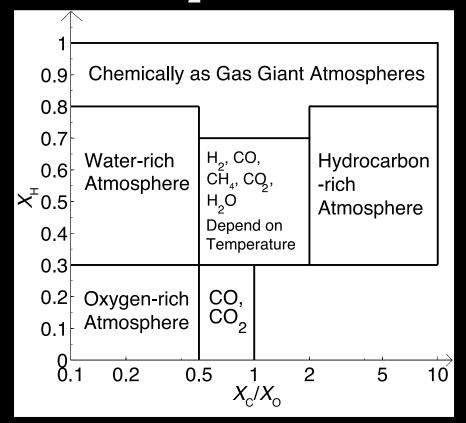
Rocky planet

## Measuring Spectra of Sub-Neptunes and Super Earths

**Diversity of Atmospheres** 

 $H_2$  – Rich





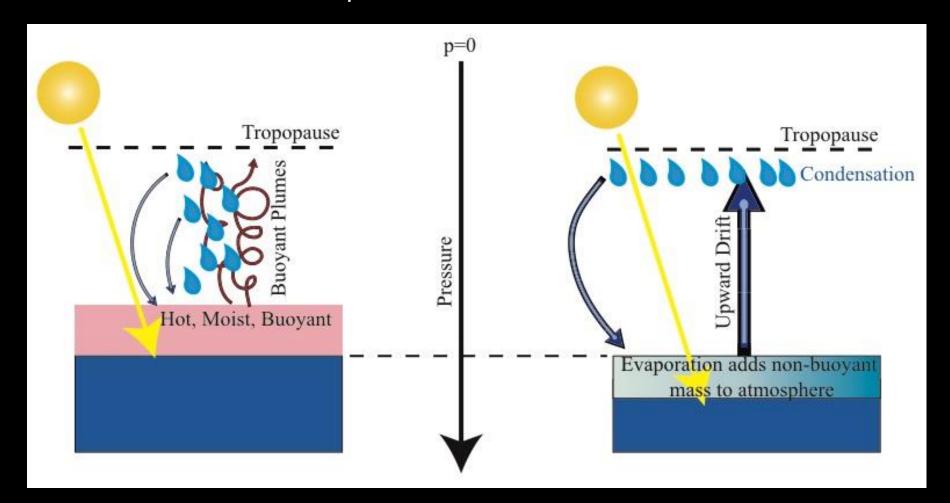
Carbon

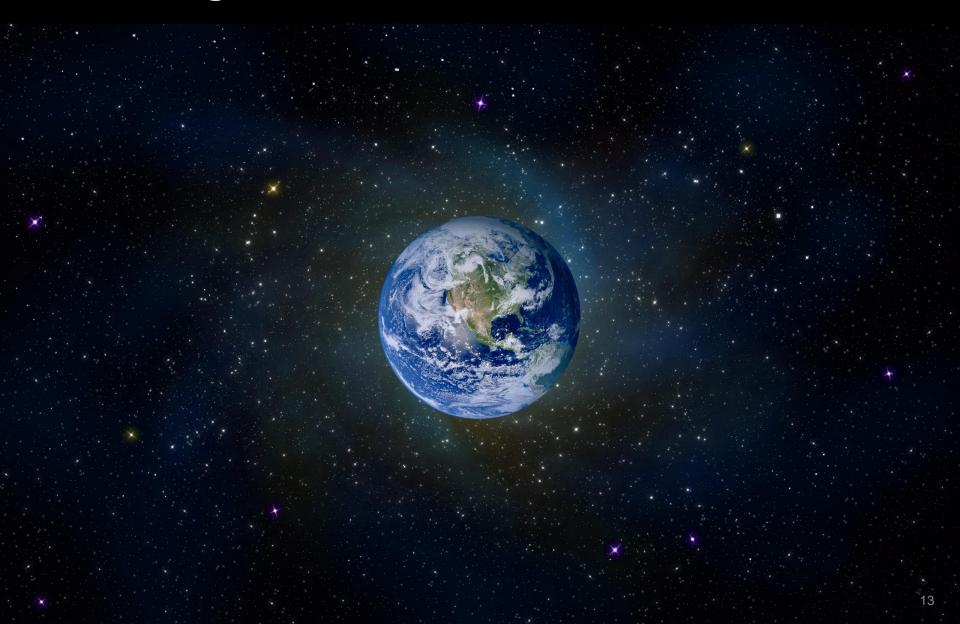
– Rich

 $H_2$  – Poor

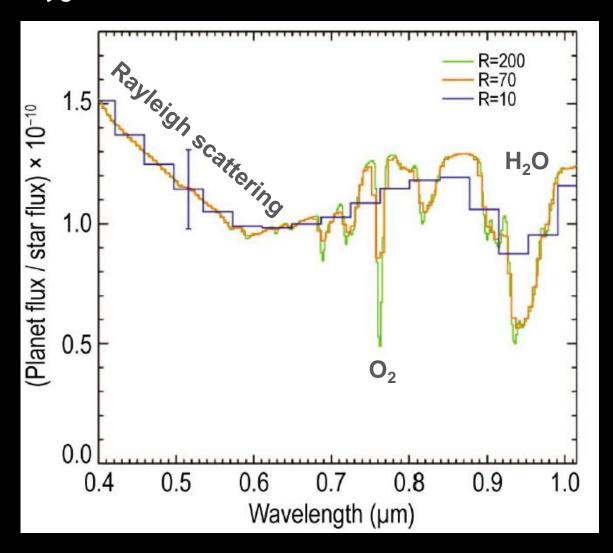
## Measuring Spectra of Sub-Neptunes and Super Earths

Condensible-Rich Atmospheres

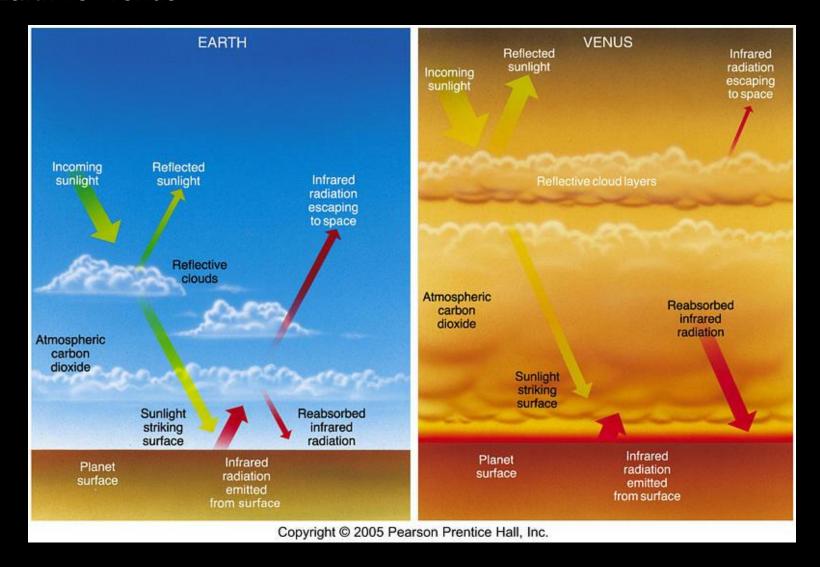




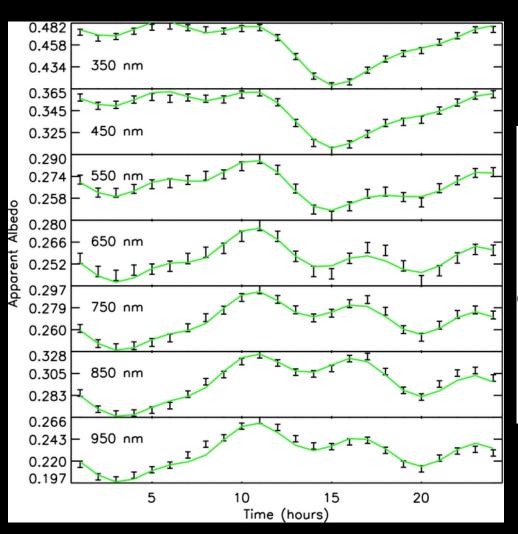
Water and Oxygen

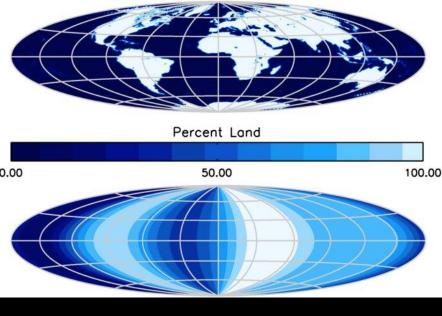


#### Earth vs. Venus



#### Land and Sea



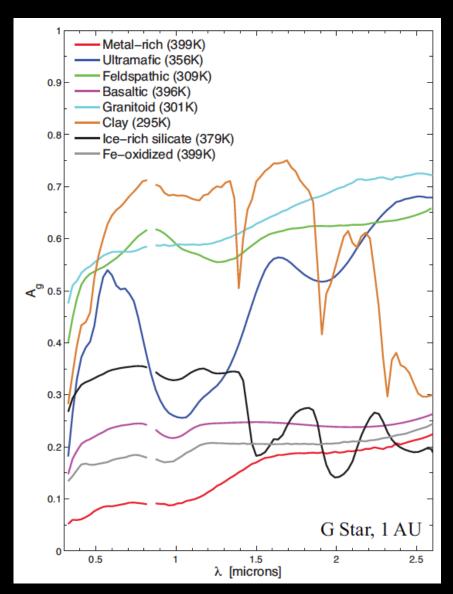


Cowan et al. 2009

Bare-Rock Exoplanets

Rocky exoplanets
 without atmospheres
 have spectral features
 in reflection

 Signature absorption features in 1-2 µm are characteristic of water ice and hydrated minerals



Hu et al. 2012





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