

# Don't Heckle My Speckle: A Coronagraph Design Study for the SEAL testbed

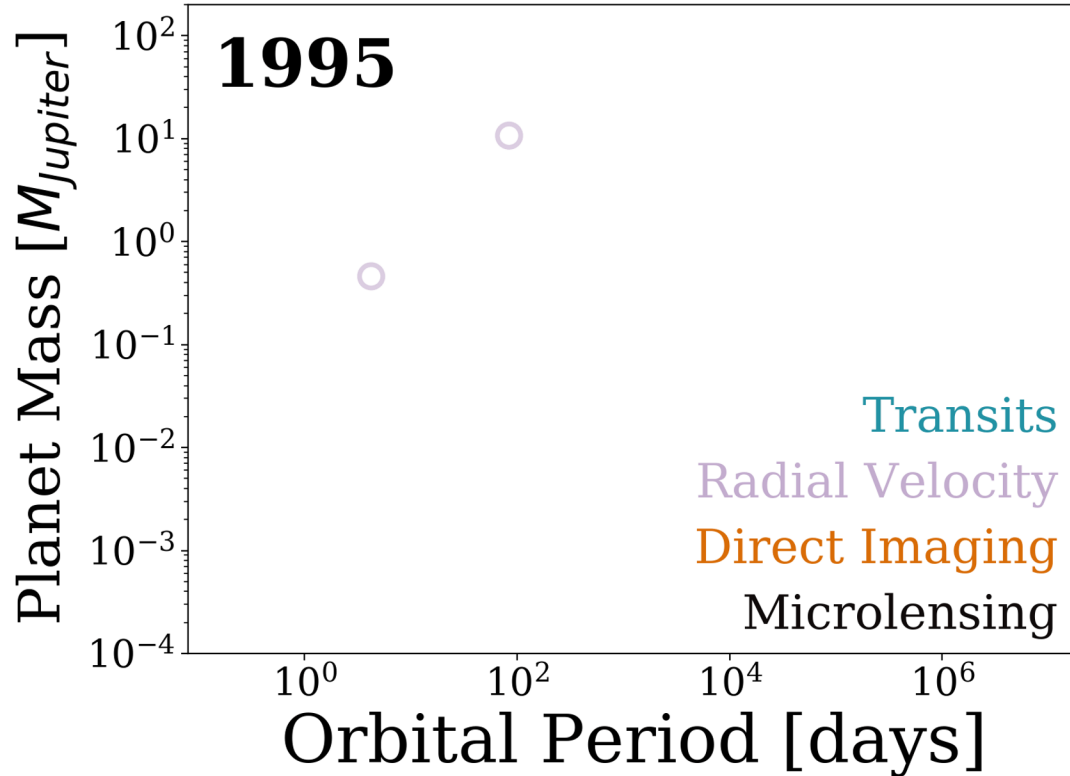
Jules Fowler . ExoExplorer Seminar Series . 5/14/21

# Talk Outline

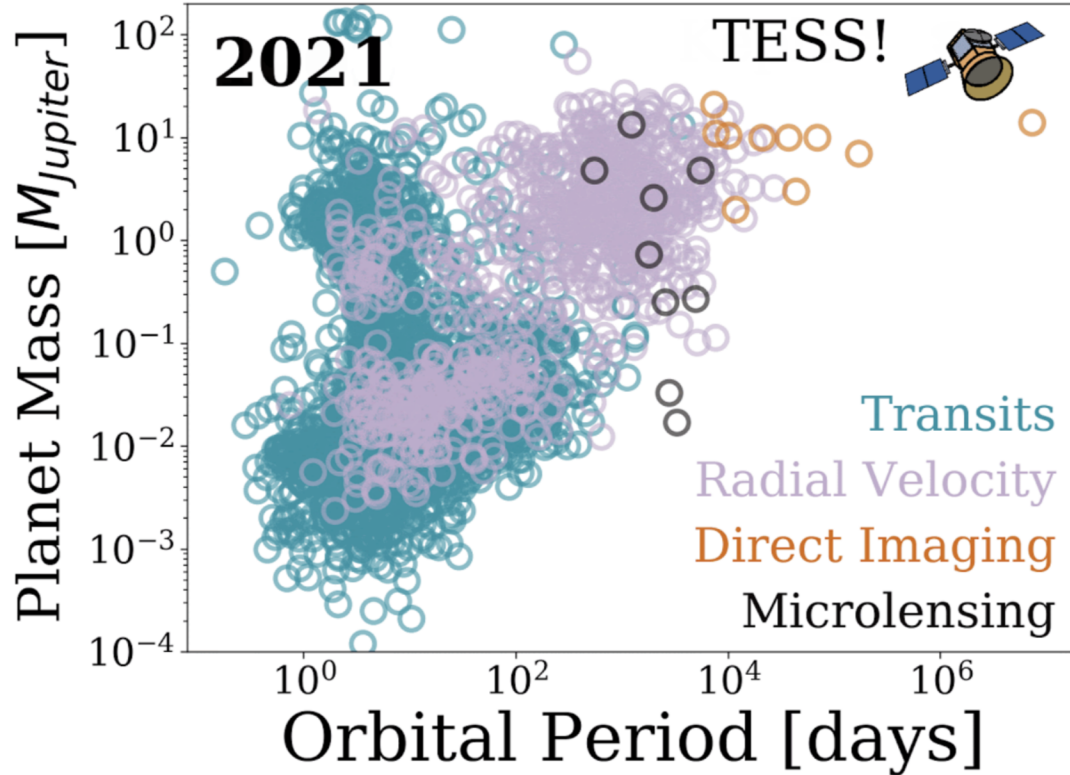
- Extreme Adaptive Optics
  - SEAL testbed
  - Optics refresher
  - Coronagraphy primer
  - Simulations
  - Current state of the coronagraph design
  - Next steps
-

# Direct Imaging and Extreme Adaptive Optics

# Directly Imaged Planets are Few and Far Between

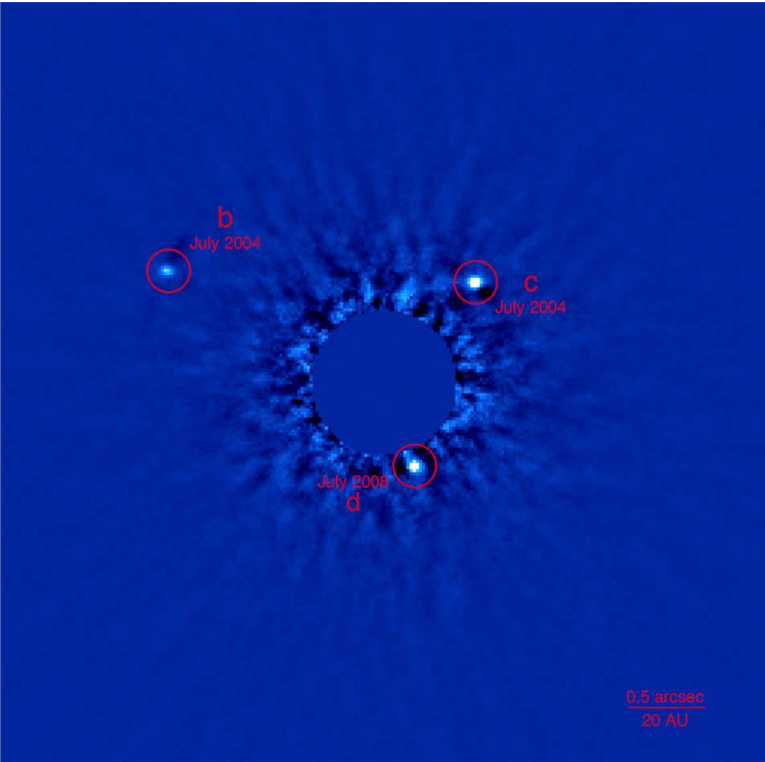


# Directly Imaged Planets are Few and Far Between



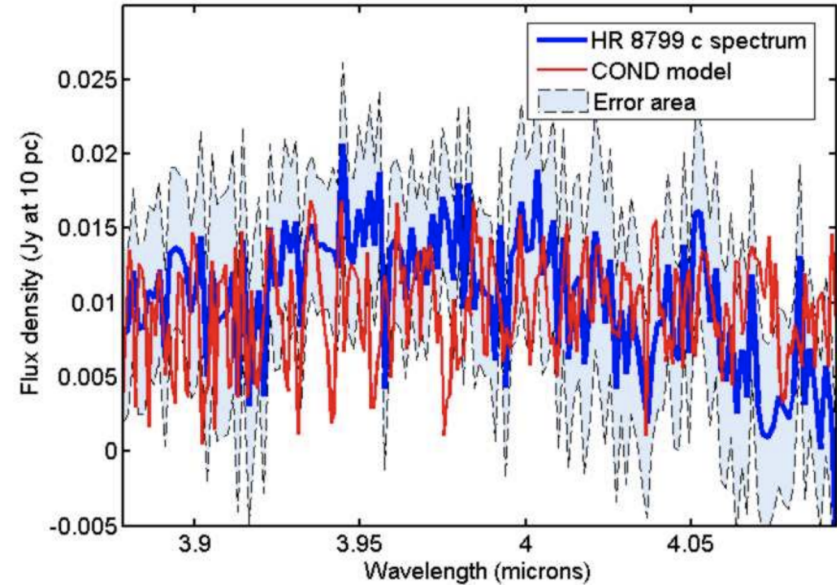
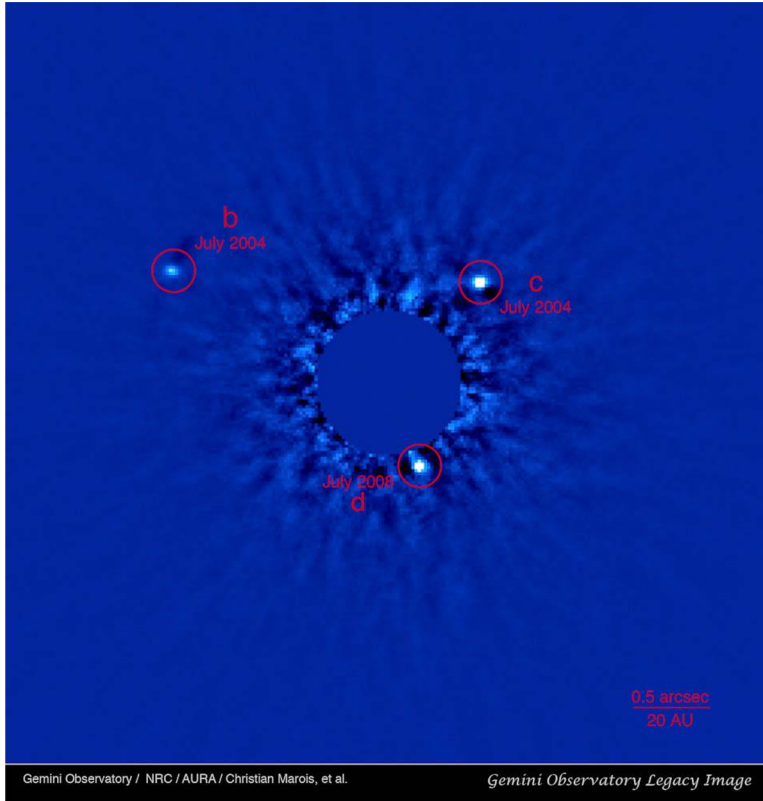
# Direct Imaging Resolves the Light of a Planet

## HR8799 Planetary System



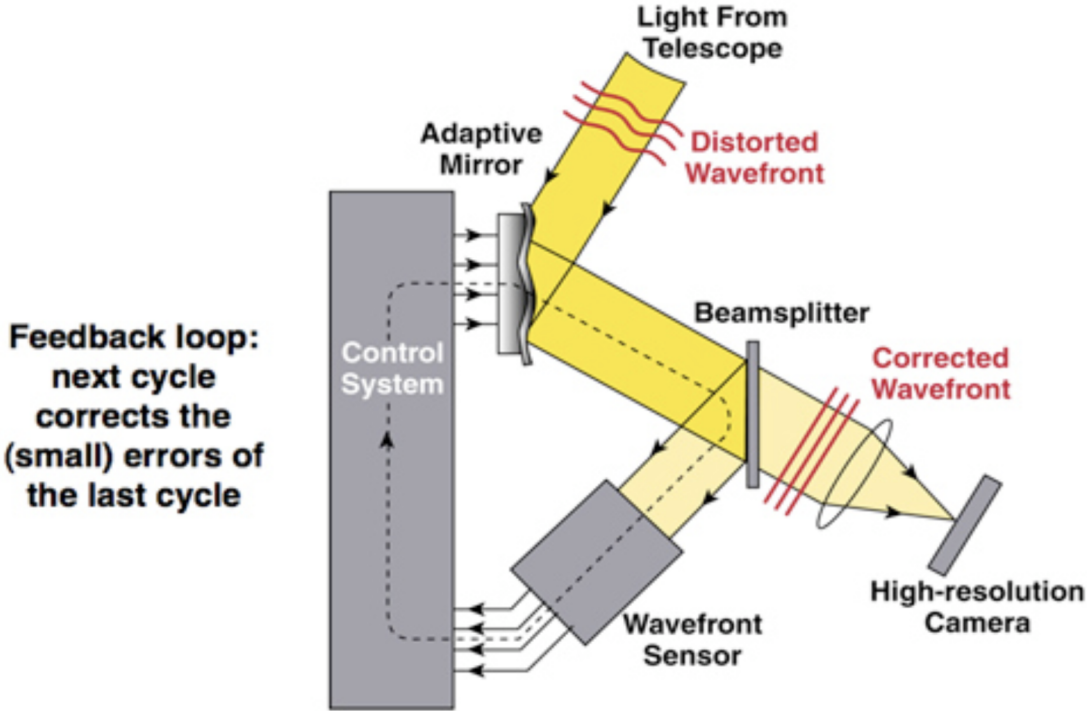
# Direct Imaging Unlocks Exoplanet Characterization

## HR8799 Planetary System



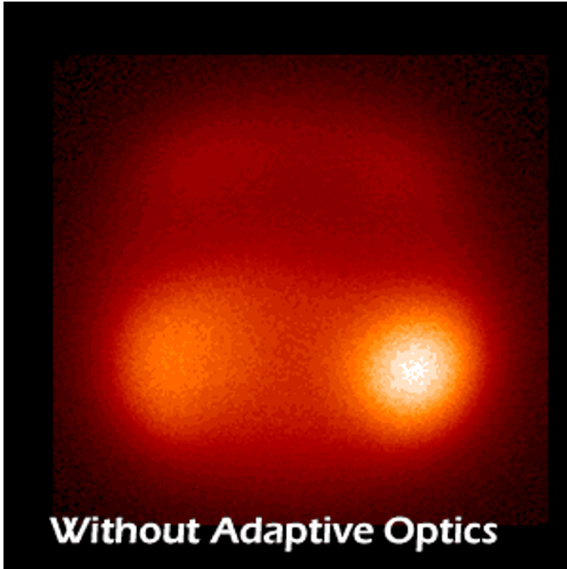
(Janson, 2010)

# Ground-based Imaging Requires Adaptive Optics

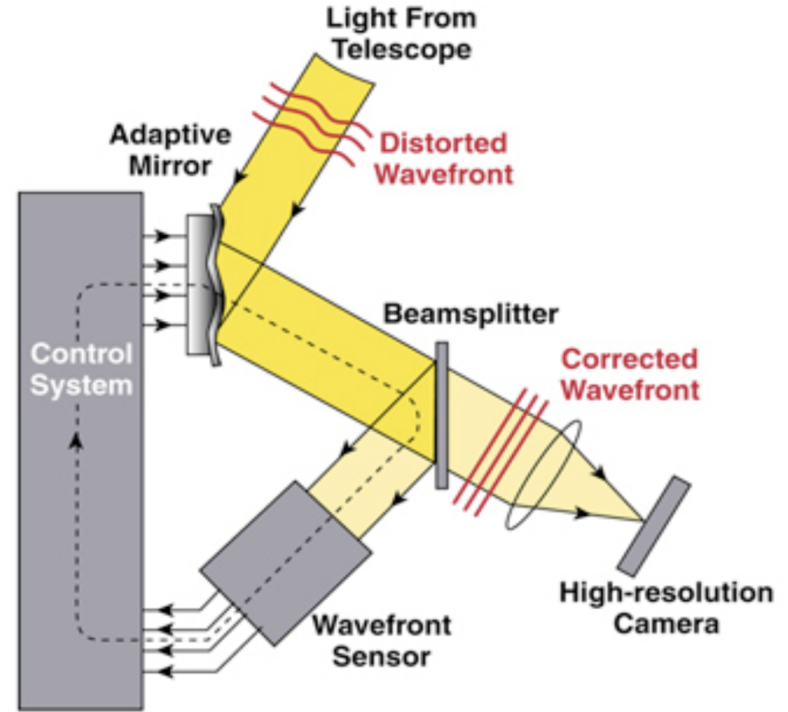




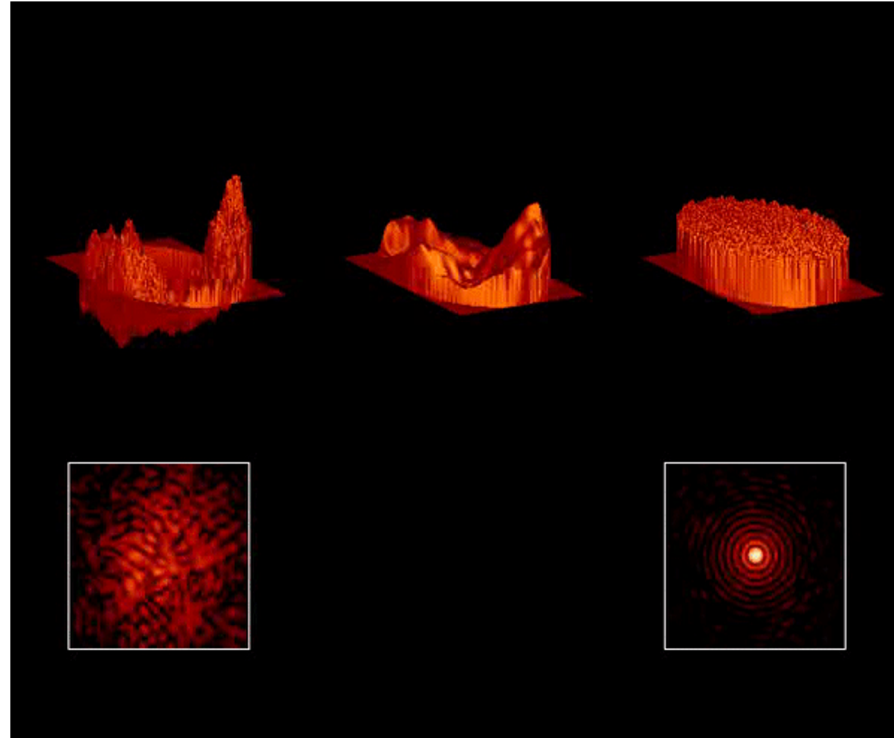
# Ground-based Imaging Requires Adaptive Optics



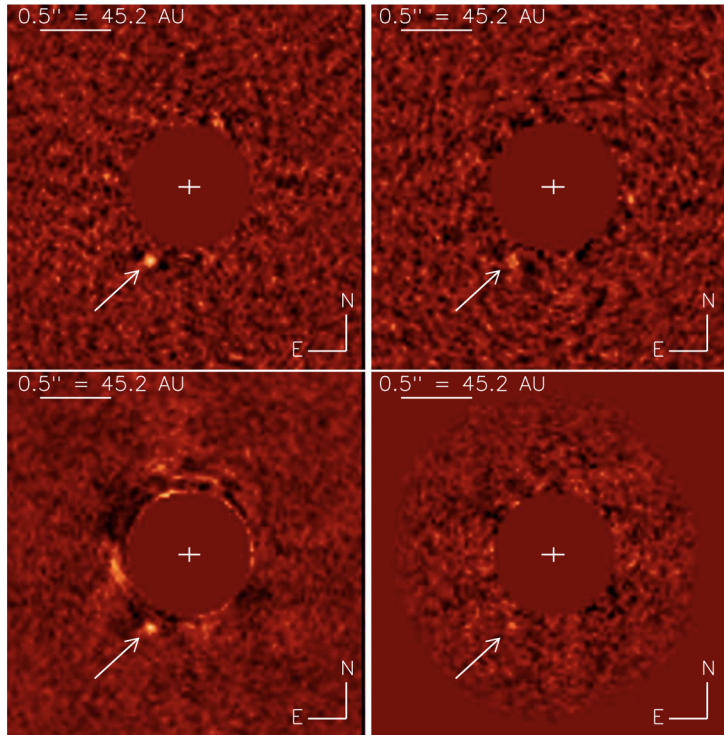
Feedback loop:  
next cycle  
corrects the  
(small) errors  
of the last cycle



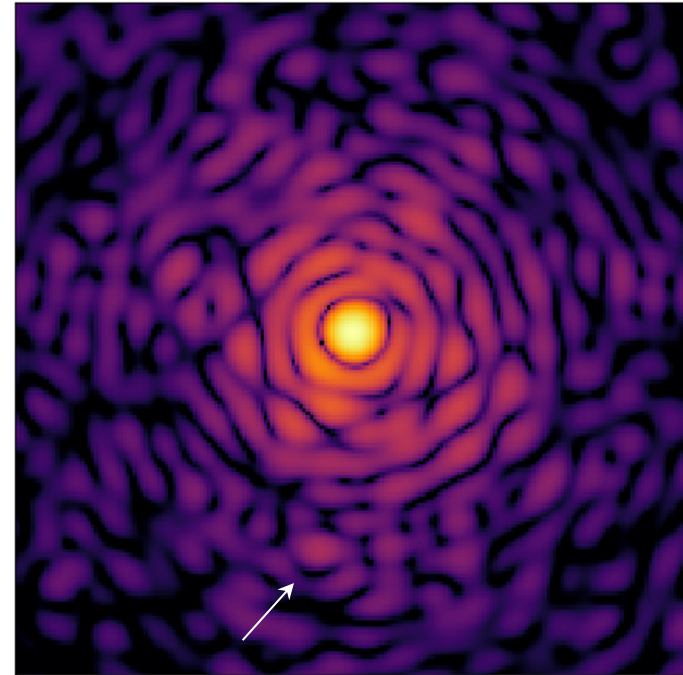
# Ground-based High Contrast Imaging Requires **EXTREME** Adaptive Optics



# HD 95086b and a Speckle are Hard to Distinguish



(Rameau, 2013)



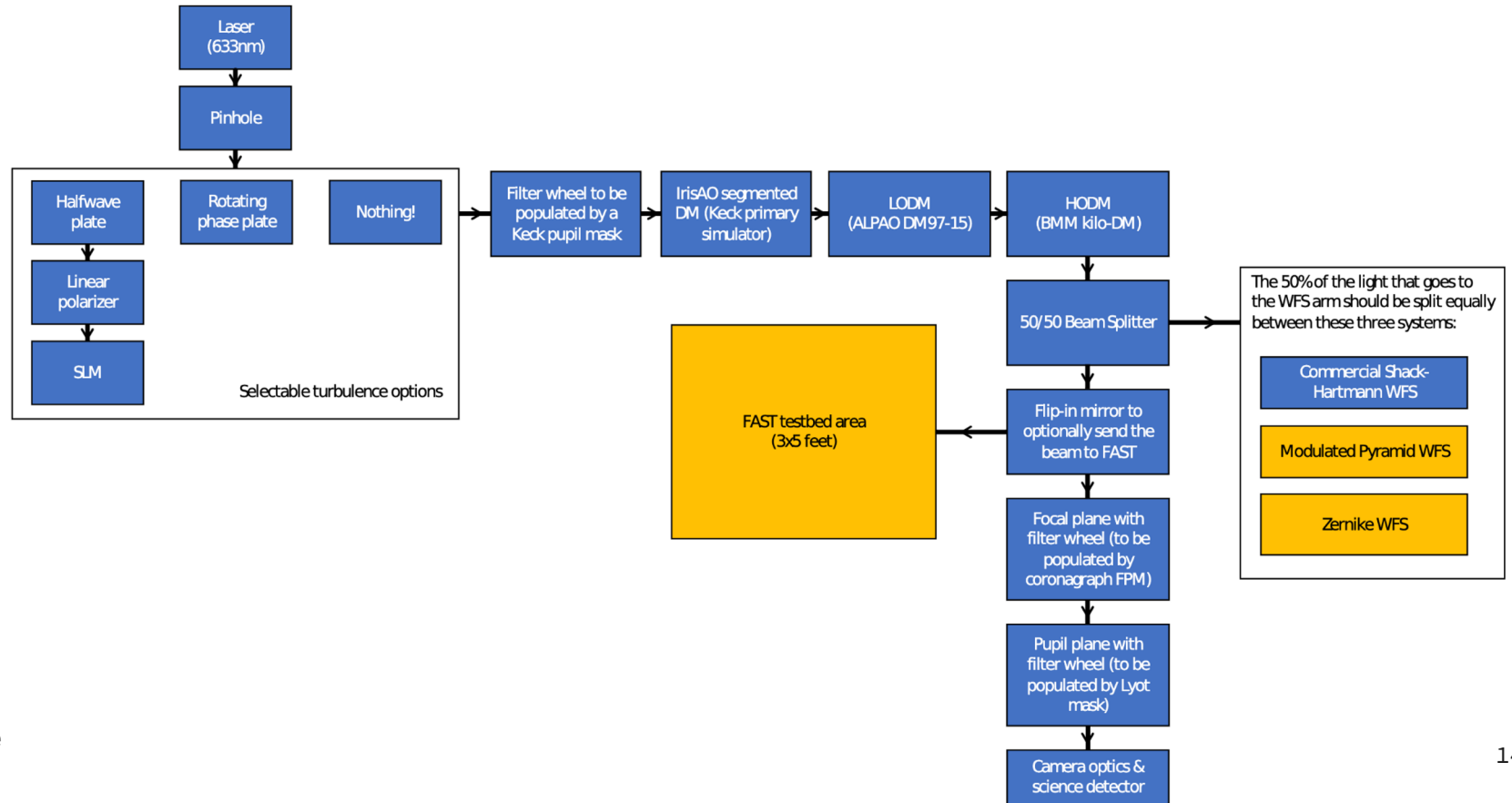
The Santa  
Cruz Extreme  
AO Laboratory  
(SEAL)

# SEAL Tests Novel Wavefront Sensing, Wavefront Control, and Coronagraphy in Synergy with Keck

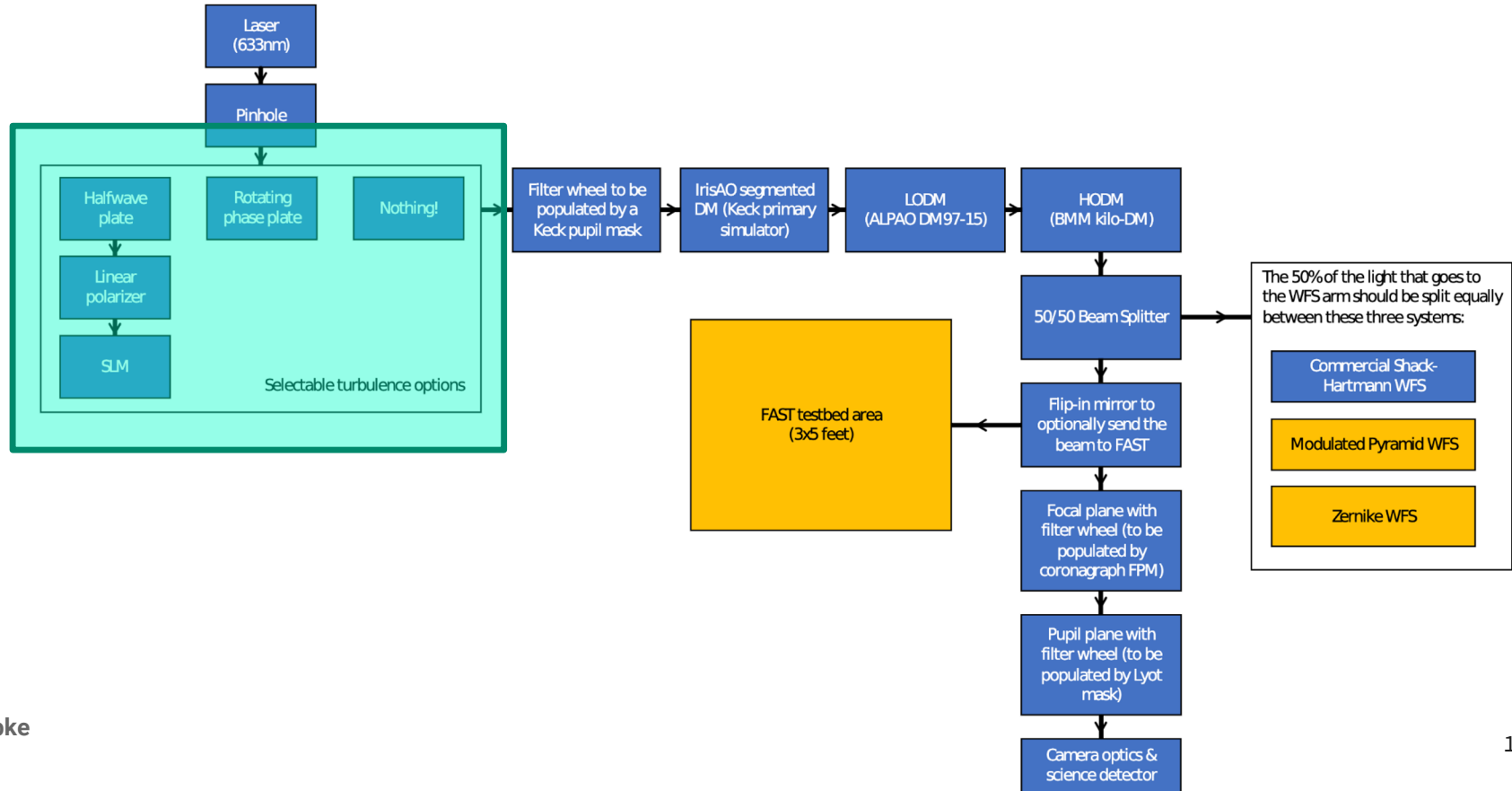


- Part of the UCSC Lab for Adaptive Optics
- PI: Rebecca Jensen-Clem
- Supported by Renate Kupke, Daren Dillon, and Sylvain Cetre
- Predictive wavefront control (Maaïke Van Kooten)
- Focal plane wavefront sensing (Benjamin Gerard)
- Pyramid wavefront sensing (Dominic Sanchez)
- Zernike wavefront sensing (Maaïke Van Kooten, soon to be Maïssa Salama, and Jules Fowler)
- **Coronagraph Design (Jules Fowler)**

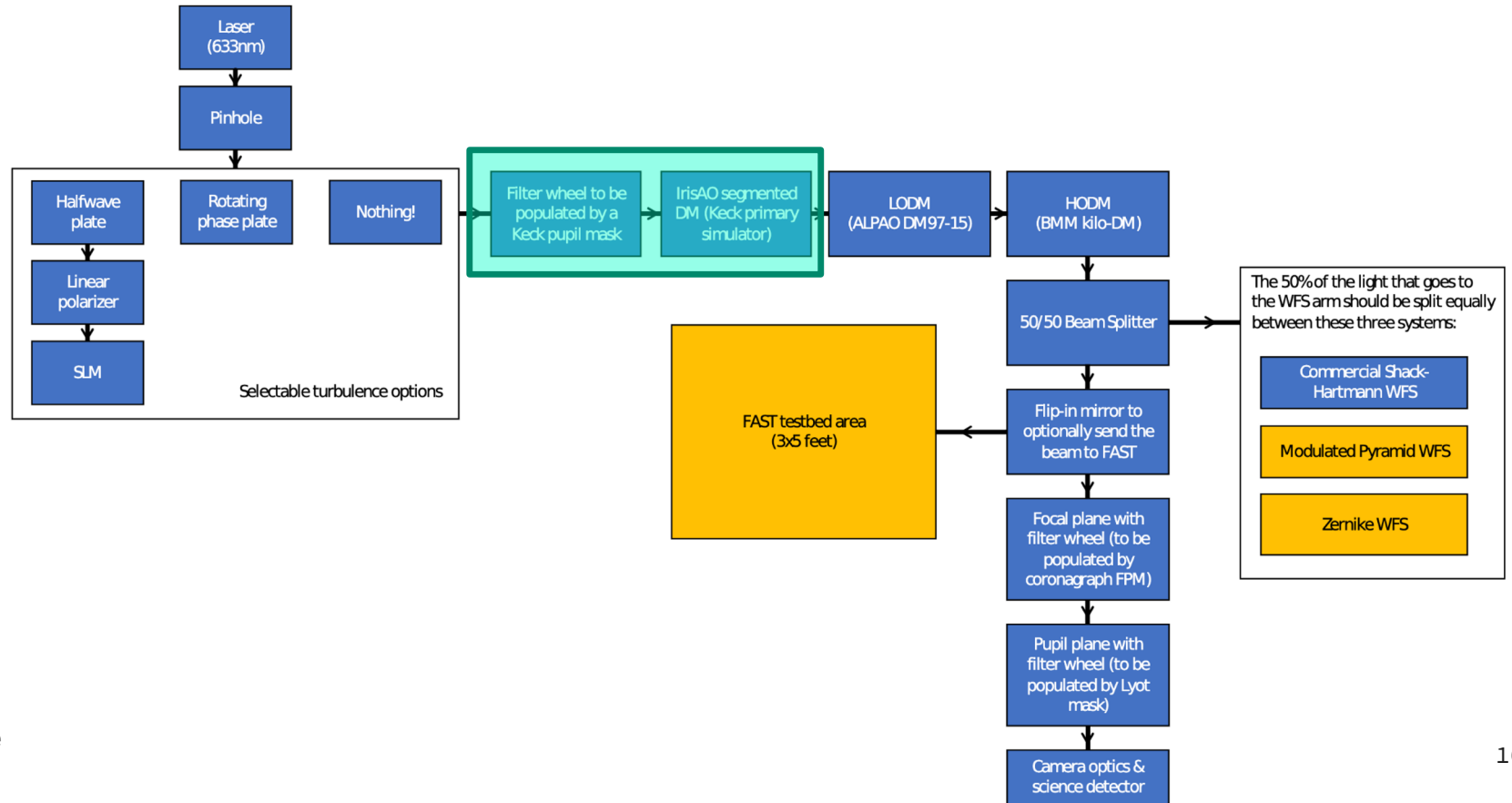
# SEAL Includes ...



# SEAL Includes Atmospheric Turbulence Generation

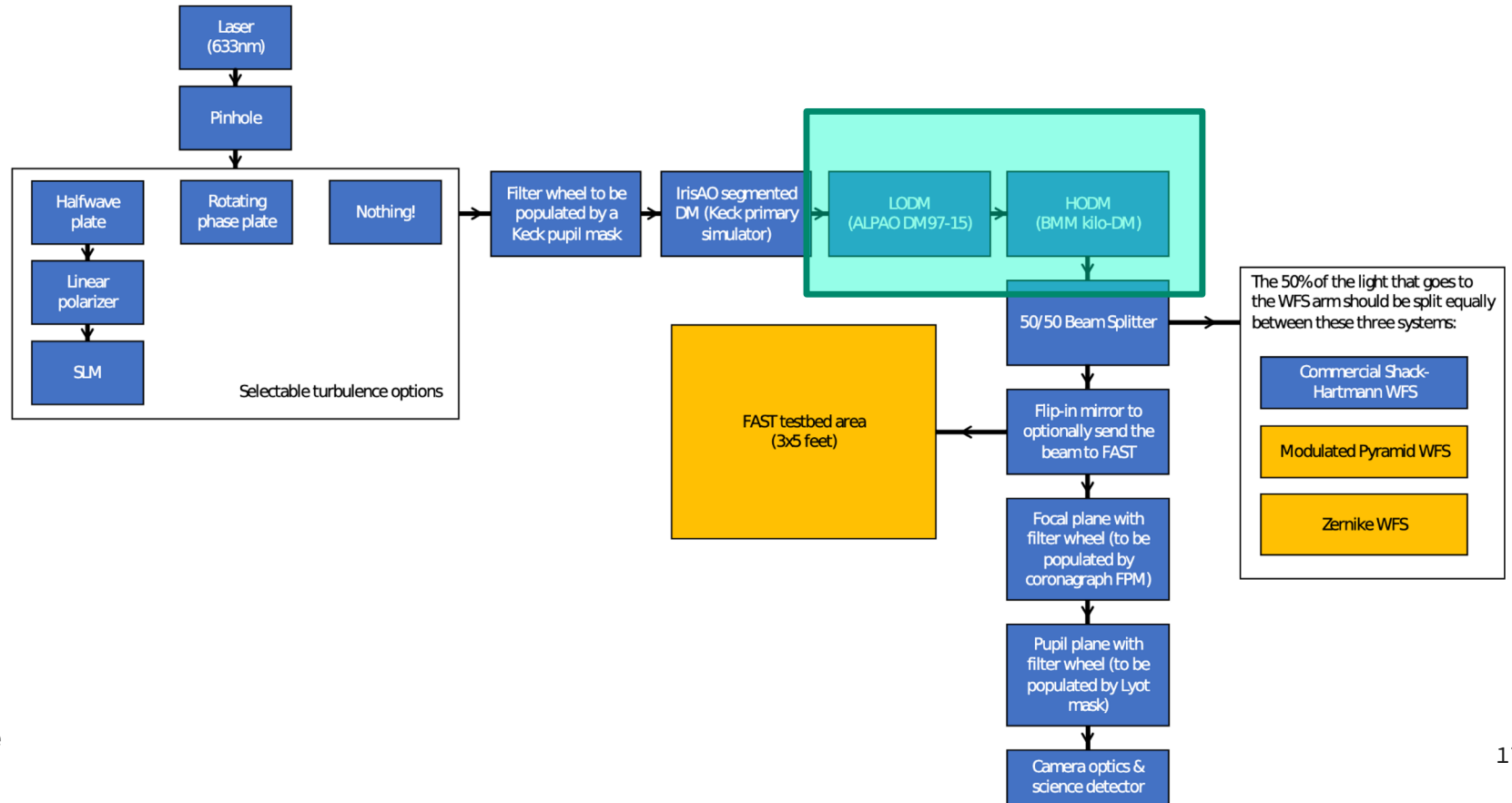


# SEAL Includes Keck-like Pupil Shaping

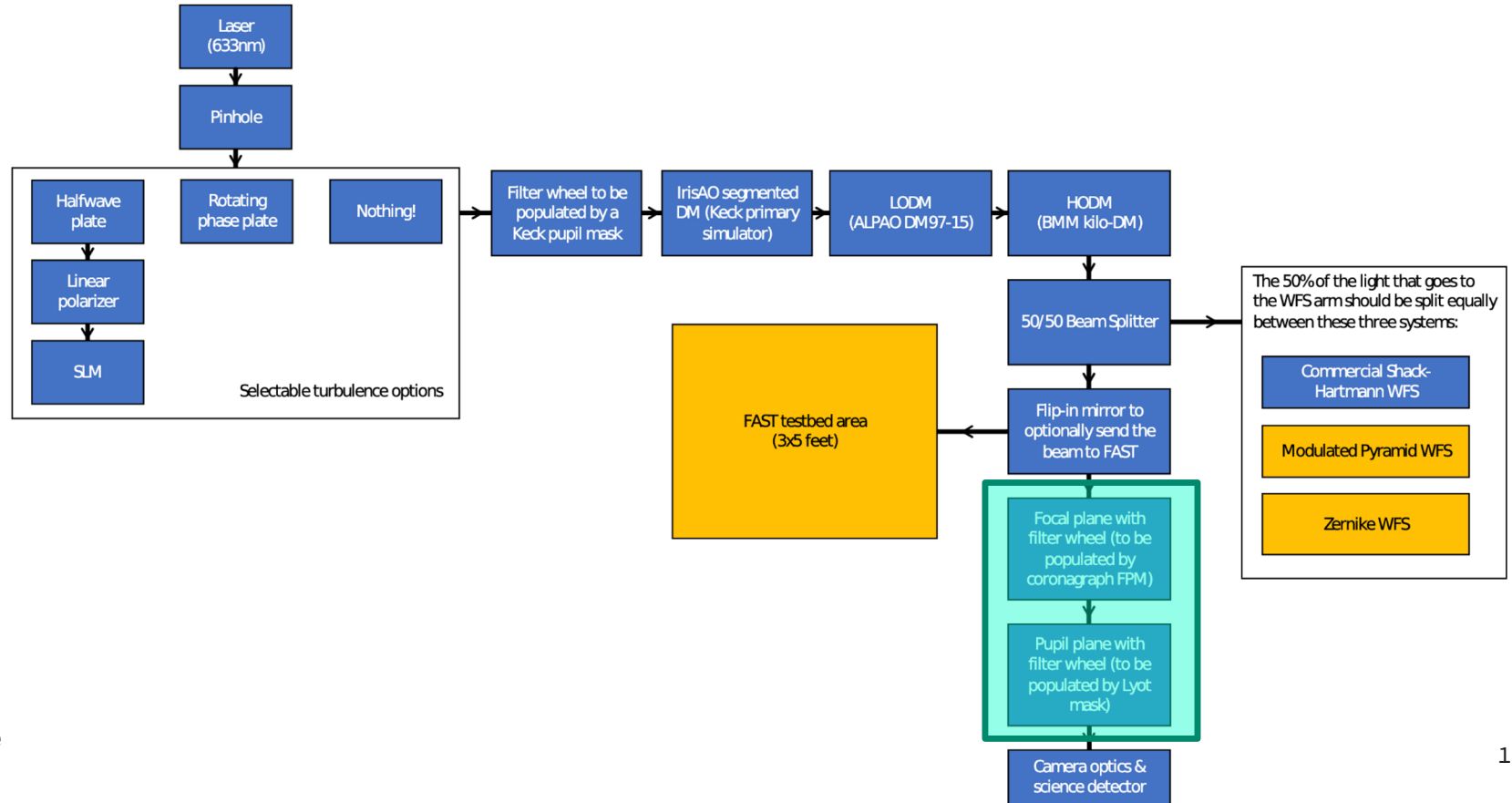




# SEAL Includes High/Low Order Deformable Mirrors



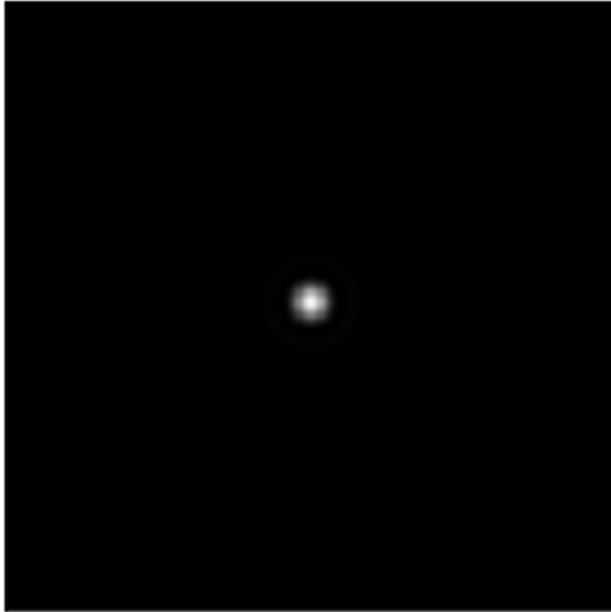
# SEAL Includes This Very Coronagraph!



Everything You  
Wanted to  
Forget About  
Optics

# Pupil and Focal Images are a Fourier Transform Apart

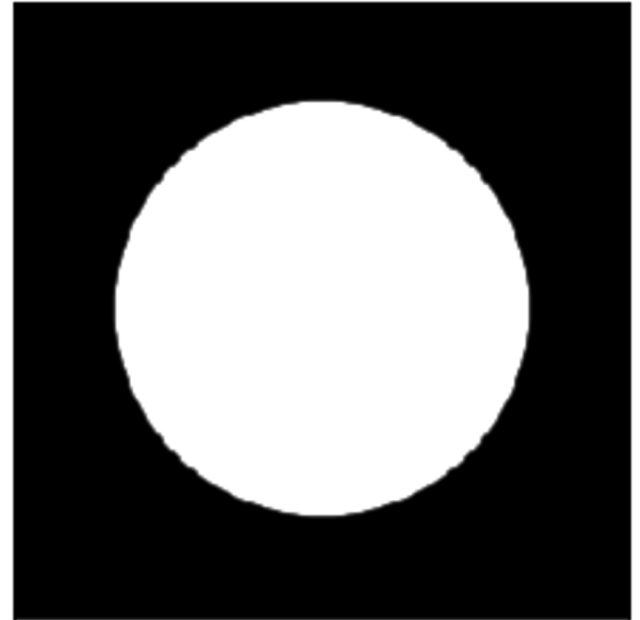
Focal Plane



$$\leftarrow \mathcal{F}(f(x,y))$$

$$\mathcal{F}^{-1}(F(\theta_x, \theta_y)) \rightarrow$$

Pupil Plane

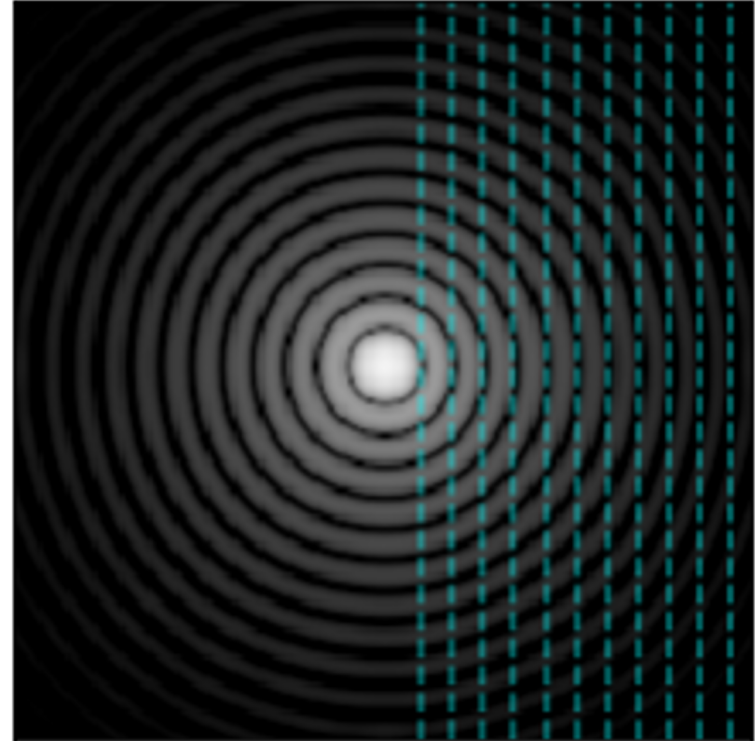


# Natural Units Intuitively Describe the Focal Plane

Airy Rings at  $1\lambda/D$  intervals

Diffraction Limited Resolution  
of a Telescope System:

$$\theta = \lambda/D$$



# Coronagraphy Basics

# Starlight Is Suppressed by a Coronagraph

Star + planet to observe



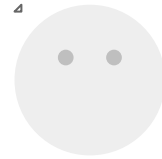
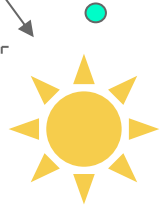
You, hoping for a



planet

# Starlight Is Suppressed by a Coronagraph

Star + planet to observe

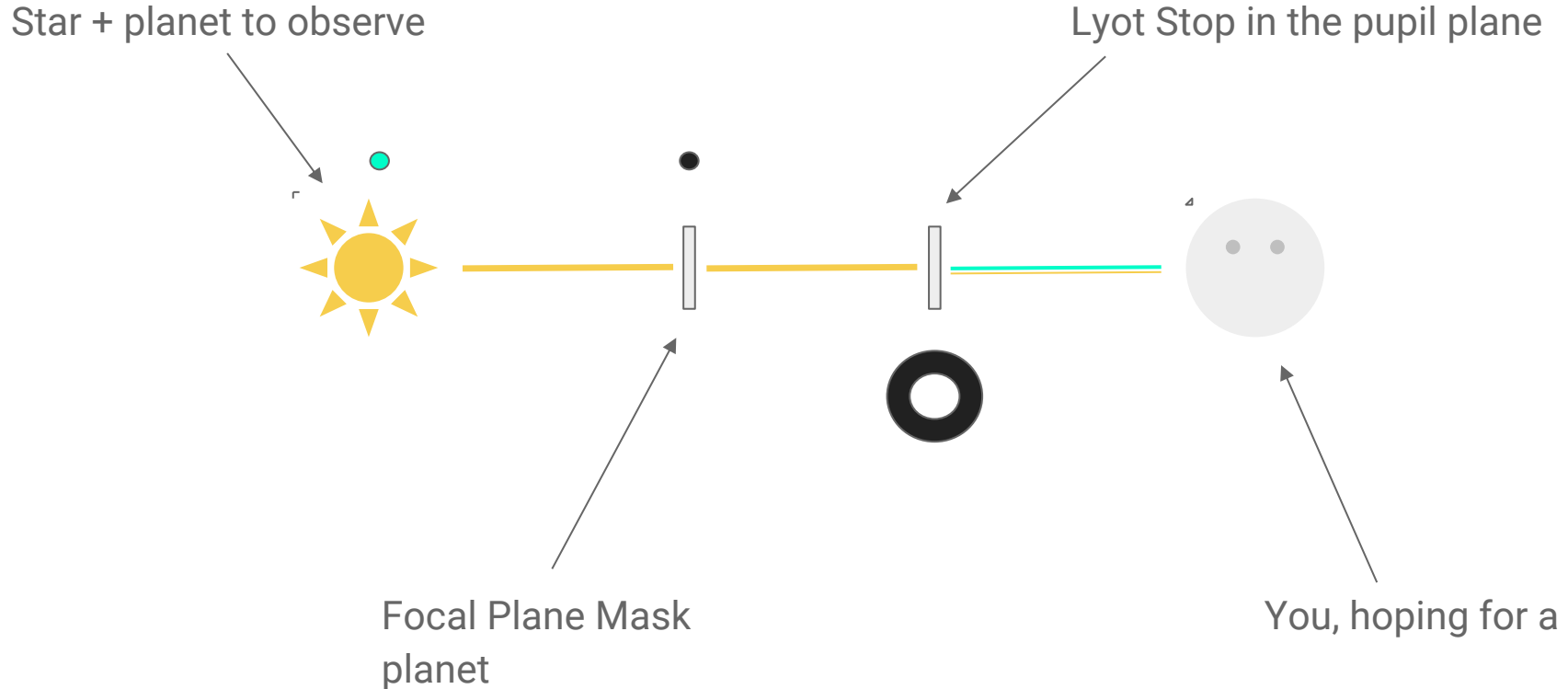


Focal Plane Mask  
planet

You, hoping for a

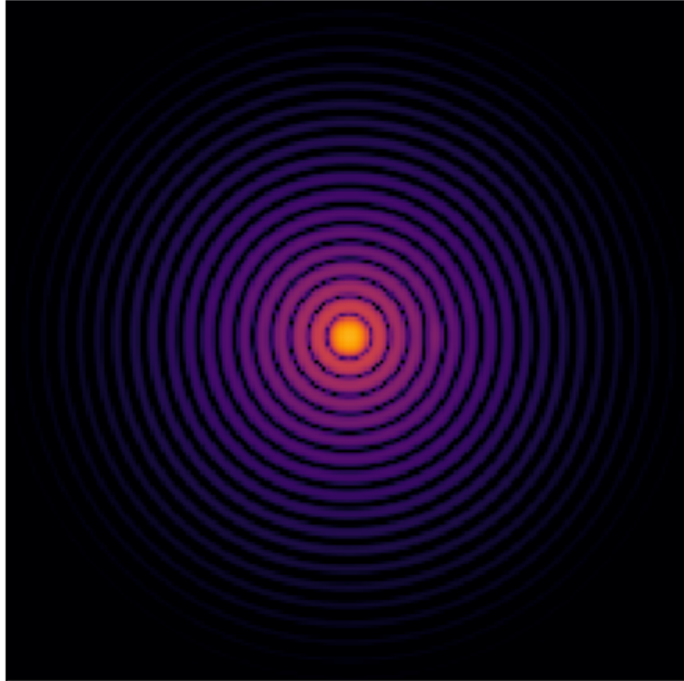


# Starlight Is Suppressed by a Coronagraph

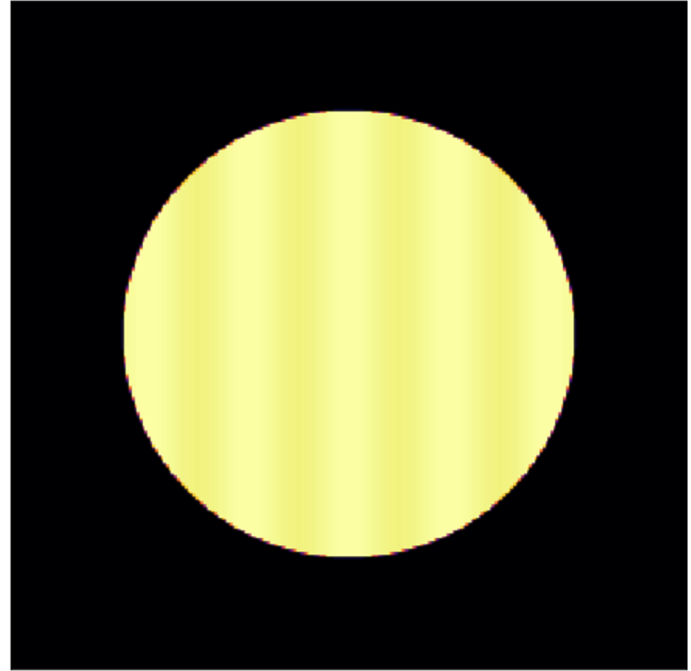


# Only Starlight is Visible, Despite an Injected Companion

Focal Plane

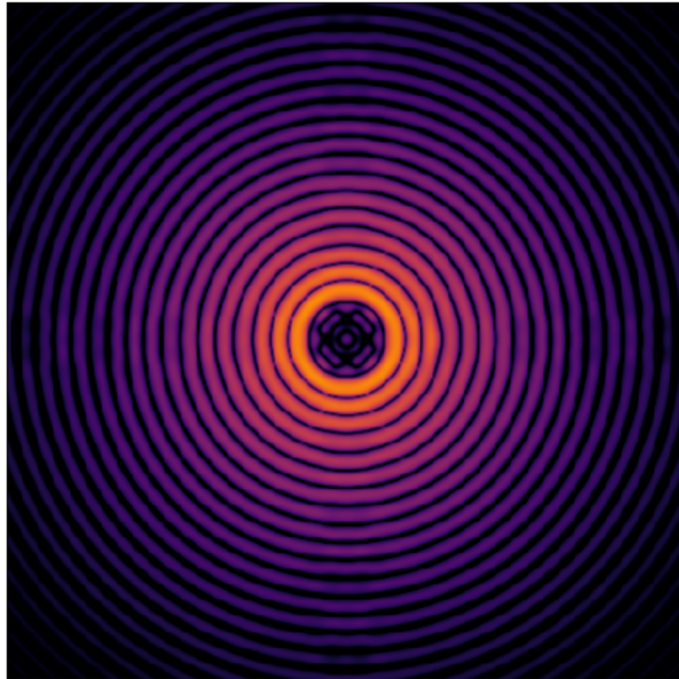


Pupil Plane



# Starlight is Diffracted to the Edge of the Aperture

Focal Plane

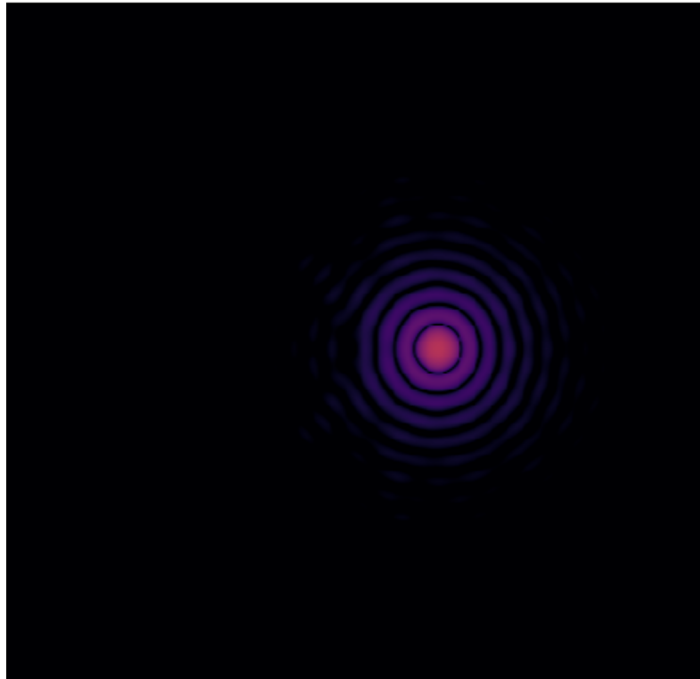


Pupil Plane

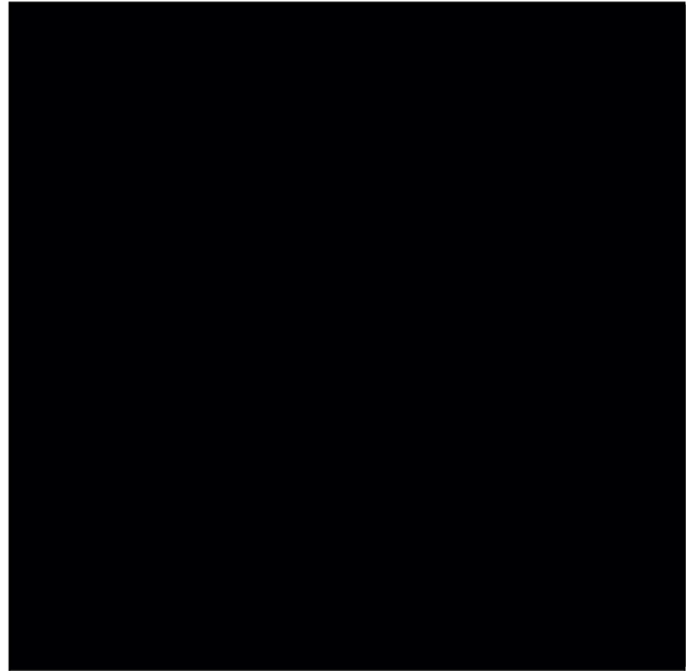


# The Planet Appears!

Focal Plane

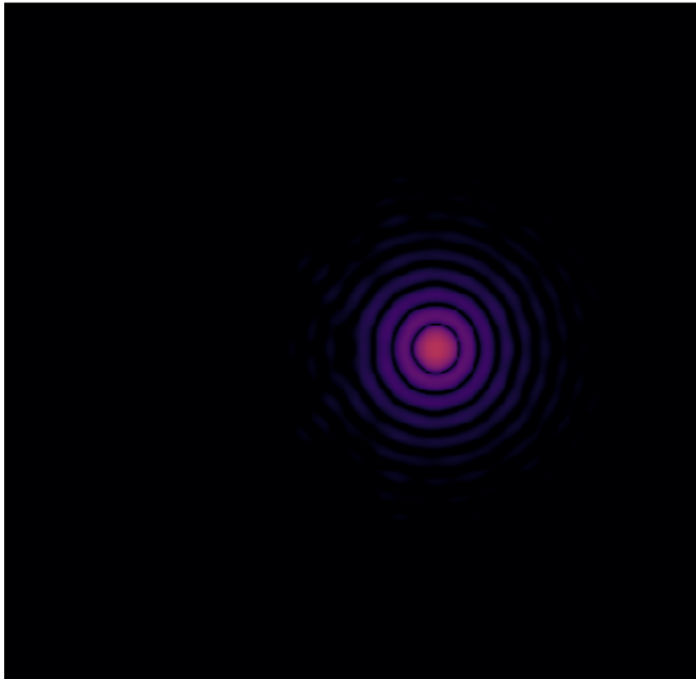


Pupil Plane

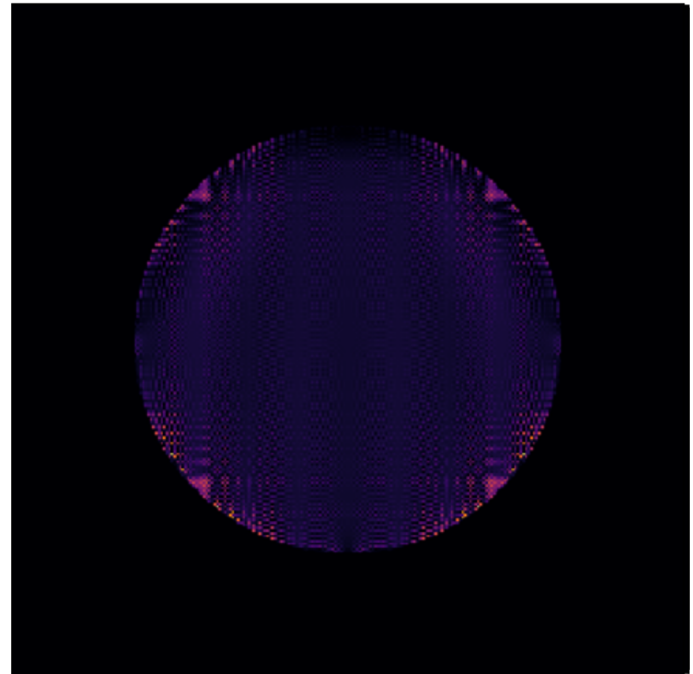


# The Planet Appears!

Focal Plane

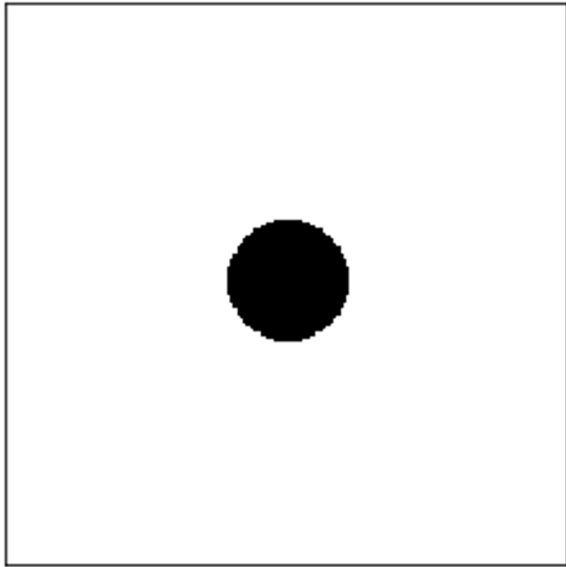


Pupil Plane

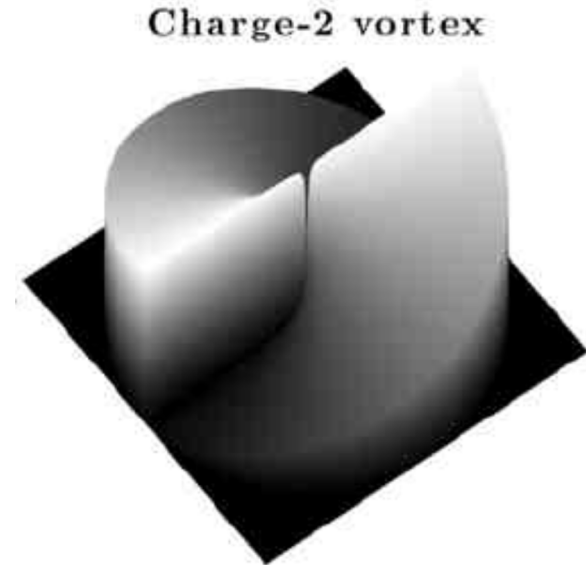


# Focal Plane Masks Vary for Classical Lyot and Vortex Coronagraphy

## Classic Focal Plane Mask



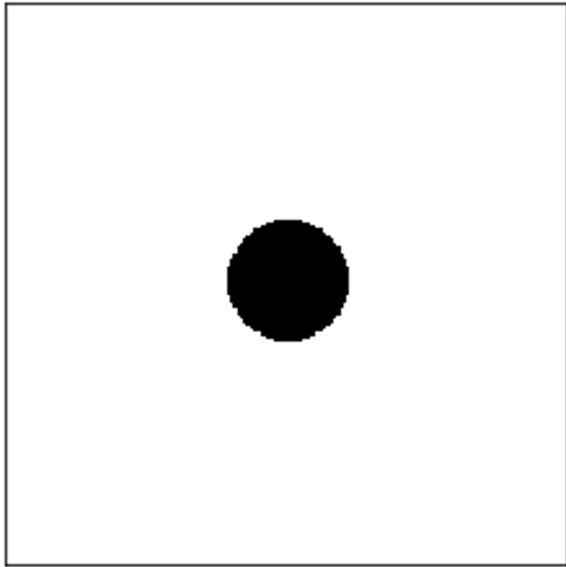
## Vortex Focal Plane Mask



(Delacroix, 2014)

# Focal Plane Masks Vary for Classical Lyot and Vortex Coronagraphy

## Classic Focal Plane Mask



## Vortex Focal Plane Mask



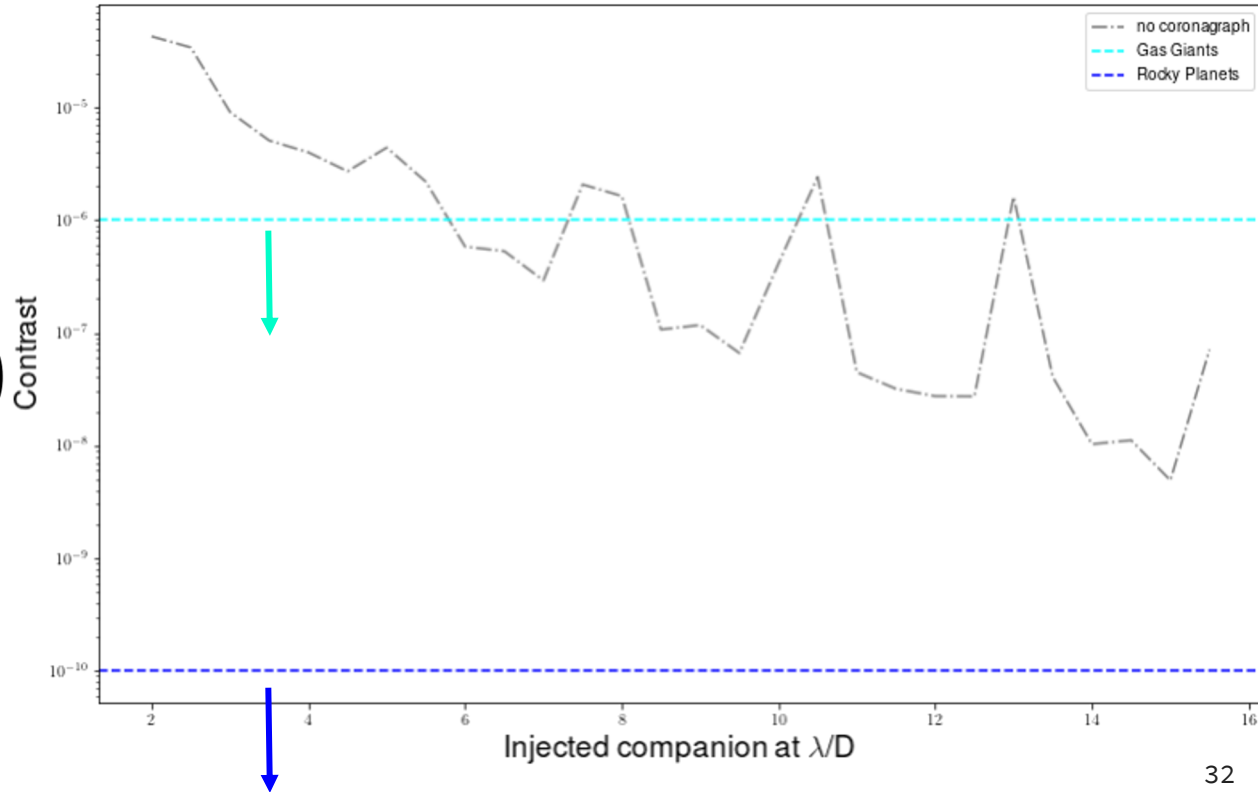
(MarbulaOne, 2020)

# Contrast Curves Across the Image Plane are a Coronagraph Performance Metric

Contrast normalized by throughput

- Jensen-Clem 2018

$$\text{contrast} = \left( \frac{\text{factor} \times \text{noise}}{\text{stellar aperture photometry}} \right) \times \left( \frac{1}{\text{throughput}} \right)$$





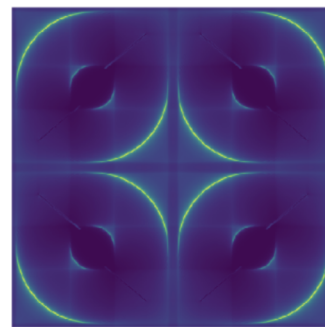
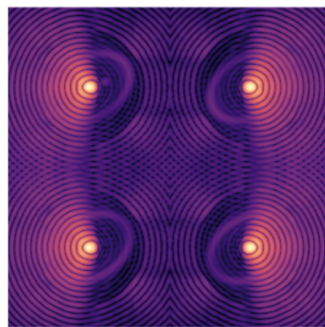
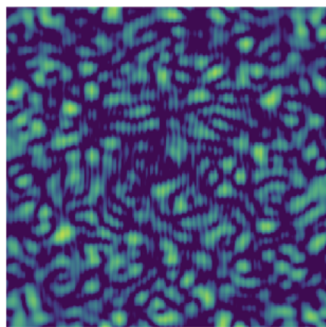
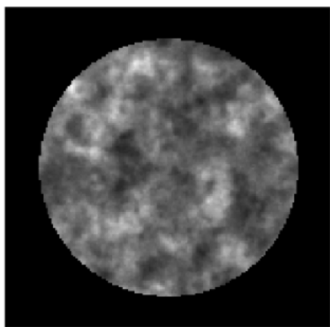
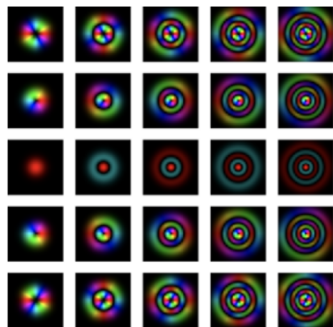
# Simulating a Coronagraph with HCIPy

# High Contrast Imaging for Python

[github.com/eorpor/hcipy](https://github.com/eorpor/hcipy)

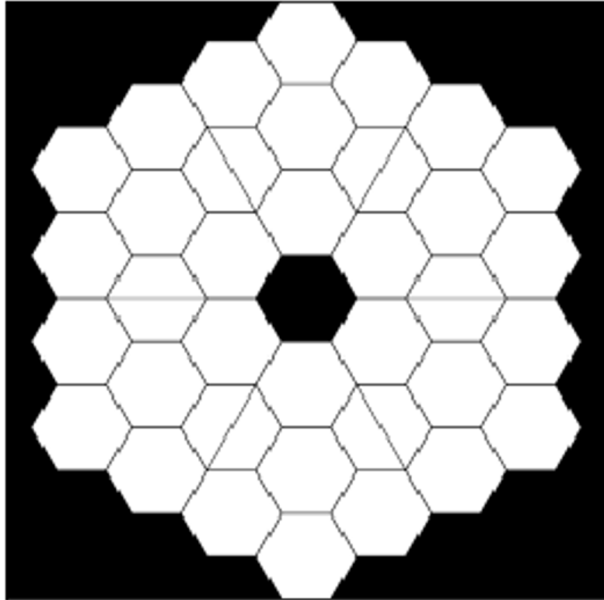


## HCIPy: High Contrast Imaging for Python

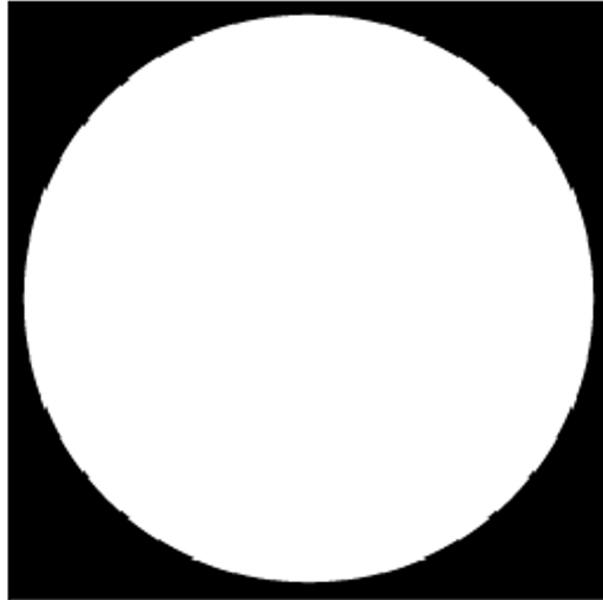


# Our Simulations Use a Circular Aperture

Keck Aperture



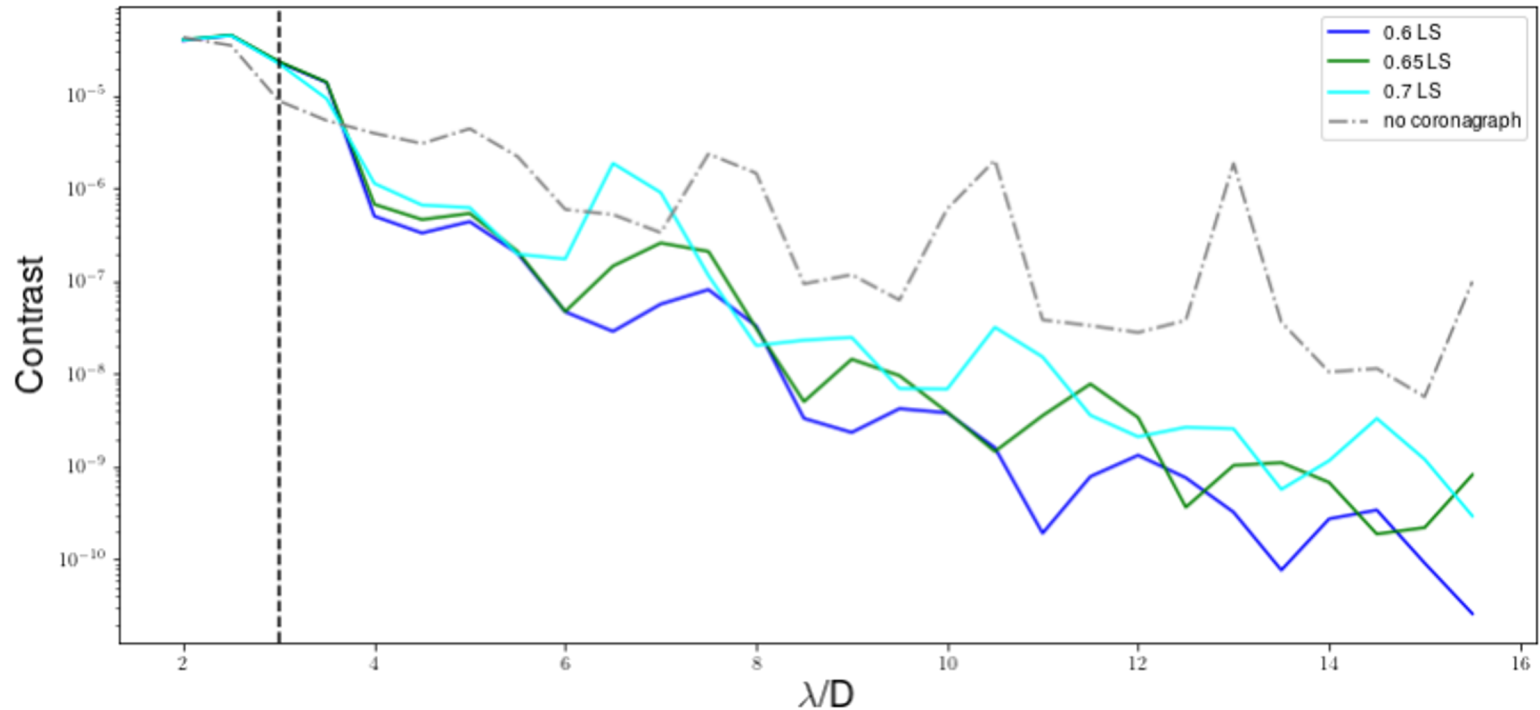
Circular Aperture



# Designing a B- Lyot Coronagraph

$$D_{LS}/D_{\text{aperture}} < 1 - 1/D_{FPM} : D_{FPM} = 3 \Rightarrow D_{LS} < 0.67$$

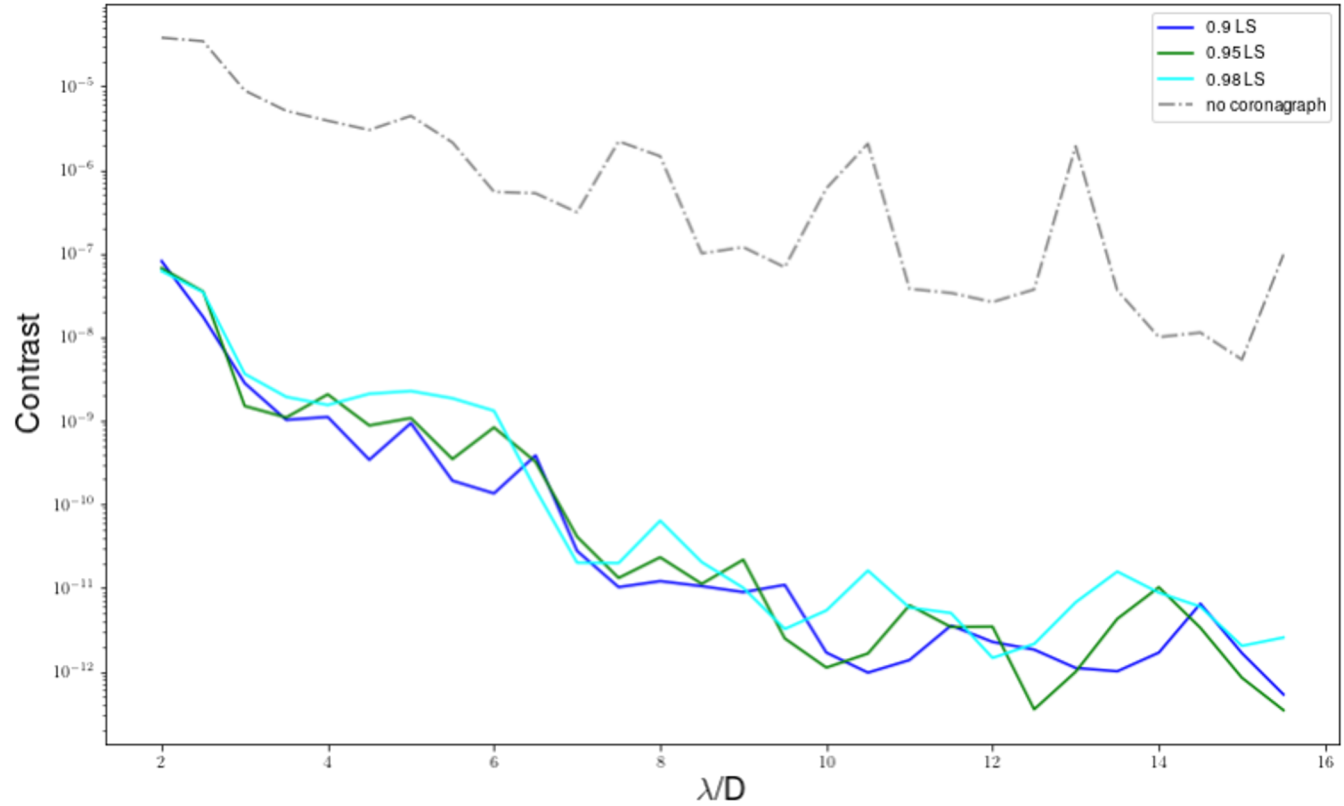
- Sivaramakrishnan, 2001



# Designing an A+ Vortex Coronagraph

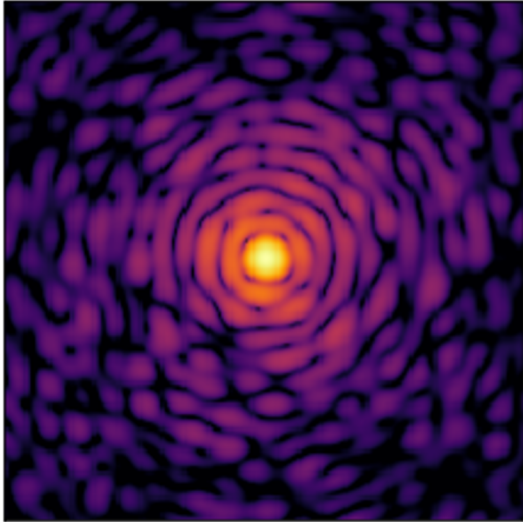
We expect to see no contrast difference past a certain threshold ( $\sim 0.98$ ).

- Ruane, 2018

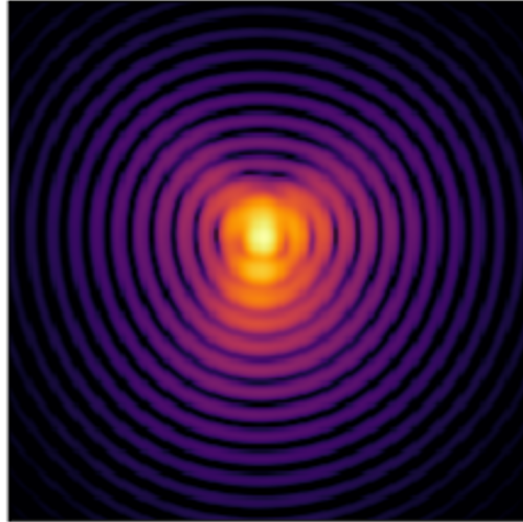


# Wavefront Error from Imperfect Optics Creates Speckles

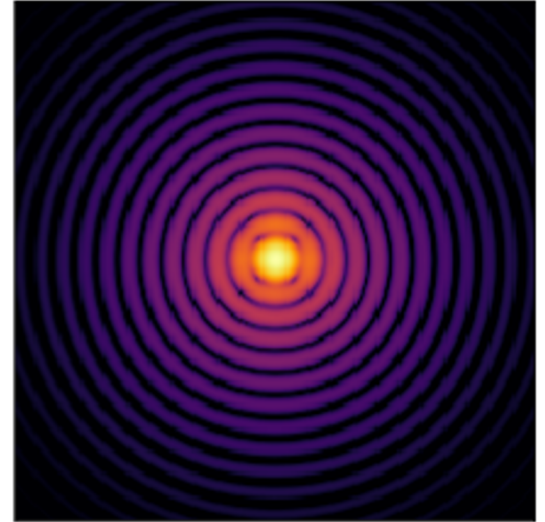
High Order WFE



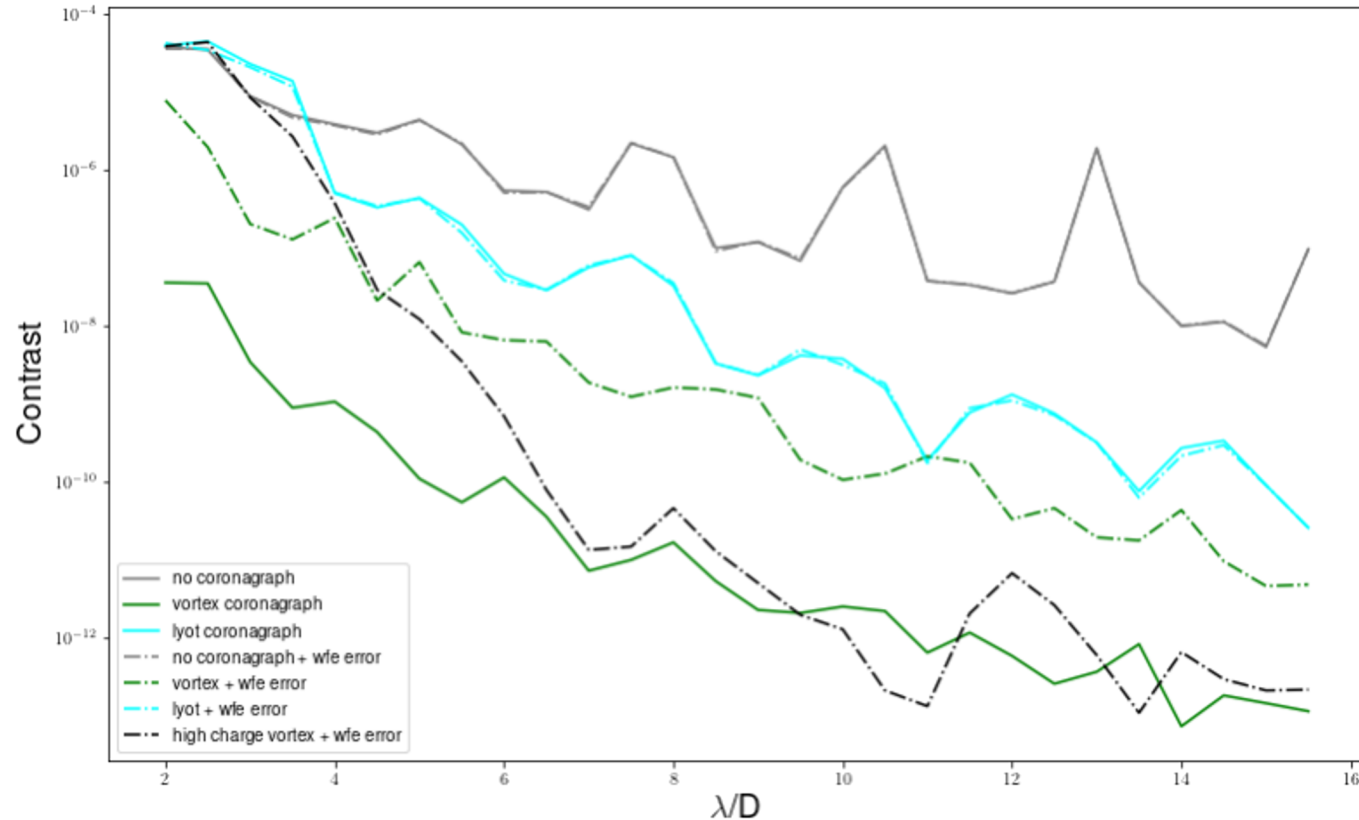
Low Order WFE



Realistic WFE

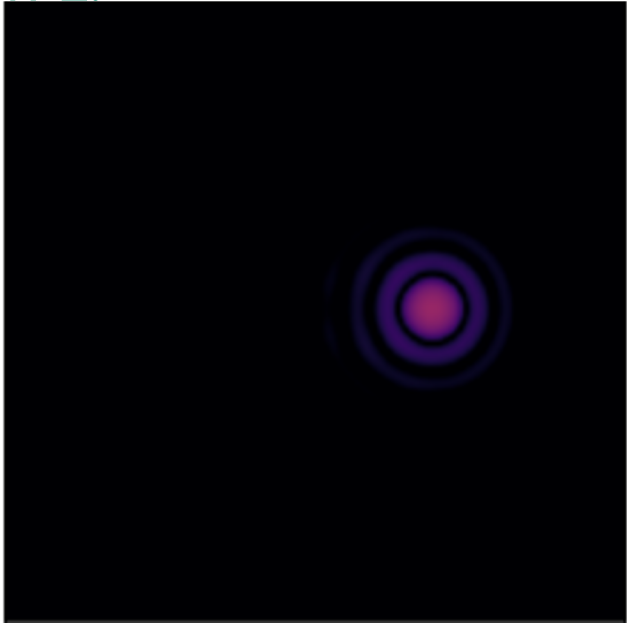


# Wavefront Error Impacts Contrast, Especially for the (low charge) Vortex

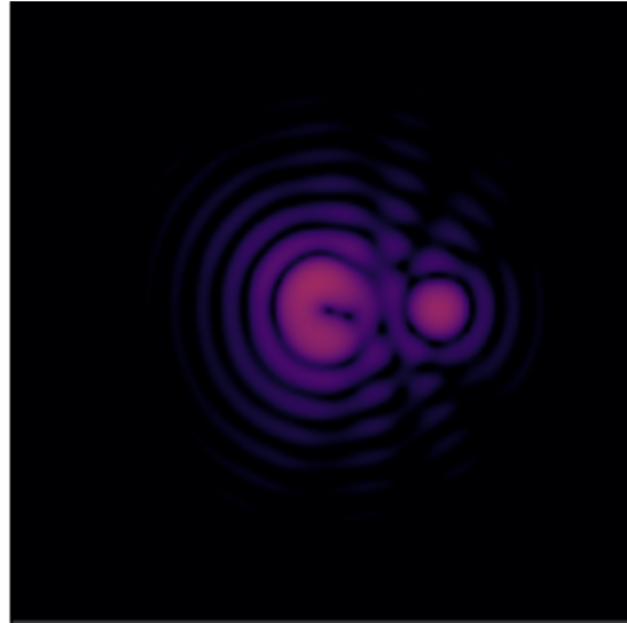


# Wavefront Error from Impacts the Final Coronagraphic Image

Perfect vortex:  
WFE:

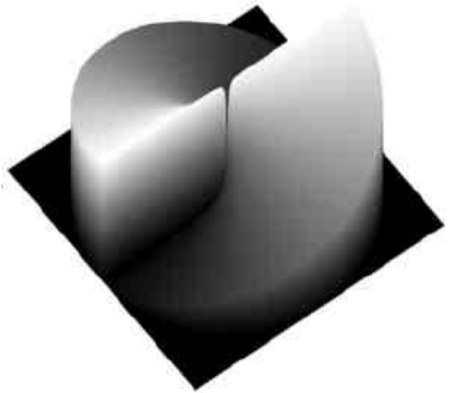
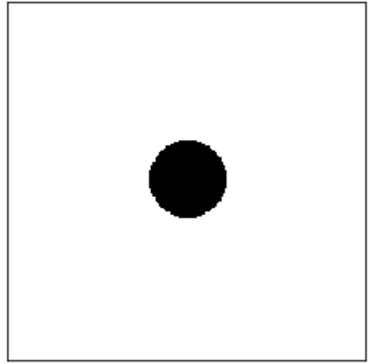


Vortex with

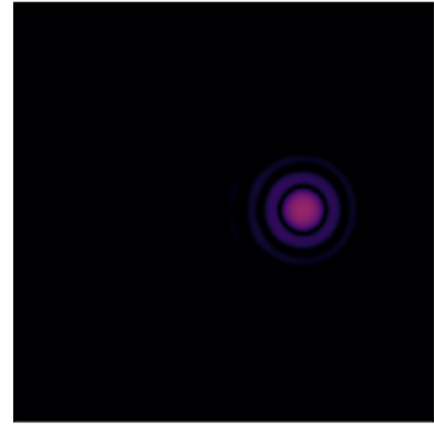




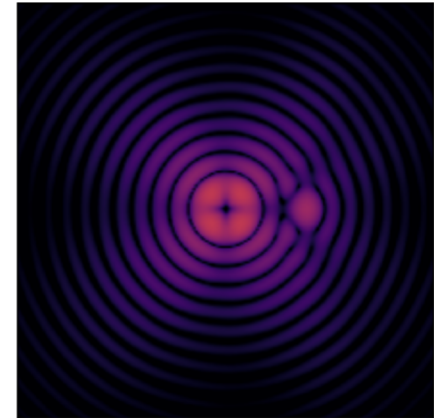
# Typical Vortex Masks are Imperfect



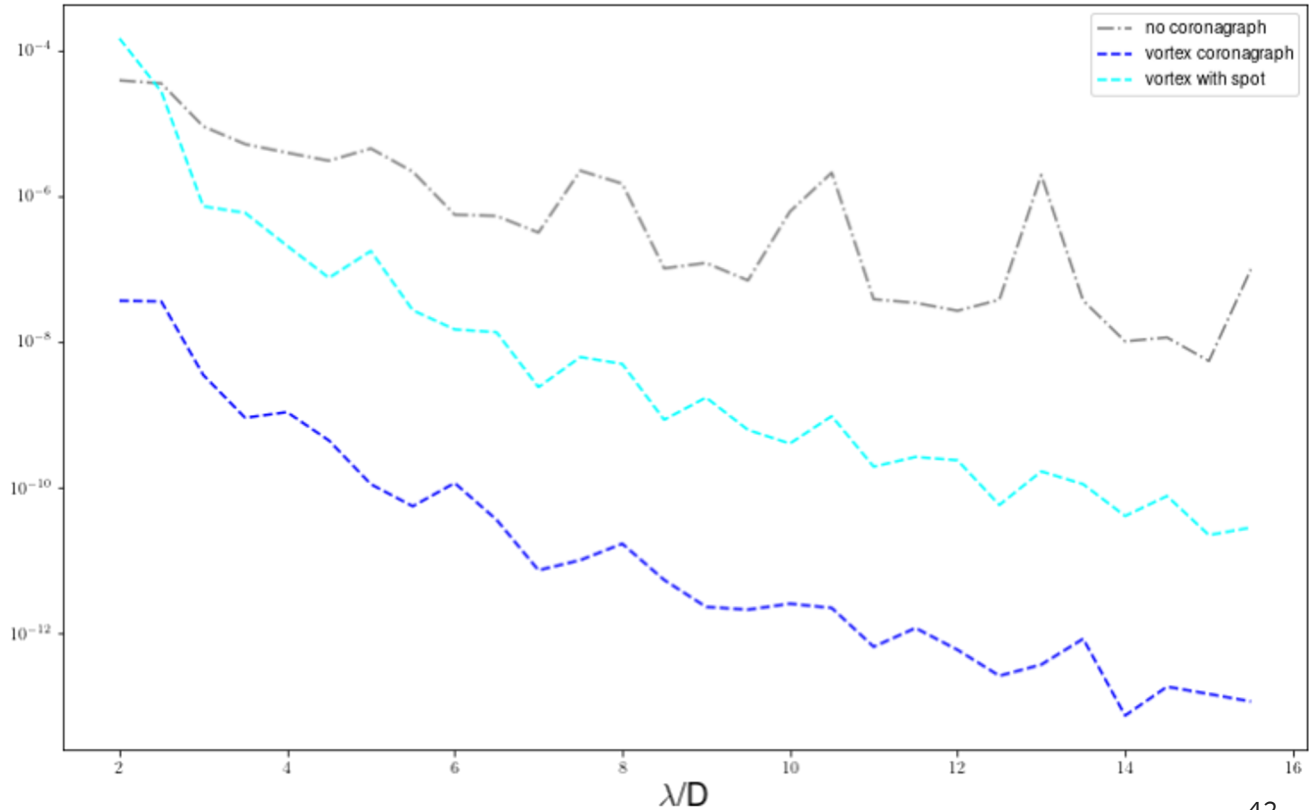
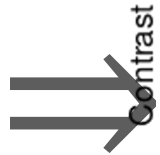
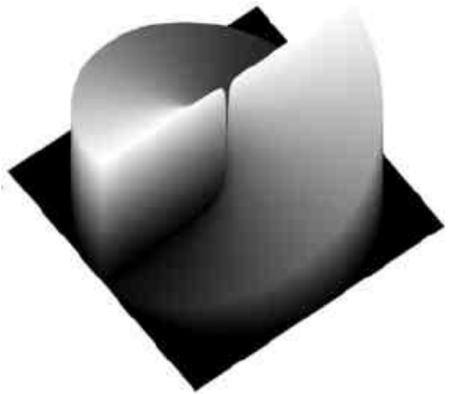
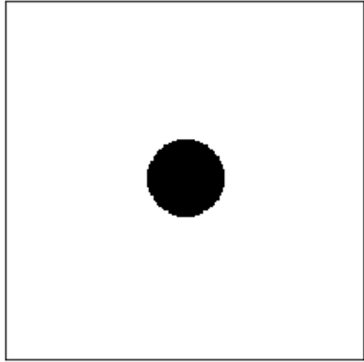
perfect vortex:



vortex + spot:



# Imperfect Vortex Masks Impact Contrast

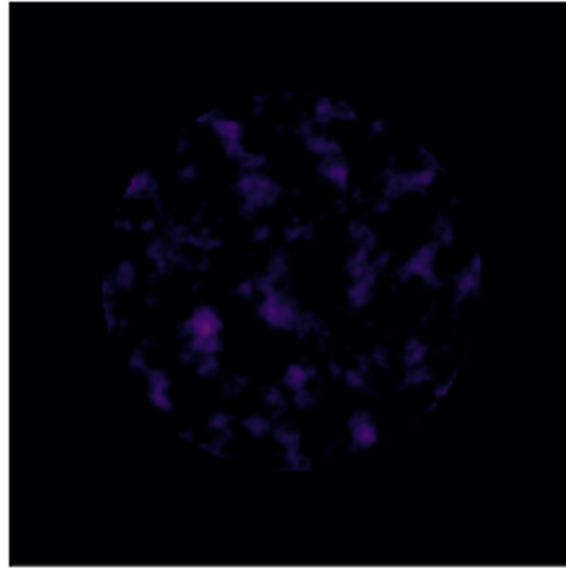
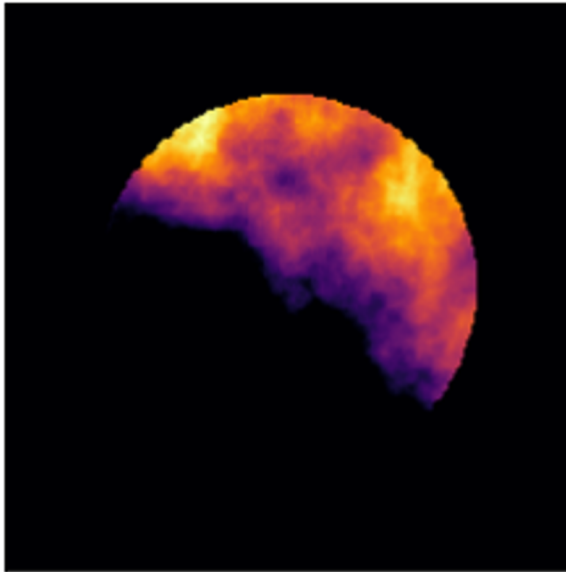


# Atmosphere Imparts Phase Errors But Can be Corrected with Deformable Mirrors

No Correction

50 Actuators

Kilo DM

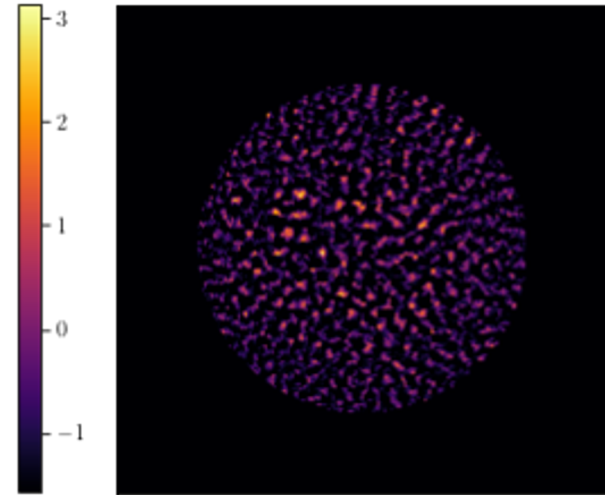
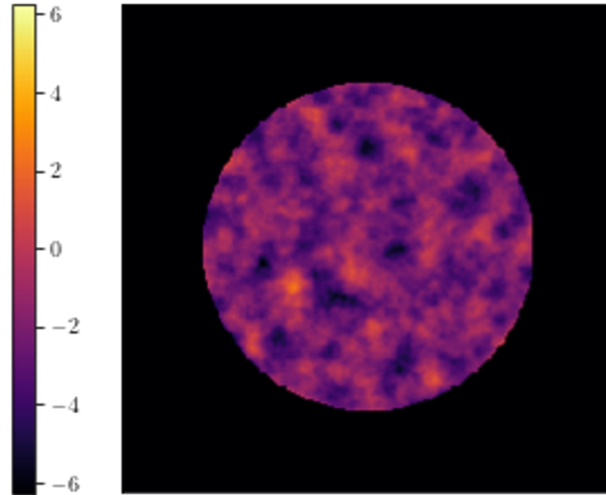
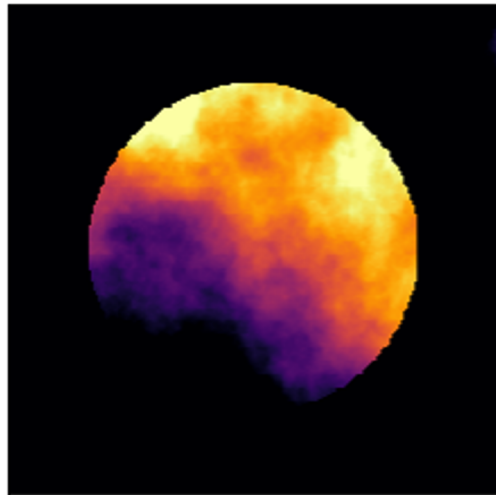


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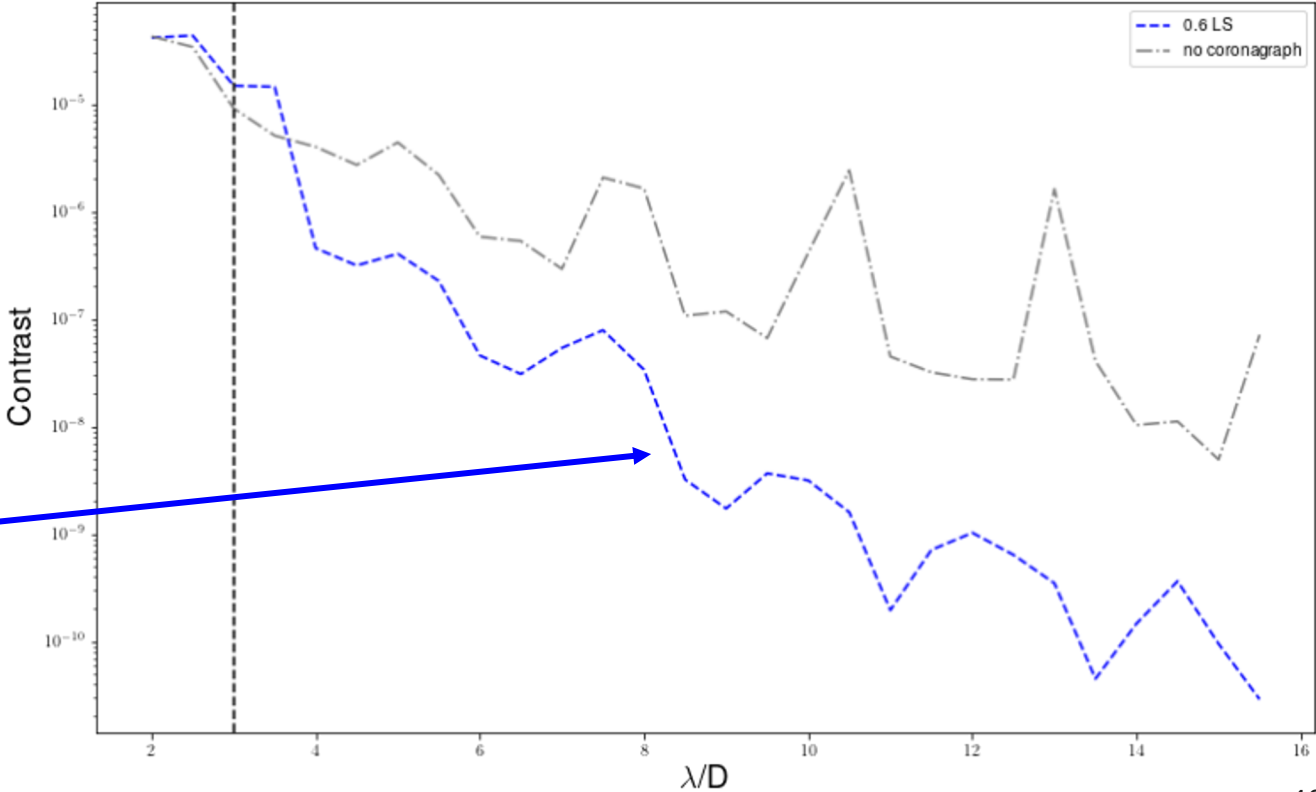
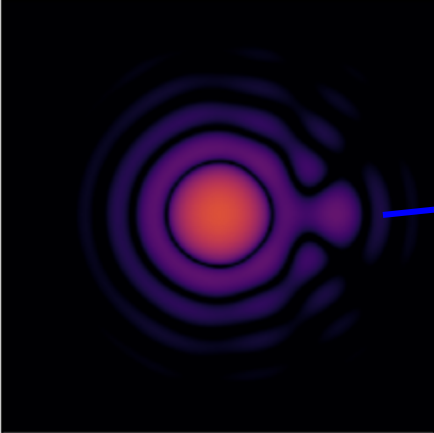


# Current State of the Design

# Large FPM Allows for Larg(er) Lyot Stop

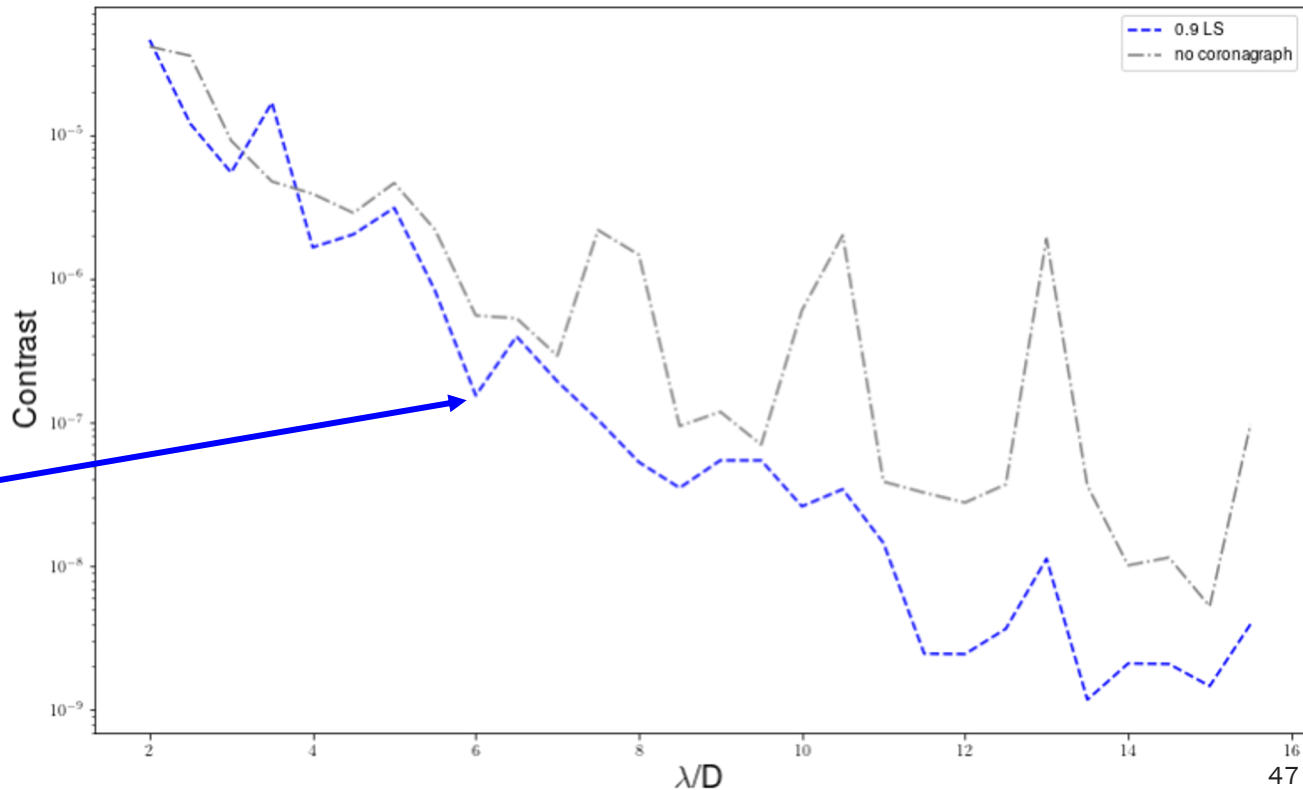
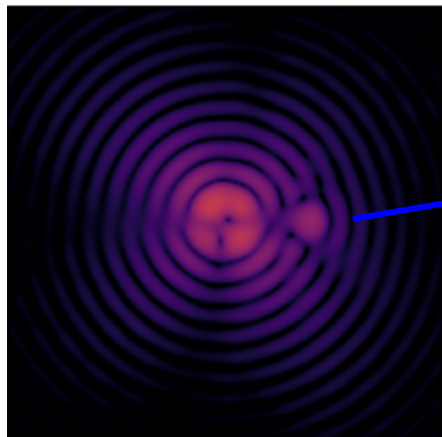
FPM:  $3\lambda/D$

LS: 0.6

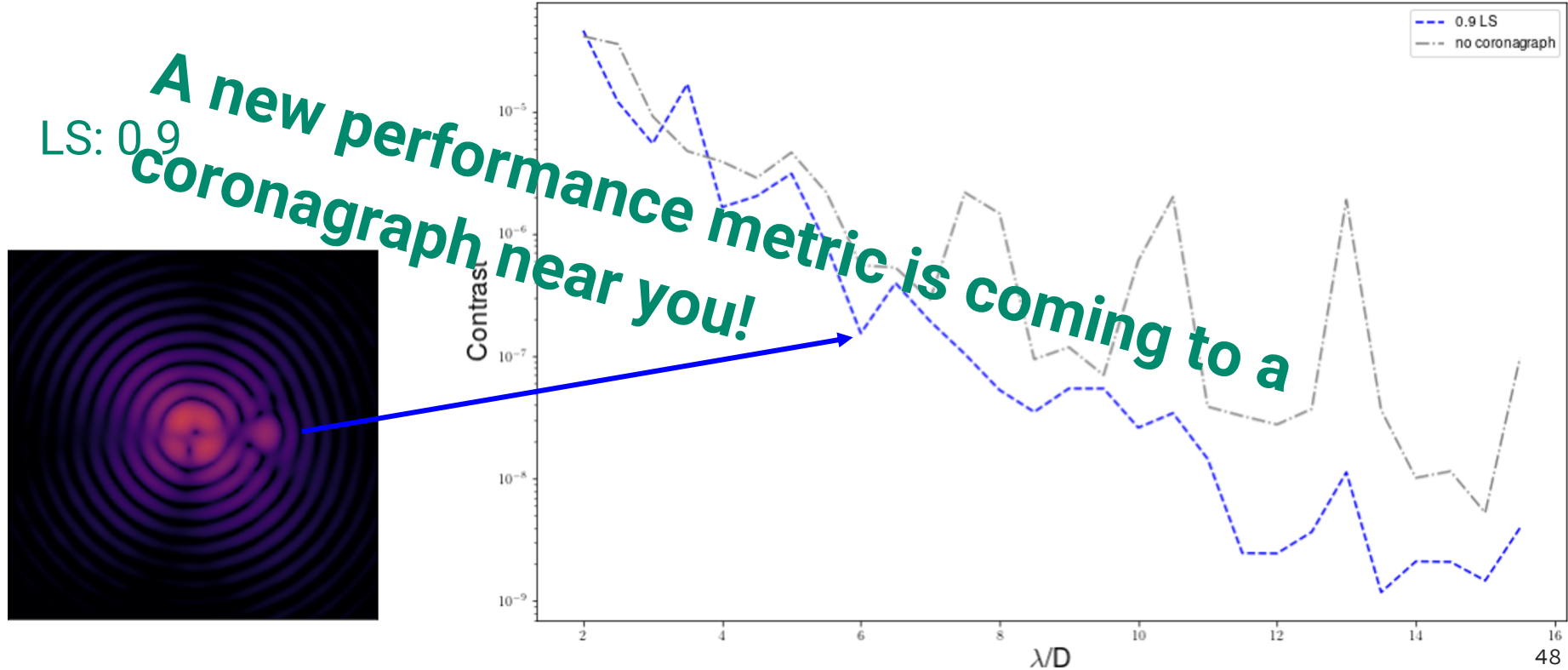


# Optimal Vortex Will be Difficult to Integrate and Align: Lyot Stop 0.98 $\rightarrow$ 0.9

LS: 0.9



# Optimal Vortex Will be Difficult to Integrate and Align: Lyot Stop 0.98 $\rightarrow$ 0.9





# Next Steps

- Incorporate cost function to minimize exposure time as secondary metric
- Model and design additional elements, including a pupil mask optimized for Keck-like apertures and an apodizing phase pattern we could apply with Deformable Mirrors
- Simulate and design a preliminary Lyot Coronagraph for the Thirty Meter Telescope (TMT)
- Use our setup to compare predictive wavefront control methods side-by-side

# Acknowledgements

Many thanks to ExEP and the ExoExplorers program, the UCSC Lab for Adaptive Optics, Becky Jensen-Clem, Maaïke Van Kooten, Ben Gerard, Anand Sivaramakrishnan, Gary Ruane, Renate Kupke, and Daren Dillon, and the other grads at UCSC.

# In Conclusion

- Lyot Coronagraphs: stand up well to system errors but rule out close in companions.
- Vortex Coronagraphs: on paper offer stunning performance but don't hold up as well to practical systems.
- Other coronagraphic elements and algorithms will be vital for high contrast, and practical integration will bring other affects we haven't yet thought to model.
- Feel free to contact me with any further questions (or heckling):



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[@121gigajules](https://twitter.com/121gigajules)