

# Welcome to the Workshop on Advanced Wavefront Sensing for Coronagraphs!

May 1, 2020

**Brendan Crill** 

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#### **Today's Workshop**

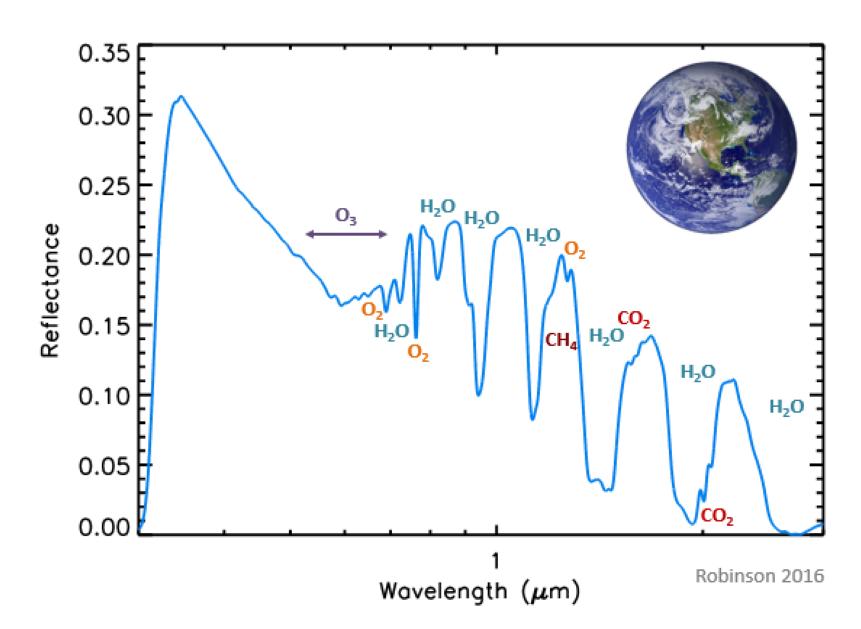


#### Agenda

- 09:00 Pacific time 10 min Brendan Crill (NASA ExEP) Wavefront Sensing and NASA's Goals for Exoplanet Direct Imaging
- 09:10 20 min Laurent Pueyo (STScI) Wavefront Sensing in a space-based coronagraph instrument
- 09:30 20 min Mamadou N'Diaye (Cote D'Azur) ZELDA results on VLT/SPHERE
- 09:50 20 min John Steeves (JPL) picometer sensitivity demo
- 10:10 20 min Garreth Ruane (JPL) Decadal Survey Testbed demonstration of a Zernike WFS
- 10:30 20 min David Doelman (Leiden Observatory) vector Zernike WFS
- 10:50 10 min break
- 11:00 20 min Kent Wallace (JPL) Vector Zernike WFS progress at JPL: liquid crystal and metasurface devices and applications
- 11:20 20 min Emiel Por (Leiden Observatory) Simultaneous wavefront sensing demonstration at Paris Observatory
- 11:40 20 min Jeff Jewell (JPL) New concepts in wavefront sensing for high contrast imaging
- 12:00 50 min Q&A, open discussion
- 12:50 10 minute wrapup: actions, next steps

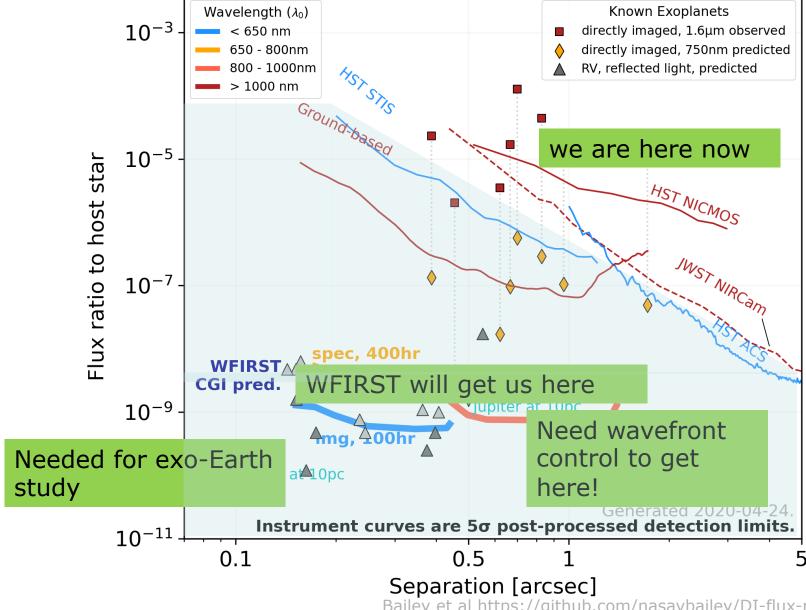
### Searching for Life in the Universe – reflected light spectroscopy of terrestrial exoplanets





#### **Extreme Starlight Suppression with a Coronagraph** requires Wavefront Control





#### **NASA Exoplanet Exploration Program's Focus on Exoplanet Direct Imaging from Space**



#### 2010 Decadal Survey

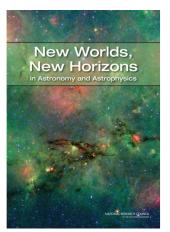


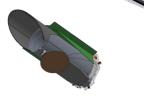
TABLE ES.4 Space: Recommended Activities—Medium-Scale (Priority Order)		
Recommendation	Science	Appraisal of Costs <sup>a</sup>
1. New Worlds Technology Development Program	Preparation for a planet-imaging mission beyond 2020, including precursor science activities	\$100M to \$200M
2. Inflation Probe Technology Development Program	Cosmic microwave background (CMB)/ inflation technology development and preparation for a possible mission beyond 2020	\$60M to \$200M

#### 2018 Exoplanet Science Strategy

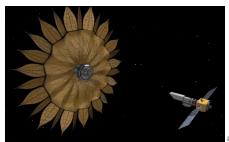


Recommendation: NASA should lead a large strategic direct imaging mission capable of measuring the reflected-light spectra of temperate terrestrial planets orbiting Sun-like stars.









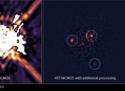
#### V-NIR Coronagraph/Telescope Technology Gaps

**Contrast** (ExEP Technology Gap List)

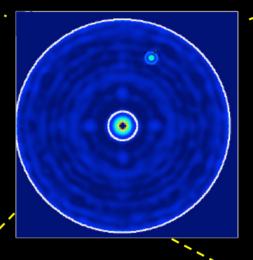




CG-3: Deformable Mirrors



CG-4: Data Post-Processing

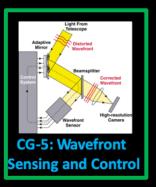


Angular Resolution



CG-1: Segmented Mirrors

#### **Contrast Stability**

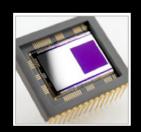


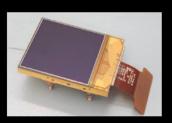




CG-7: Telescope Vibration
Sensing and Control or Reduction

#### **Detection Sensitivity**





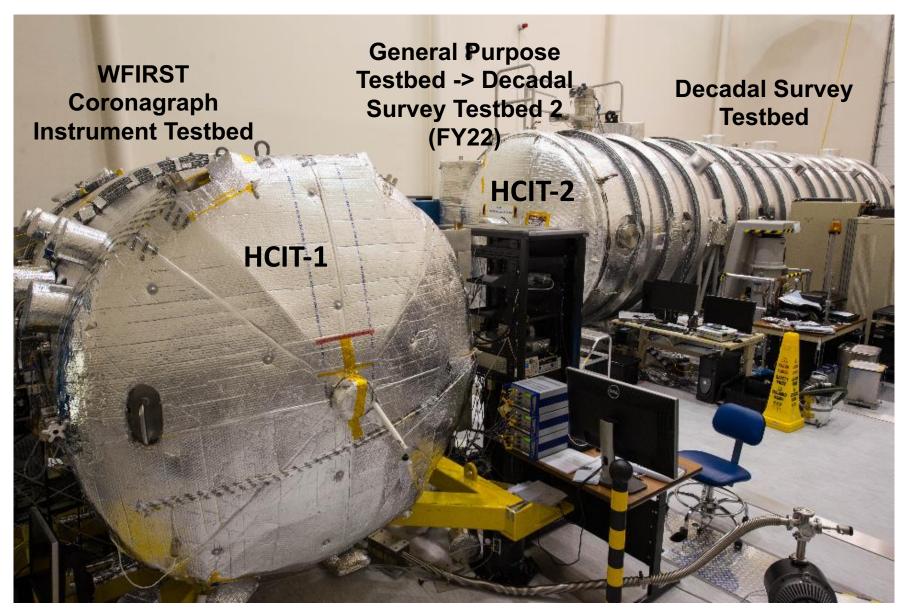
Ultra-low Noise Visible (CG-8) and Infrared (CG-9) Detectors

**CG-6: Mirror Segment Phasing** 

https://exoplanets.nasa.gov/exep/technology/gap-lists/

#### **ExEP's High Contrast Imaging Testbed**





#### **Decadal Survey Testbed**



- Available to investigations funded by Strategic Astrophysics Technology grants
- Commissioned using a Lyot coronagraph  $3.8x10^{-10}$  contrast, working angles 3-9  $\lambda/D$ , 9% bandwidth (Seo et al 2019)
- A new Zernike Wavefront Sensor has been installed and demonstrated (Garreth's talk)
- Decadal Survey Testbed roadmap here:
   <a href="https://exoplanets.nasa.gov/internal\_resources/1170/">https://exoplanets.nasa.gov/internal\_resources/1170/</a>
- Aim is to commission a testbed capable of demonstrating < 10<sup>-9</sup> contrast with a simulated dynamic environment

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- Type questions into WebEx chat (I'll monitor that)
- Remember to keep your mic muted
- Keep video feed off (except for the presenter)



## 10-MINUTE BREAK RESTARTING AT 11:10 PACIFIC TIME

#### **Q&A / Discussion**



- If you ask a question, please identify yourself and state to whom you are directing your question
- Remember to go on mute if you're not talking