Diamonds in the Rough Finding Thousands of New Planet Candidates Around Faint Stars

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NASA/JPL-Caltech

New planets may be rare outcomes of planet formation

or evolution

Introduction

NASA/ESA/CSA/Leah Hustak (STScl)

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> Finding many planets lets us explore exoplanet demographics

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NASA's Transiting Exoplanet Survey Satellite (TESS) 2018 - Present

All-sky survey designed to find exoplanets around nearby, bright stars amenable to follow-up



TESS sky coverage over the Prime Mission (2018 - 2020)













More efficient automated vetting is needed for large-scale TESS planet searches!

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- there is no fully automated vetting pipeline available for TESS planet searches
- what worked for Kepler won't necessarily work for TESS
- humans are inherently biased
- fully automated vetting is needed for robust statistical studies of exoplanet populations (demographics)

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False Positives







False Alarms









Current series of tests, inspired by the Kepler *Robovetter* (Coughlin et al. 2017; Thompson et al. 2018) but fine-tuned for TESS

- Model fits: straight line, trapezoid, transit model, sinusoid
- Model-shift uniqueness tests: assesses the uniqueness of events, including secondaries
- Odd-even depth comparison tests: compare depths of odd/even events
- V-shape test: checks if an object is both grazing and large
- Individual event tests: checks consistency and validity of individual events
- **Difference images:** pixel-level data products to identify off-target signals (in collaboration with **Dr. Steve Bryson**, NASA Ames)
- ... many others ...

The performance of the vetting pipeline has been tested on simulated planets and false alarms:



After implementing these tests into the pipeline:



Reviewed all signals and refined vetting algorithm with fellow postdoc Tansu Daylan (MIT/Princeton)





Results



Results

~2000 hot Jupiter TOIs



Results



Results

Several of these confirmed giants orbit M dwarfs





TOI-3884 b: First sub-Saturn found transiting a mid-M dwarf (Almenara [+ Kunimoto] et al. 2022)



TOI-3884 b: First sub-Saturn found transiting a mid-M dwarf (Almenara [+ Kunimoto] et al. 2022)

TOI-5205 b: First Jupiter found transiting a mid-M dwarf (Kanodia et al. 2023)

Astronomers Discover Bizarre "Forbidden" Planet That Should Not Exist

Scitech Daily

Katharine Cain/Carnegie Institution for Science



TOI-3884 b: First sub-Saturn found transiting a mid-M dwarf (Almenara [+ Kunimoto] et al. 2022)

TOI-5205 b: First Jupiter found transiting a mid-M dwarf (Kanodia et al. 2023)

TOI-3757 b: Lowest density transiting planet known to orbit an M dwarf (Kanodia et al. 2022)

Katharine Cain/Carnegie Institution for Science





TOI-4010: Metal-rich K-dwarf ([Fe/H] = +0.37)

Highlighted Discoveries













Cool things about TOI-4010

1. Comparative Planetology

TOI-4010 has THREE large planets in close-in orbits, all with atmospheres



Mass-radius curves from Zeng et al. (2018)

20 - 104 10 · 10³ Planet radius (R_®) Š TOI-4010 d flux 10² TOI-4010 c Insolation 101 3 TOI-4010 b - 10⁰ 0 10^{-1} 10² 10^{0} 10^{1} Orbital period (days)

Planets in the NASA Exoplanet Archive with density measured to better than 50% precision

Cool things about TOI-4010

2. A planet in the hot Neptune desert

TOI-4010 b is one of the only hot Neptunes in a multi-planet system

Cool things about TOI-4010

3. Particularly massive sub-Saturns

TOI-4010 d is one of the most massive sub-Saturns found in a multi-planet system



Mass-host star metallicity diagram for sub-Saturns (4 - 8 R)

Cool things about TOI-4010

4. Moderately eccentric super-Jupiter companion

Still monitoring RVs to complete the orbit and determine the correct period



Summary & Future Plans



- automated vetting is needed for large-scale planet searches
- I'm close to a fully automated TESS vetting pipeline, with applications to both planet searches and exoplanet demographics
- will be made publicly available for use by the exoplanet community
- 2922 TOIs so far many more upcoming!
 - several TOIs are rare outcomes of planet formation and promising targets for follow-up