

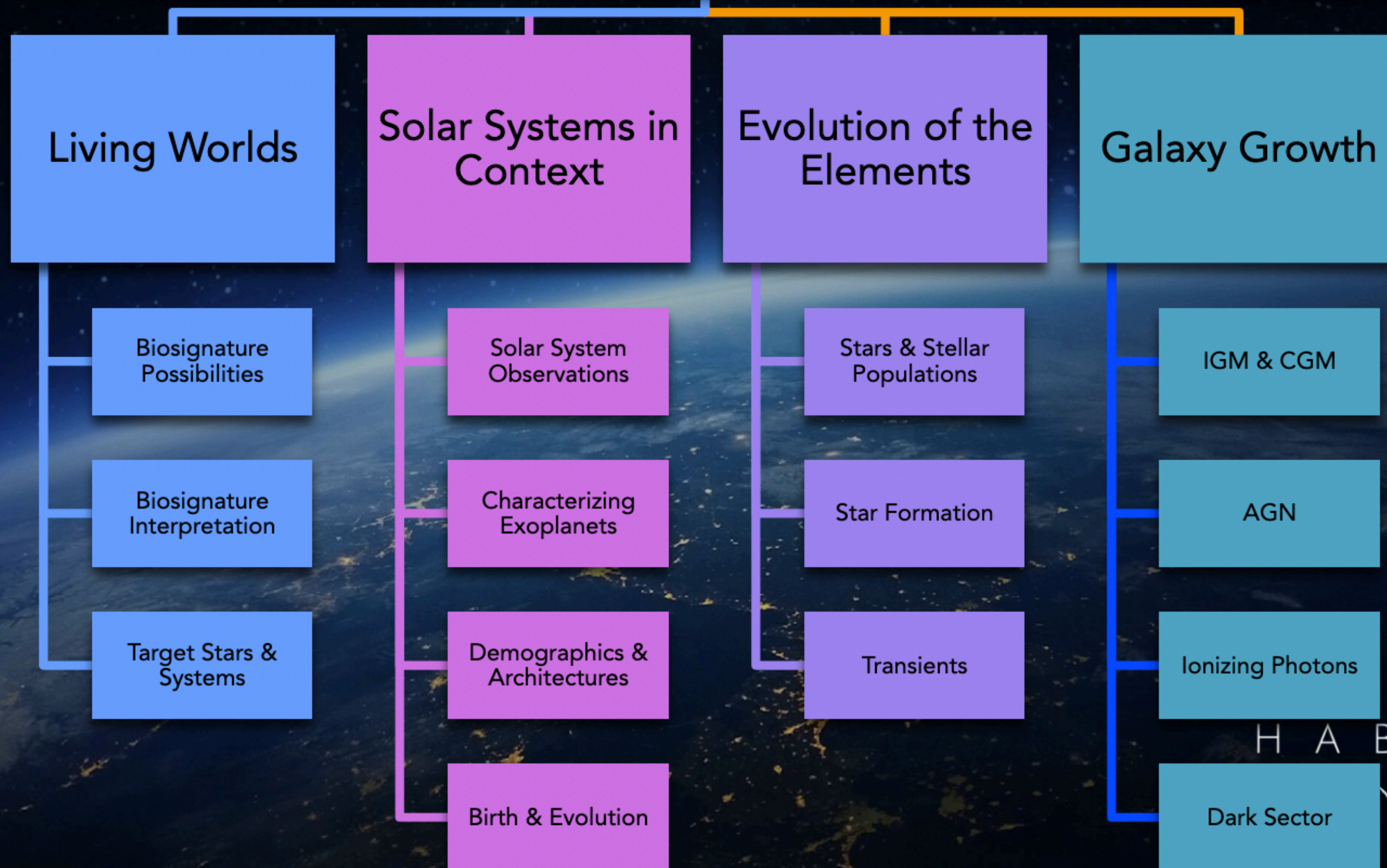
Exoplanet Demographics & Architectures with the Habitable Worlds Observatory

1/11/2025

Malena Rice, on behalf of the HWO D&A sub-WG

H A B I T A B L E
W  R L D S
O B S E R V A T O R Y

Science Working Groups



H A B I T A B L E
W O R L D S
O B S E R V A T O R Y

Solar Systems in Context (SSiC)



Evgenya Shkolnik (Arizona State)
Tyler Robinson (U Arizona)

Ex Officio
Courtney Dressing (Berkeley)



Characterizing Exoplanets
Renyu Hu (JPL)
Michiel Min (SRON)

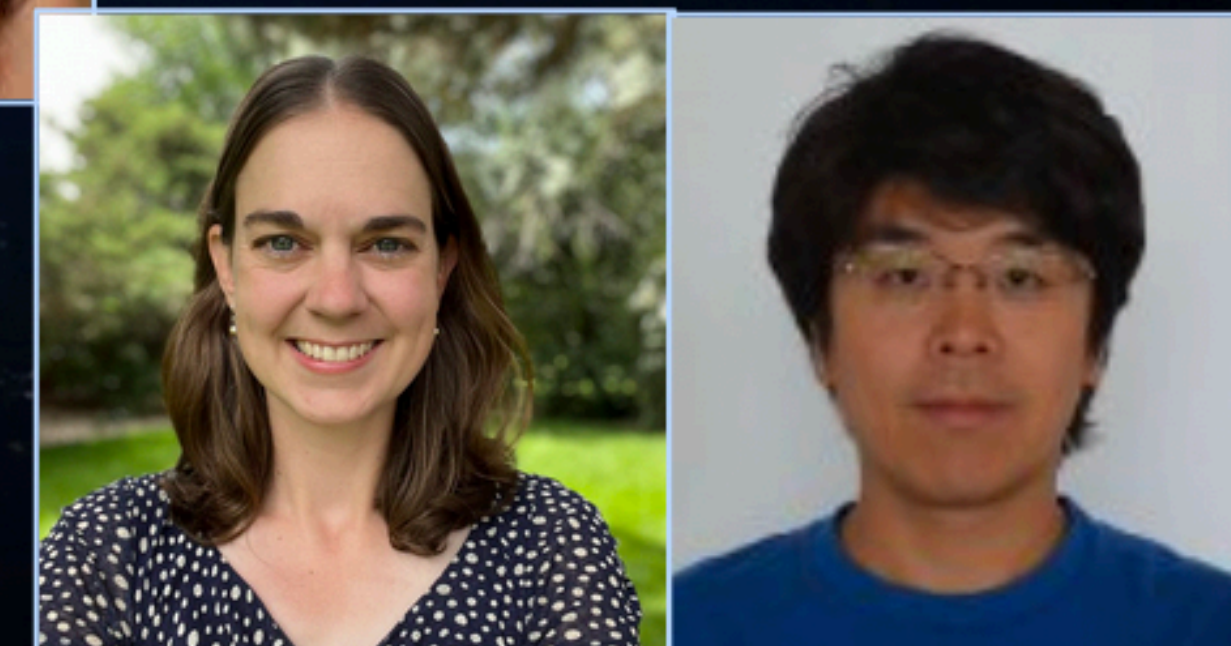
Solar System Observations
Lynnae Quick (GSFC)
Richard Cartwright (JHU/APL)



Demographics & Architectures
Jessie Christiansen (IPAC)
Malena Rice (Yale)



Birth & Evolution
Meredith MacGregor (JHU)
Yasuhiro Hasegawa (JPL)



Solar Systems in Context (SSiC)

Goal:

Develop strategy to answer Solar System and Exoplanet science questions.

- Help set HWO priorities and design.

Deliverables:

Science Case Development Documents (objectives, measurements, etc.)

Structure:

- 4 interdisciplinary science subgroups
- 800 (!!) unique expressions of interest
- 80+ people actively working on 20 SCDDs

Solar System
Observations

The Case for Venus

Mars Origins

Ocean World Habitability

Solar System Origins

Exoplanet Surface Liquid Water

Rocky Worlds vs Sub-Neptunes

Characterizing
Exoplanets

Atmospheric Escape of Small Exoplanets

Reflected Light Spec. of Giant Exoplanets

Transiting Exoplanets

Identify Cold Ocean Worlds

Identify Venus-like Exoplanets

Birth & Evolution

Habitability in Planetary System Context

Protoplanets and Protoplanetary Disks

Debris Disks and their Properties

Disk Winds and Dispersal of Protoplanetary Disks

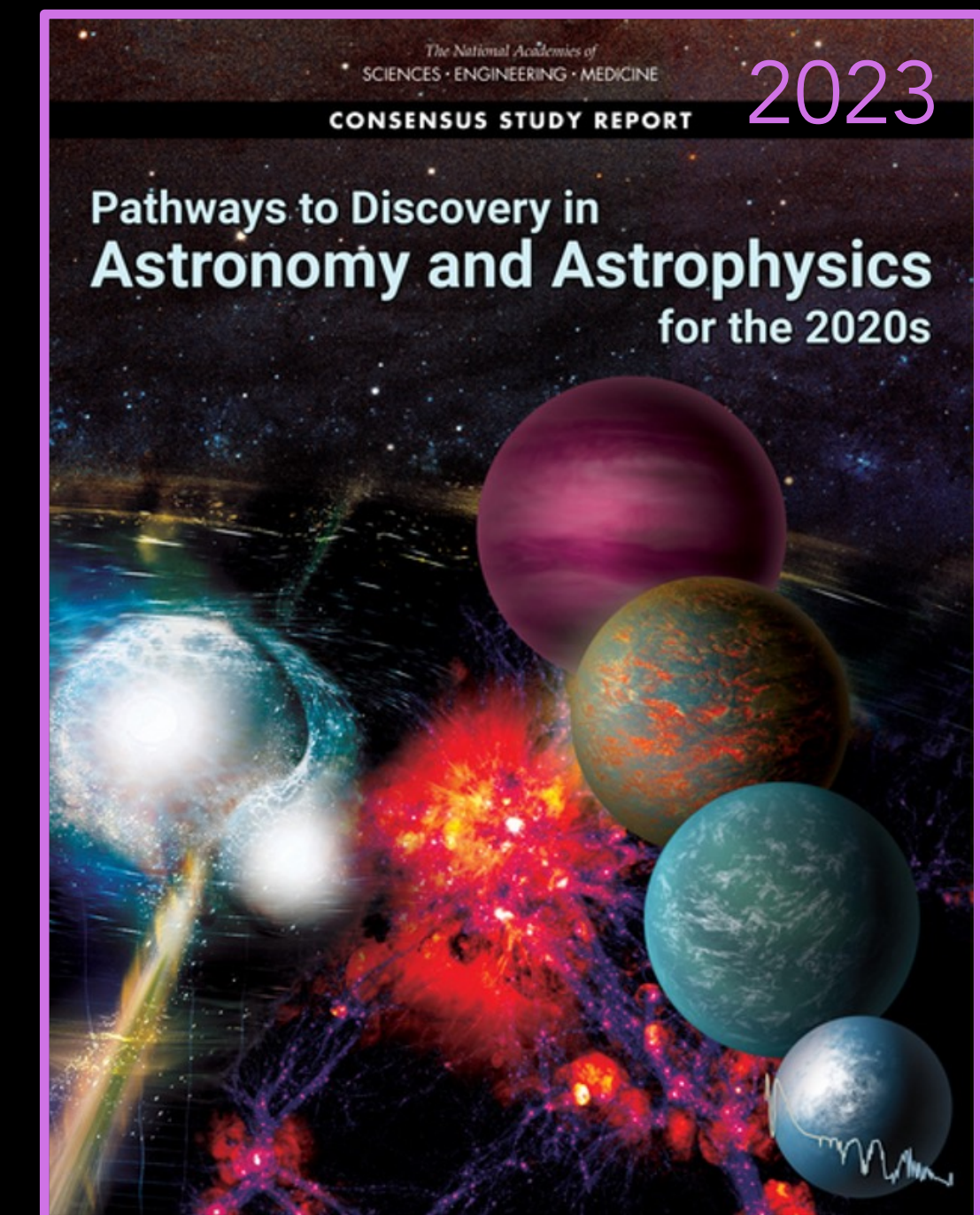
Earth-Like Atmosphere Demographics

Demographics &
Architectures

Giant Exoplanet Orbital Evolution

Occurrence Rates in Binary Systems

Occurrence Rates of Small Exoplanets



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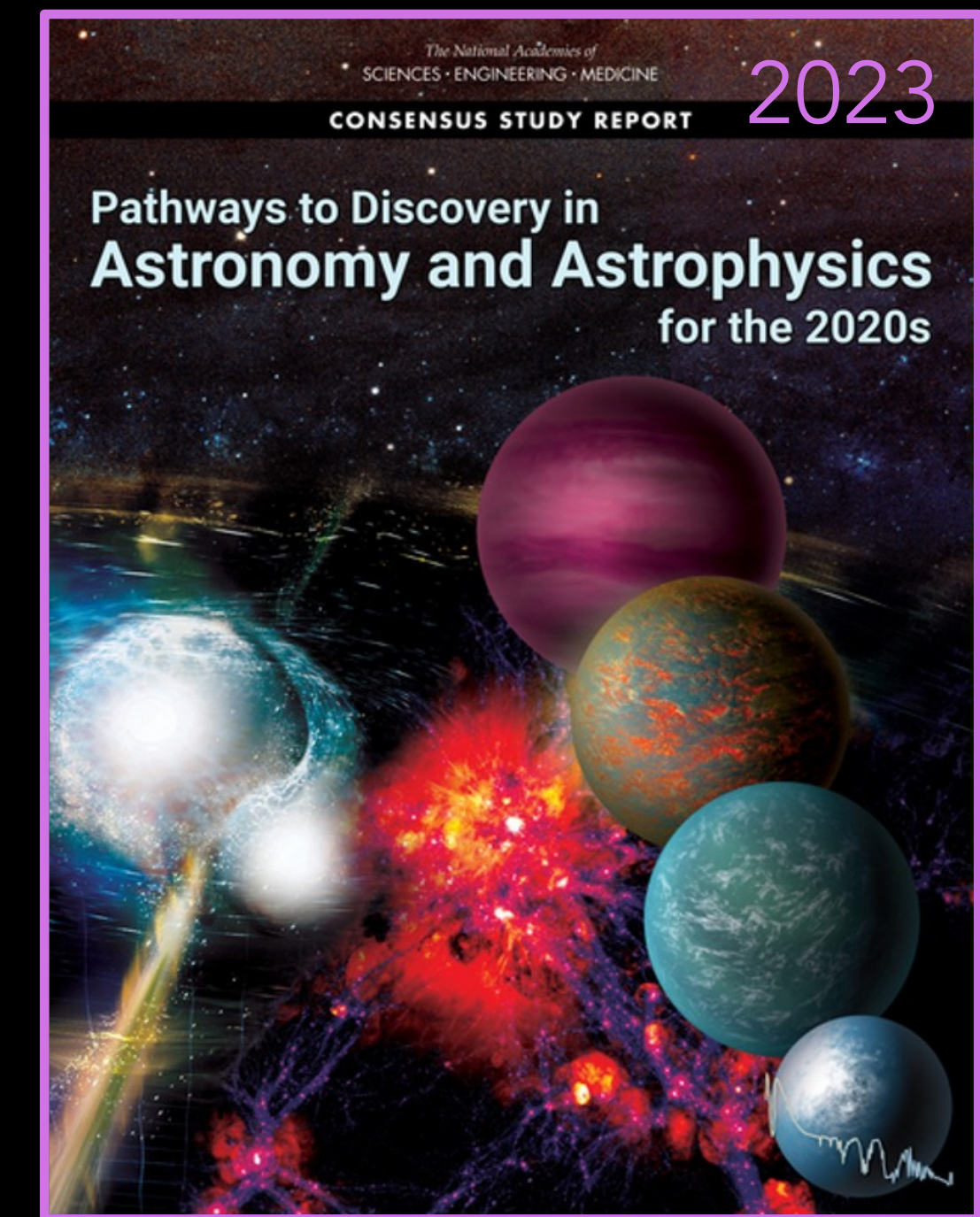
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Occurrence Rates of Small Exoplanets



HWO Exoplanet Demographics & Architectures sub-working group

Goal: synthesize current knowledge of exoplanet occurrence rates and system architectures for the types of stars that HWO will target, and assess the sensitivity and accessibility required to constrain system architectures.



HWO Exoplanet Demographics & Architectures sub-working group

HWO D&A sub-WG co-leads



Malena Rice (Yale)



Jessie Christiansen (IPAC)

Lead: Giant exoplanet orbital evolution



Sabina Sagynbayeva
(Stony Brook)



Stephen Kane
(UCR)

Lead: Occurrence rates in binary systems



Elisabeth Newton
(Dartmouth)

Lead: Earth-like atmosphere demographics



Sarah Blunt (UCSC)



Eric Nielsen (UNM)

Lead: Occurrence rates of small exoplanets



Tansu Daylan (WashU)



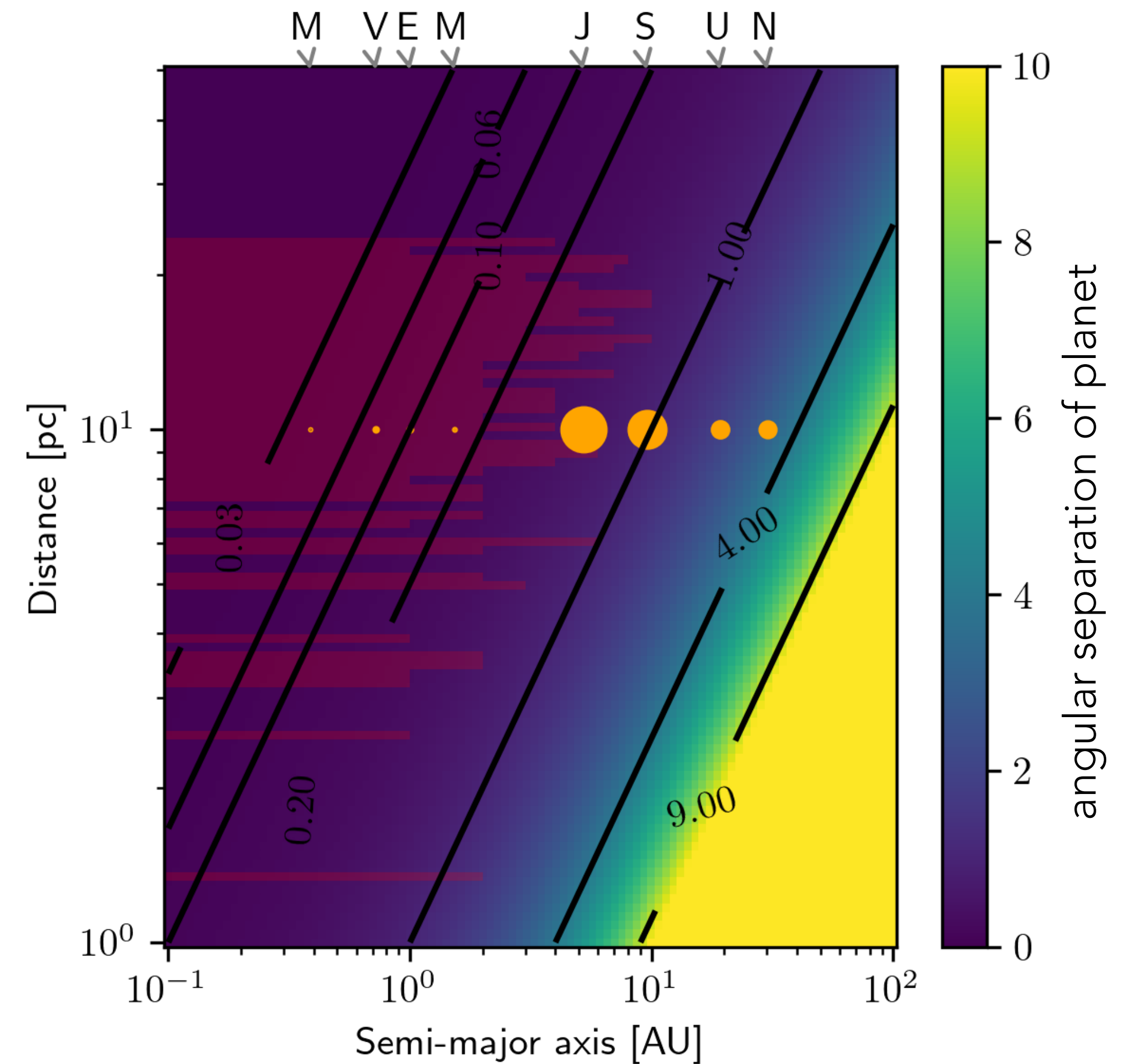
Romy Rodriguez (CfA)

Occurrence rates of small exoplanets

What fraction of small, habitable-zone planets exist in architectures similar to that of the solar system?

Breakthrough requirements:

- 0.6-1.5 microns
- 3-40 λ/D at 0.5 microns
- Contrast of 10^{-11}

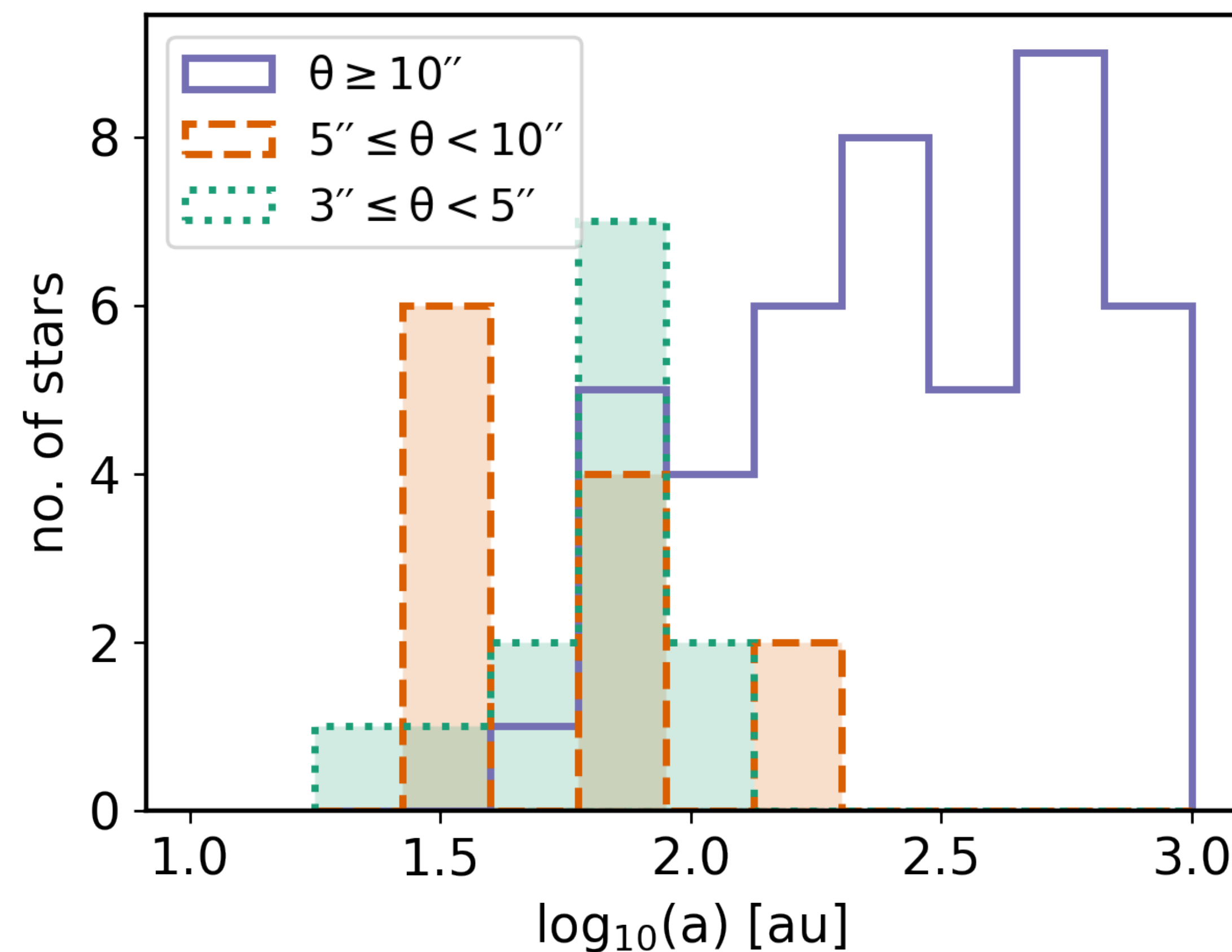


Occurrence rates in binary systems

Is the formation of potentially habitable planets suppressed, enhanced, or unaffected by the presence of a stellar companion?

Breakthrough requirements:

- Starlight suppression for angular separations $\sim 3''$ and targets 15% more distant than those on the ExEP list, or
- Starlight suppression for angular separations $\sim 5''$ and targets 60% more distant
- Census of stellar companions within 30 pc

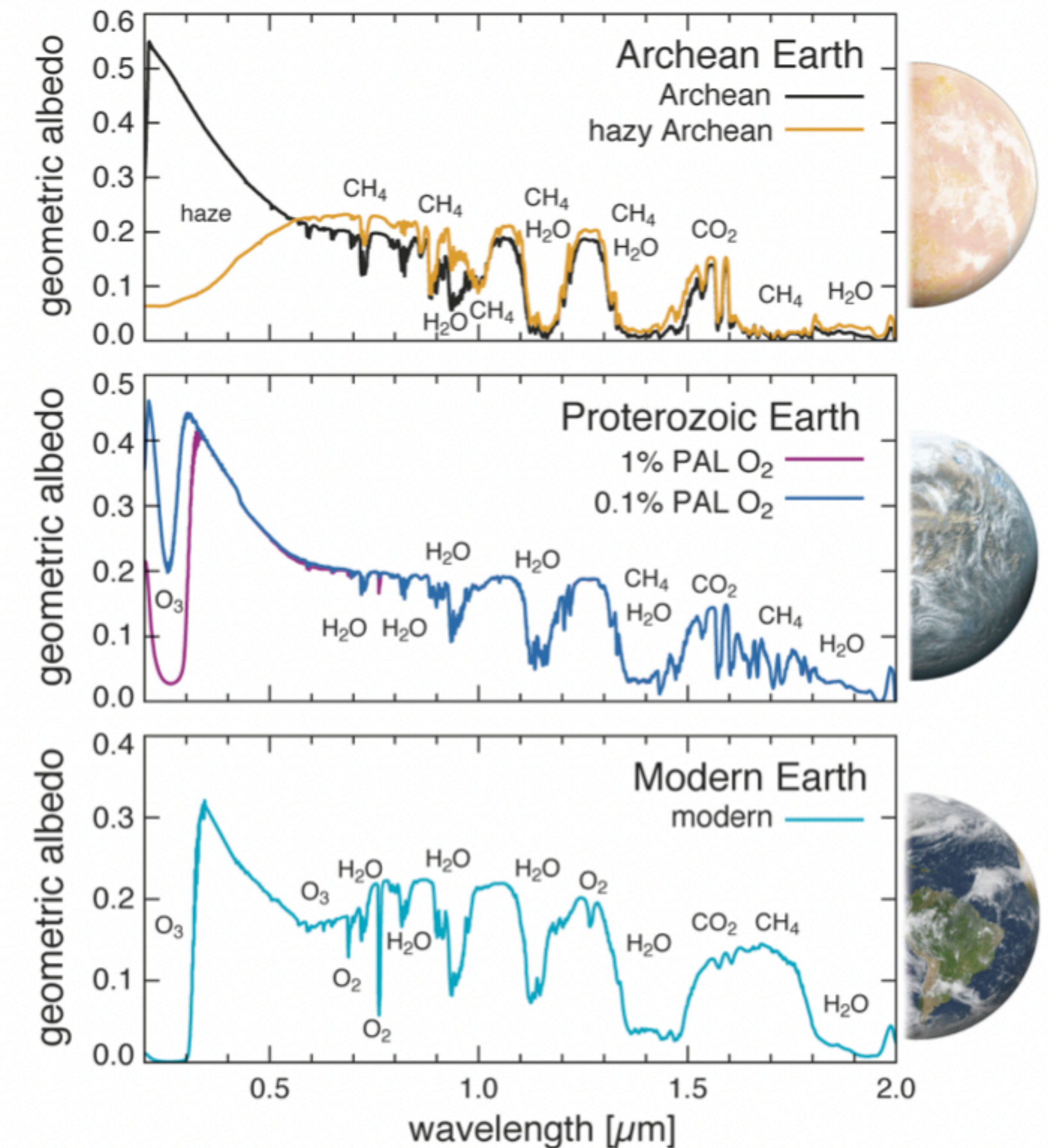


Earth-like atmosphere demographics

How long after it forms does an Earth-like planet typically get an oxygen-rich atmosphere like modern Earth's?

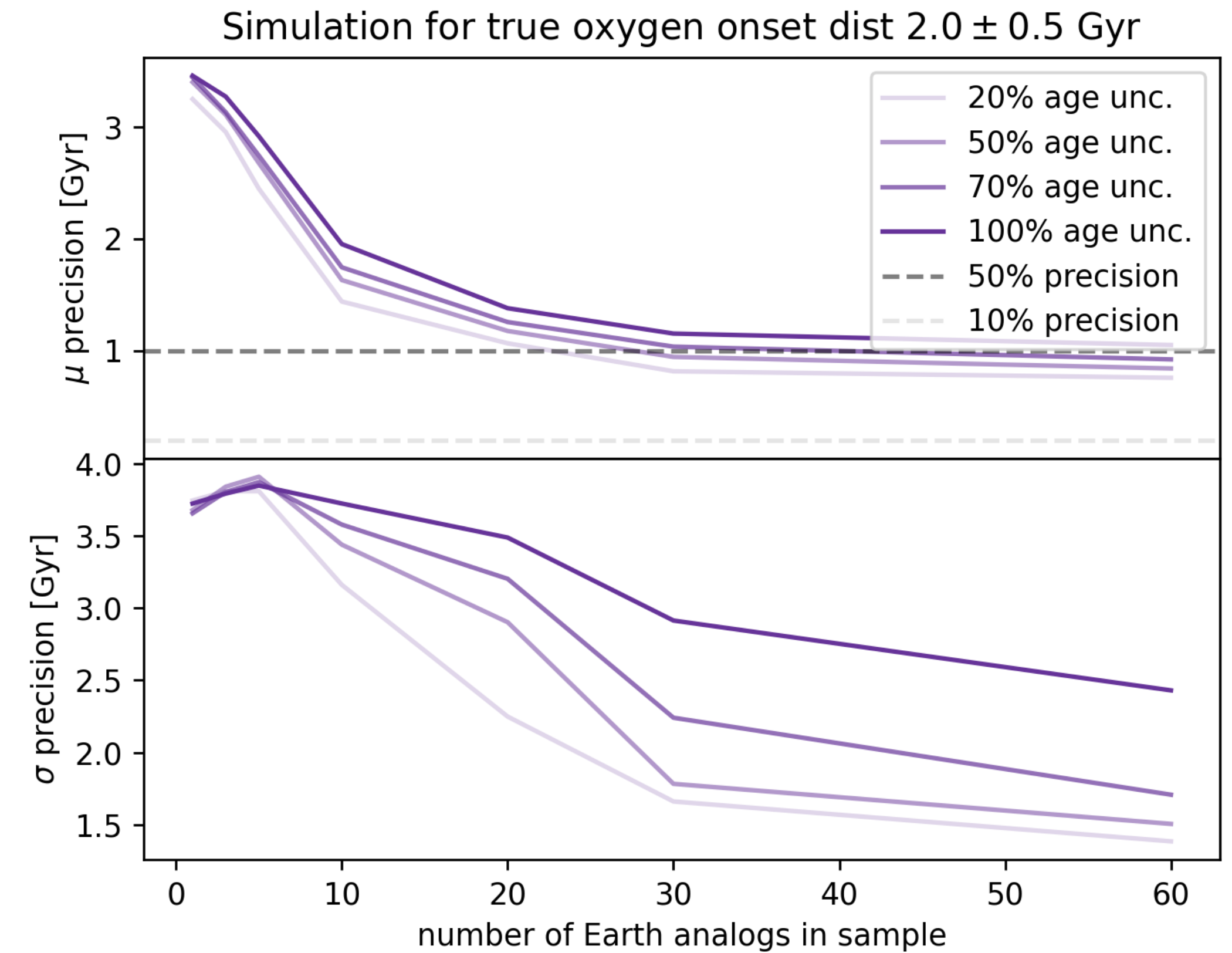
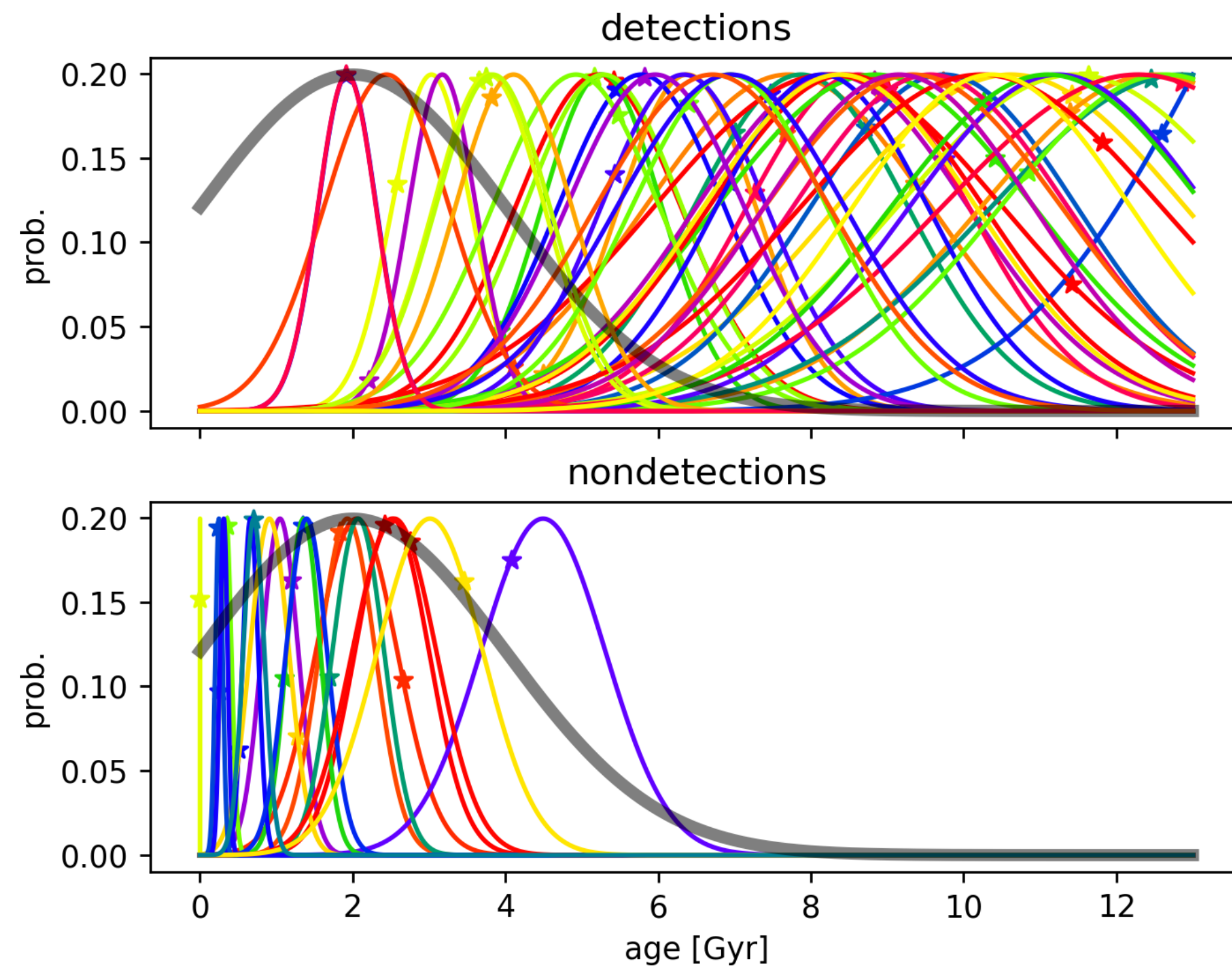
Breakthrough requirements:

- 0.25-2.5 microns
- UV coverage necessary
- 40 detected Earth analogues with detection (or non-detection) of ozone
- 20% age precision for HWO target stars



Lead: Sarah Blunt, Eric Nielsen

Earth-like atmosphere demographics



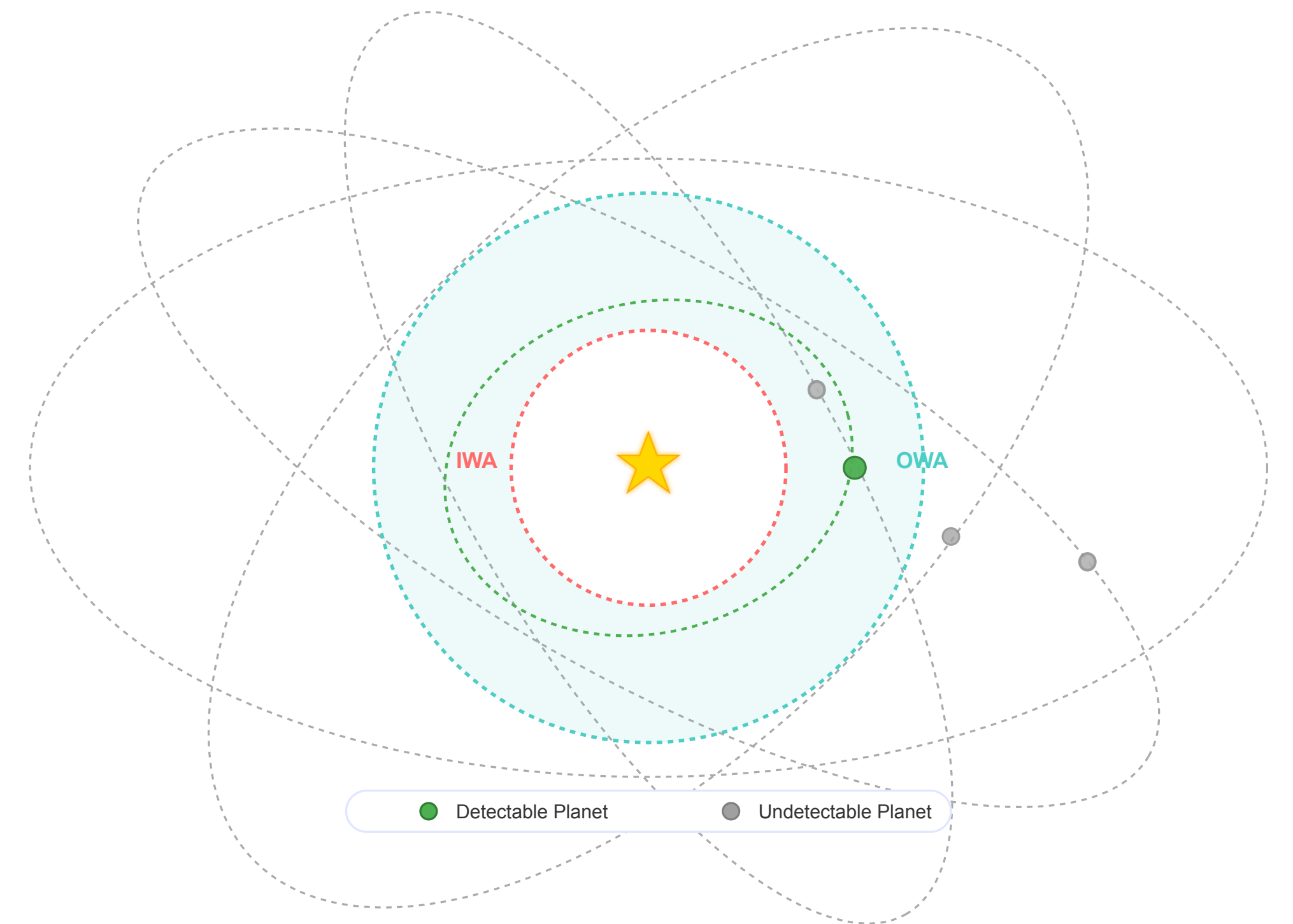
Lead: Sarah Blunt, Eric Nielsen

Giant exoplanet orbital evolution

Which dynamical processes are responsible for the observed orbital architectures of giant planets?

Breakthrough requirements:

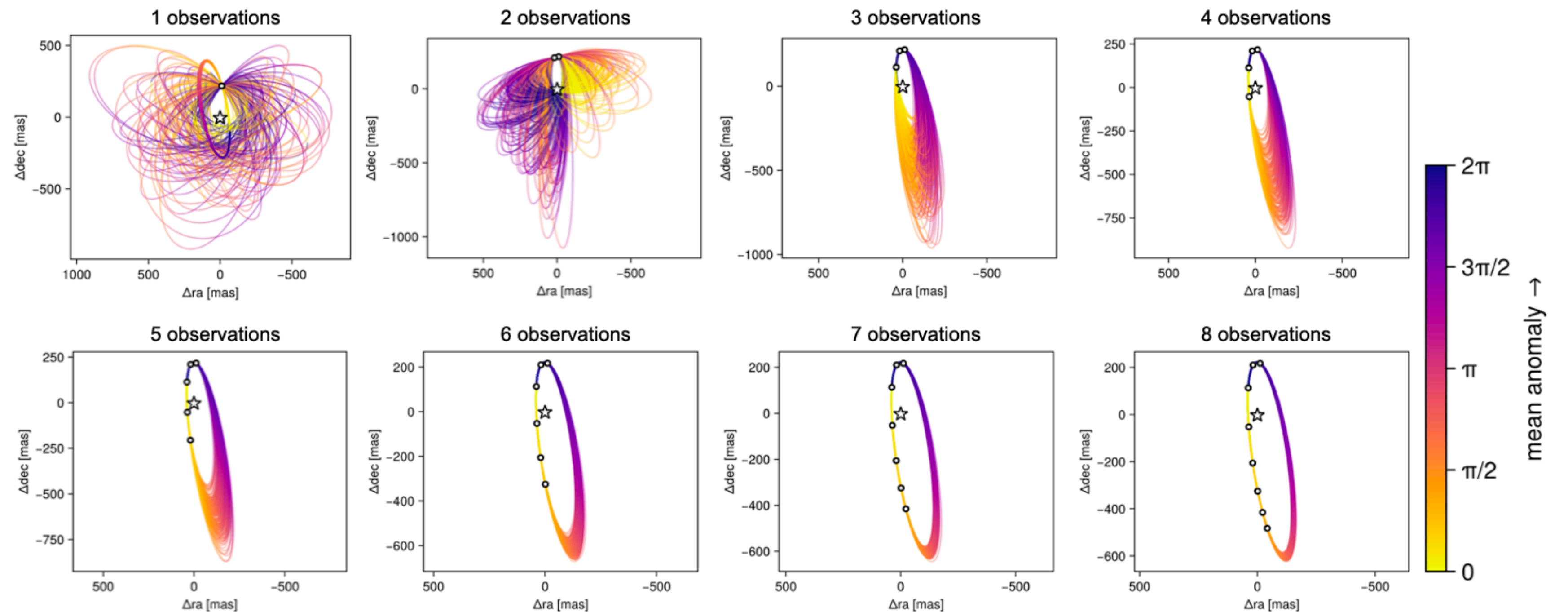
- Sensitivity: < 30.5 mag
- Contrast: 10^{-9} to 10^{-10} at 0.1 - $1''$ Separation
- IWA of $< 0.083''$; OWA of $> 0.64''$



Lead: Sabina Sagynbayeva, Stephen Kane

Giant exoplanet orbital evolution

Generally ~6-8 epochs needed to constrain orbits and provide reliable HZ characterization



Lead: Sabina Sagynbayeva, Stephen Kane

General outlook:

- HWO has the potential to provide transformative advances to not only the detection of habitable zone planets, but also our understanding of system architectures and demographics.
- This will likely require considering a broader target list than the nominal 100-star search: ideally at least ~160-200 stars to consider a diversity of systems and build statistical power.
- Precursor observations are critical to lay the foundation for future HWO observations of bright, nearby stars — ages, extended radial velocity monitoring, stellar companion characterization, etc.

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