



Exoplanet Exploration Program

Starshade Stray Light Assessment presented at Science and Industry Partnership F2F-2

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Introduction



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- **We desire ≥ 30 mag planet sensitivity at SNR ≥ 4 to require noise sources ≥ 31.5 mags**
 - 30 mag planet has 1×10^{-10} planet contrast relative to a magnitude 5 star
 - Brighter noise sources can be calibrated down to 31.5 mag residuals
- **Here we try to identify all starshade created stray light sources & assesses compliance**
 - Can be considered “foreground” stray light
 - Stuart Shaklan separately presents solar edge scatter
 - Glenn Sellar presents secondary reflections (telescope facing surface with view to sun face surface)
 - Compliance presented for WFIRST Rendezvous but Habex delta’s are also discussed
- **Astronomical background sources are thought to set the noise floor at ~ 31.5 mags**
 - Can the TSWG confirm this or recommend another limit ?



Foreground and Background Stray Light Sources



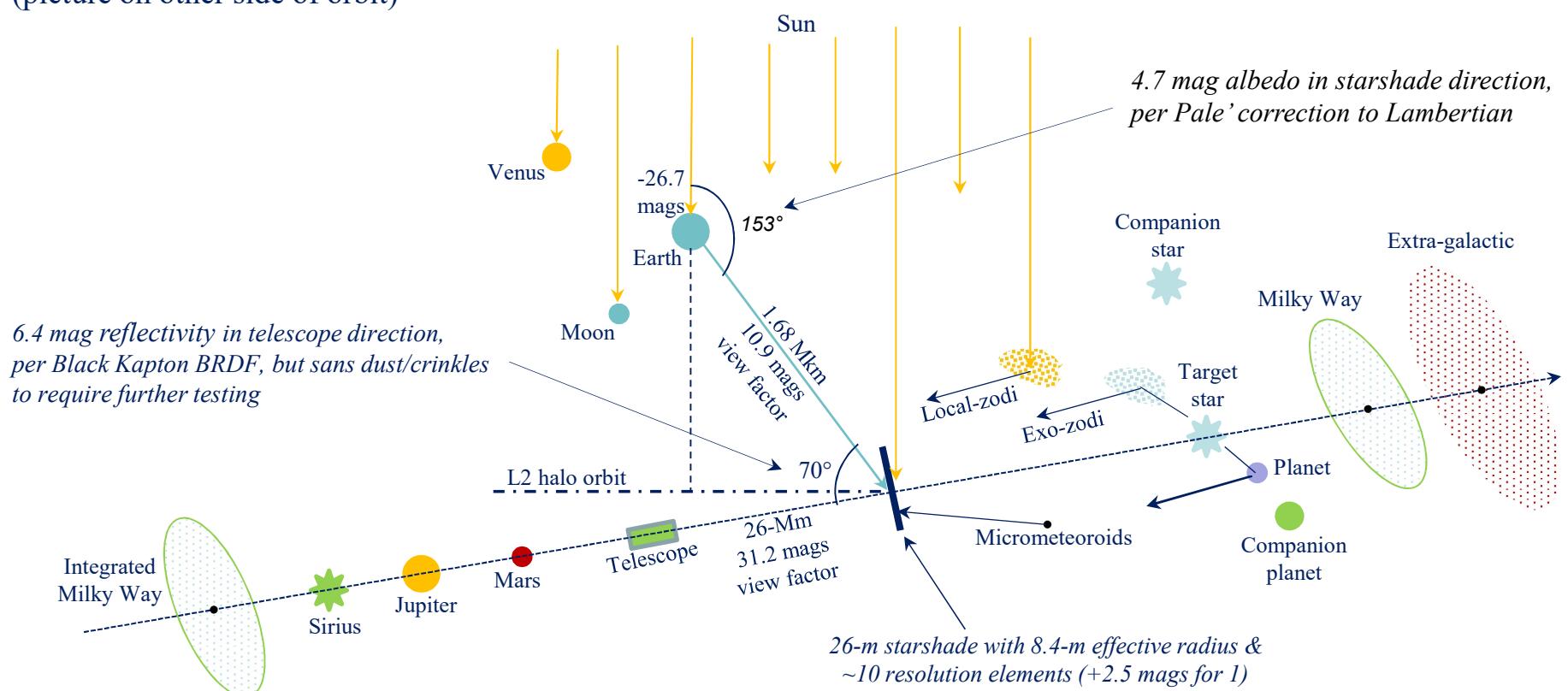
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Starshade reflected bright-bodies
Off-axis Earth, Moon and Venus
On-axis Jupiter, Mars and Sirius
On-axis distributed Milky Way
Secondary solar reflections

Starshade solar leakage
Optical edge scatter
Direct material transmission
Optical shield close-out at petal-disk I/F
Transmission thru micrometeoroid holes

Astrophysical Background
Galactic & Extra-galactic
Exo-zodiacal light
Local-zodiacal light
Companion stars & exoplanets

Telescope straylight
Bright bodies direct to telescope
(picture on other side of orbit)





Summary of Results



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**2 sources
detailed
here**

Bright-body case	Max expected level (mags per telescope res element)	Comments
Off-axis Earth	29	Requires calibration to 25% accuracy plus darker starshade &/or target pointing constraints
Off-axis Moon	≥ 31.6	Gives ≥ 30 mag planet sensitivity without calibration
Off-axis Venus	34.6	
On-axis Jupiter	32.8	
On-axis Mars	32.8	
On-axis Sirius	34.3	
On-axis integrated Milky Way	32.7	
Sun thru micrometeoroid holes	37	Large prelim. margin covers large uncertainty
Material opacity	$>> 32$	Black Kapton is highly opaque but rapid 2010 measurements are not documented & retest is desired (Princeton testbed)
Sun thru optical shield close-outs	TBD	Will design seals to required level
Optical Edge Scatter	See Shaklan's presentation	
Secondary Reflections	See Sellar's presentation	



Reflected Earth Light



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- Max expected reflected Earth light at the telescope is 29 mags per resolution element
 - At extreme of orbit and target pointing
 - See table below (all numbers from figure above)
- Calibration accuracy is probably limited by time varying Earth albedo to ~25%, which yields a planet sensitivity of only 29 mags (see table below) for a SNR of 4
- A darker starshade (AR coating) combined with minor pointing restrictions should achieve the 30 mag planet sensitivity goal

Solar incidence on Earth	-26.7	
Earth albedo at 153 deg	4.7	In direction of starshade
Earth-starshade view factor	10.9	
Starshade reflectivity at 70 deg	6.4	In direction of telescope
1 of ~10 resolution elements	2.5	
Starshade-telescope view factor	31.2	
Earth light at telescope	29.0	mags per resolution element

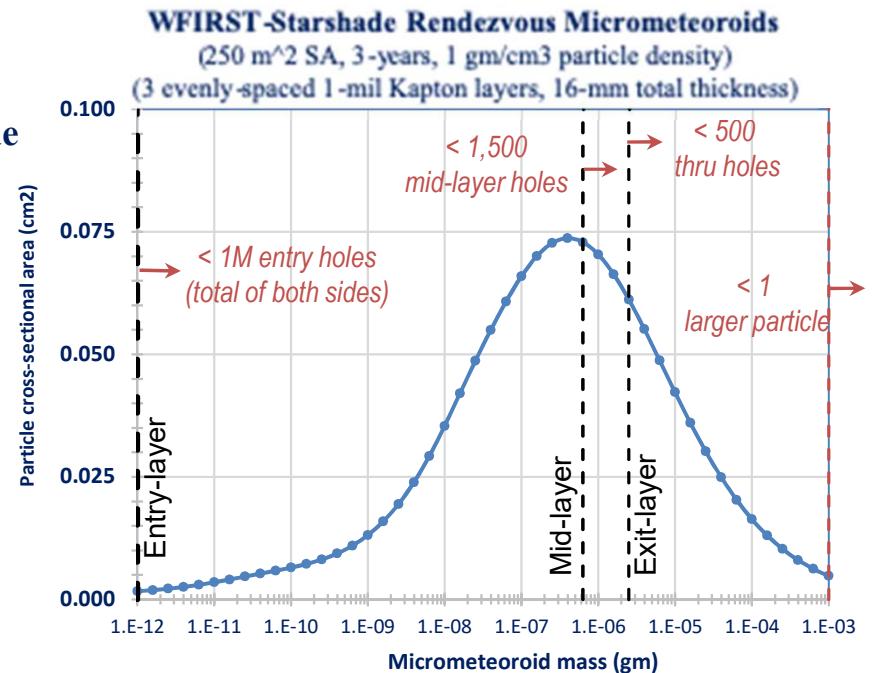


Sun thru micrometeoroid holes



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- Optical shield consists of 3 evenly spaced layers of 1-mil thick Kapton and total thickness of 16-mm
- Heaviest particle stopped at each layer per published formulae
 - NASA Std. PD-EC-1107 for single layer (entry holes)
 - SMAD for 2 spaced layers (3rd layer exit hole calc neglects mid-layer)
 - 50 km/s particle velocity (near max)
 - 2.5 gm/cm³ particle density (only occurs for smaller particles)
 - Experimental constants for Al shield scaled with density to Kapton (Kapton melts at much higher temps, so should be conservative)
- Particle flux per Grun model (same as JWST)
- Max expected hole area is ~100 ppm in each layer
 - Particle areas per 1 gm/cm³ density (only occurs for large aggregates)
 - Integrate area under curve to right for cum area of larger particles and multiply by hole expansion factors relative to particle diameters
 - 1.25X entry holes per observed craters
 - 20X mid-layer holes per semi-analogous simulations at much lower velocities (lateral gas cloud expansion is less at higher velocities)
 - 40X exit layer holes (2X spacing but neglects benefit of mid-layer)
- 37 mag per resolution element solar leakage at telescope
 - 1E-12 (30 mags) porosity based solar transmission (100 ppm)³
 - See table to right



Solar incidence on starshade	-26.7	
Starshade hole porosity	30	100 ppm ³ = 1E-12
1 of ~10 resolution elements	2.5	
Starshade-telescope view factor	31.2	
Earth light at telescope	37.0	mags per resolution element



Habex Comparison



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- All reflected bright body levels are ~0.7 mag dimmer for HabEx due to smaller telescope-starshade view factor (radius doubles, but separation distance triples) and fewer resolution elements
- Sun thru micrometeoroid holes is ~2 mags brighter due to larger surface area (900 vs. 250 m²) and longer mission duration (5 vs. 3-years) and fewer resolution elements
 - Still plenty of margin



Backup Slides



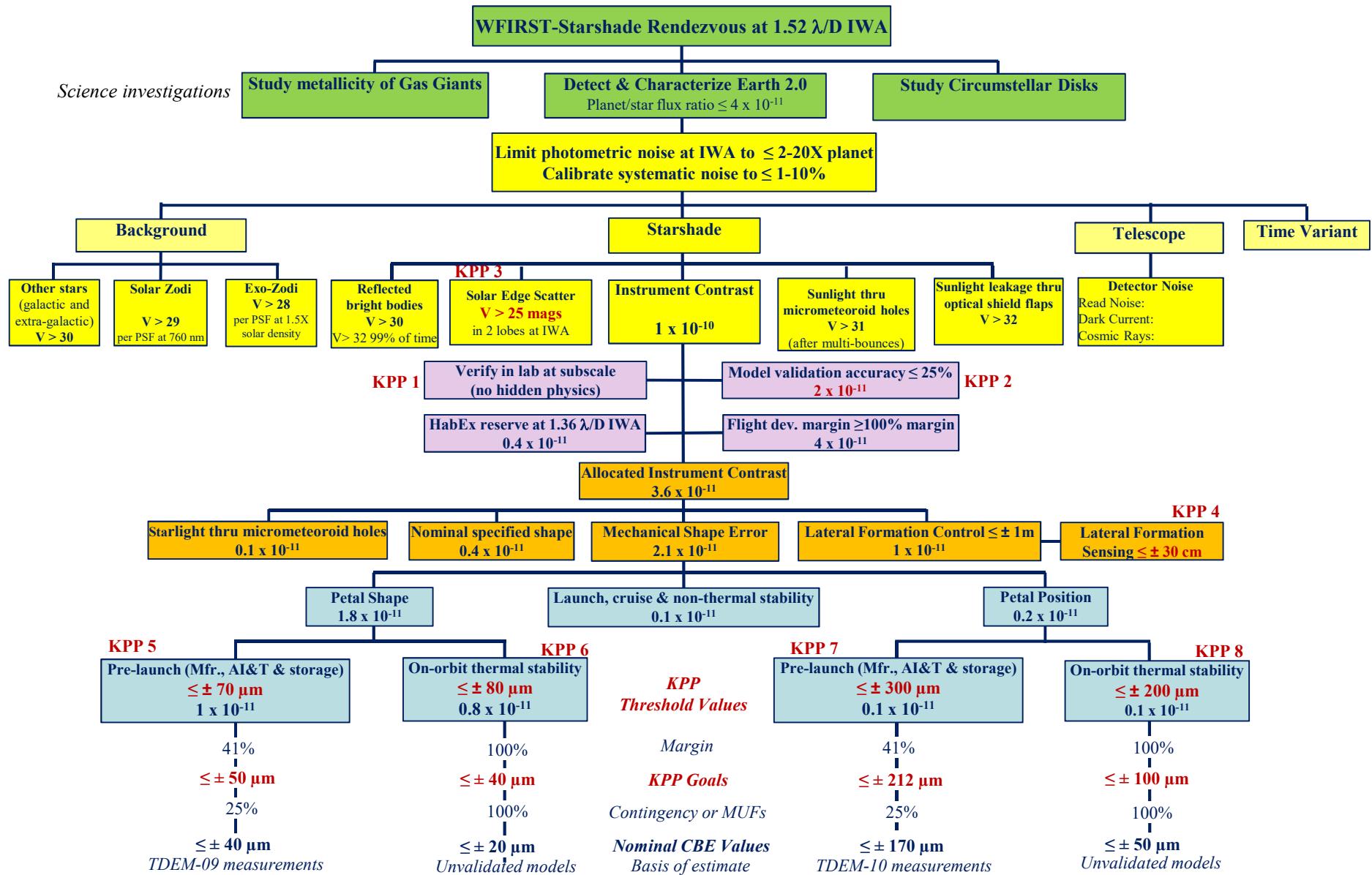
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S5 Top Level Error Budget

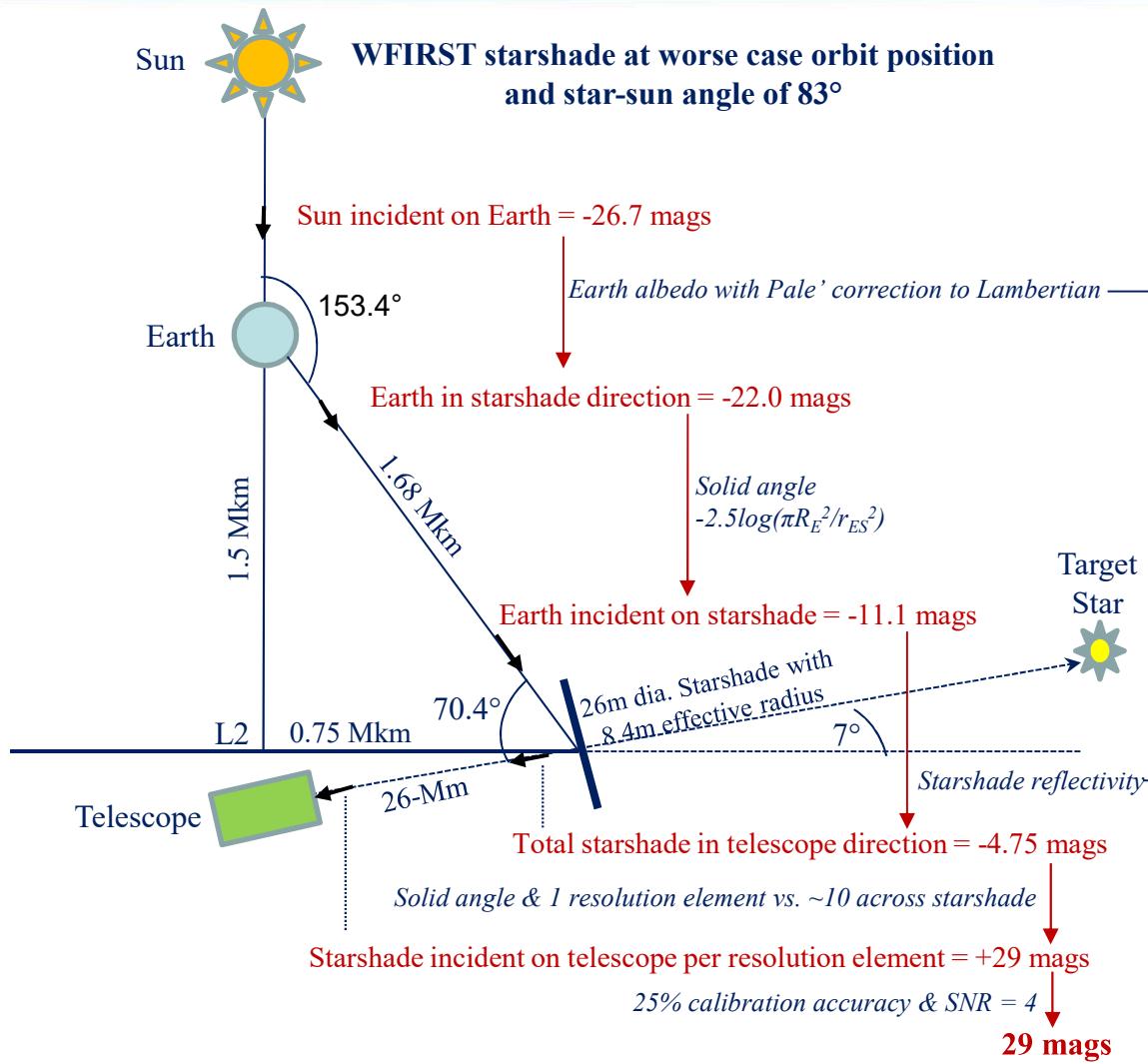


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Reflected Earth Light



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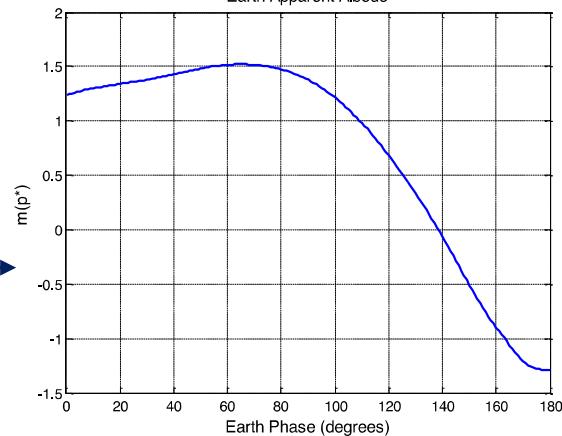
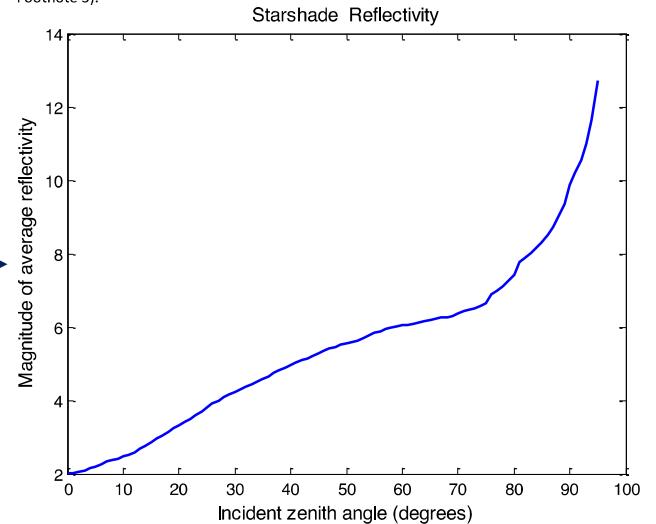


Figure 2: Magnitude of apparent albedo of Earth. Generated using data from Figure 6 of Pallé (see Footnote 3).



Both curves from 2014 Martin Regehr memo

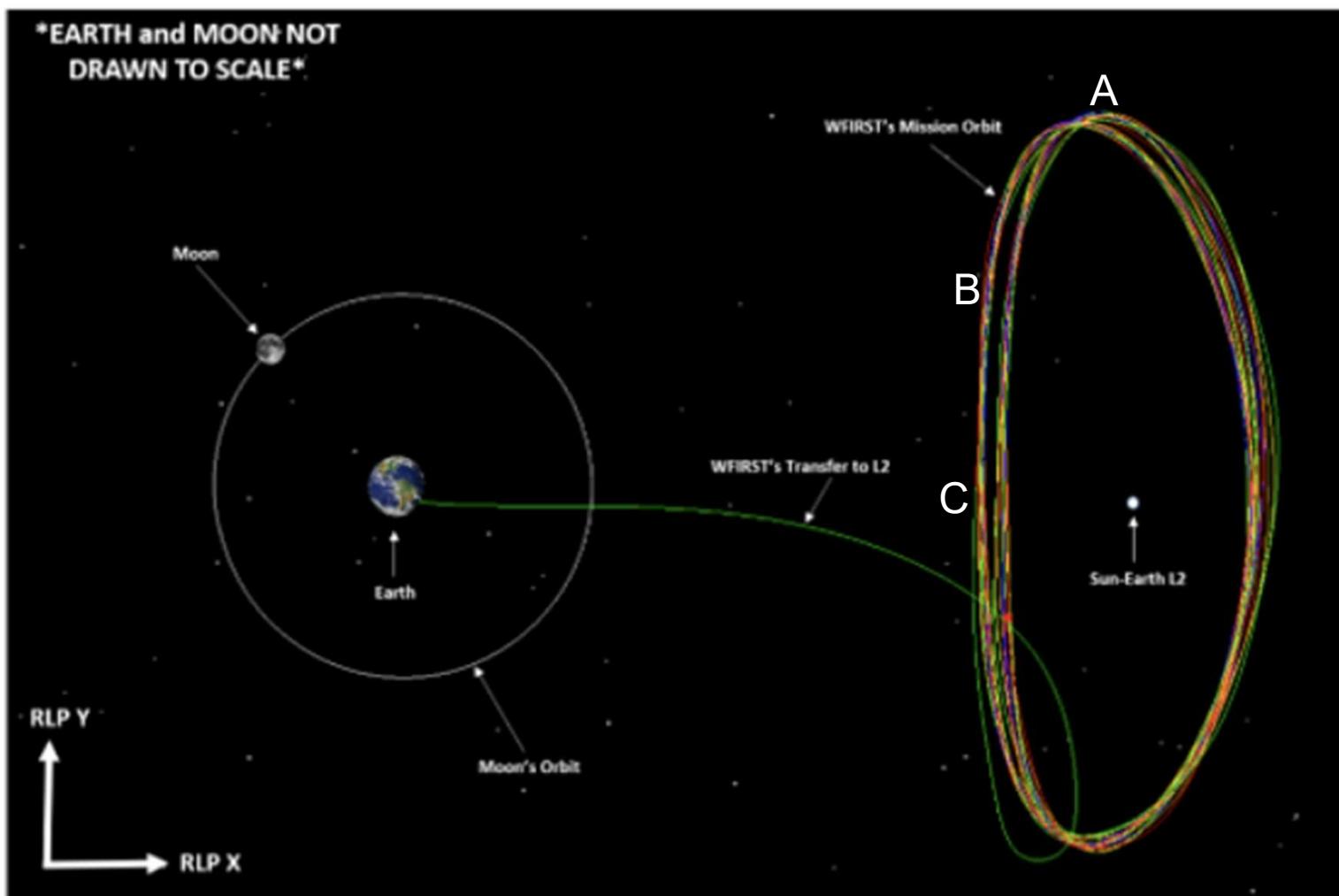
At extreme orbit positions and star-sun angles, Earth-shine exceeds the 30 mag goal and requires a darker starshade and/or restricted pointing



Earth-shine orbit geometry



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Reflected Earth Spreadsheet



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Orbit position	Orbit Extreme & 83° Sun angle
Distance along Sun-Earth line (Mkm)	1.5
Distance normal to Sun-Earth line (Mkm)	0.75
Starshade angle from Sun-Earth line (deg)	26.6
Earth phase angle (deg)	153.4
Earth phase angle (rad)	2.6779
Lambertian illumination factor	0.0069
Lambertian illumination (mags)	5.40
Pale' adjustment to Lambertian (mags)	-0.7
Earth albedo in starshade direction (mags)	4.70
Solar magnitude at 1AU	-26.7
Earth brightness in starshade direction (mags)	-22.0
Earth radius (m)	6.40E+06
Distance Earth-Starshade (m)	1.68E+09
Solid angle view factor (mags)	10.85
Earth incident on starshade (mags)	-11.15
Star-Sun observing angle (deg)	83.0
Earth incidence angle at starshade (deg)	70.4
Starshade reflectivity at Earth incidence (mags)	6.4
Starshade in direction of telescope (mags)	-4.75
Fraction within 1 resolution element	0.10
Fraction within 1 resolution element (mags)	2.5
Effective SS radius (m)	8.4
Separation distance (m)	2.60E+07
Solid angle view factor (mags)	31.21
Earth mags at telescope per resolution element	29.0



Micrometeoroids- Shielding



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- Starshade optical shield provides opacity, limits micrometeoroid holes and limits the solid angle view factor to Sun through through-holes
- It consists of 3 evenly spaced 1-mil thick Kapton layers with 16-mm total thickness
- Stopped particle mass is computed at near-max particle speed (V) of 50 km/s and 2.5 gm/cm³ particle density (larger particles are aggregates with lower density)
- Particles masses > 1 E-12 gm are expected to enter the shield per:
 - $m = t * (K r^{1/6} V^{7/8})^{-2.84}$ (NASA Std. PD-EC-1107), where t is layer thickness (cm), K is an experimental factor of 1.1 (scaled with density for Kapton vs. Al test data), r is radius in cm and V is velocity in km/s
- Particles masses > 6.3 E-7 gm are also expected to pass the mid-layer per:
 - $m=t / C * S'^2 / V$ (SMAD), where C is an experimental factor of 100 (scaled with density for Kapton vs. Al test data, S' is path distance between layer 1-2 at 45 deg trajectory)
- Particles masses > 2.5 E-6 gm are also expected to exit the shield per:
 - Same as mid-layer computation but with S' for layer 1-3
 - Mid-layer is conservatively neglected in this computation



Micrometeoroid Spreadsheet



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Parameter			Value	Comments
Flux model	Rendezvous starshade surface area (m ²)	250	26-m starshade, HabEx area is 900 m ² for 52-m starshade	
	Mission duration (years)	3	~160 Ms, Rendezvous prime mission is 2 yrs, HabEx is 5	
	$F = c_0 [(c_1 m^{0.306} + c_2)^{4.38} + c_3 (m + c_4 m^2 + c_5 m^4)^{-0.36} + c_6 (m + c_7 m^2)^{-0.85}]$		Grun model, as used on JWST, gives # larger particles per	
	c_0	3.16E+07	m ² per year vs. particle mass in grams.	
	c_1	2200		
	c_2	15	Flux direction is concentrated in 6 sectors, but effectively	
	c_3	1.3E-09	isotropic after considering variable starshade pointing.	
	c_4	1.0E+11		
	c_5	1.0E+27		
	c_6	1.3E-16		
	c_7	1.0E+06		
Particle	Particle density for shield perf., r (gm/cm ³)	2.5	Solid silica (<1e-6 gm), applies to entry hole calcs.	
	Particle density for particle sizing, r (gm/cm ³)	1	Aggregates (>1e-6 gm), conservatively applies to all particle sizing	
	Particle velocity (km/s)	50	10-70 km/s full range, 20 km/s avg, ~30% flux at 55 km/s	
	Particle mass bin size factor	1.58489	gives 5 mass bins per decade	
Shield	Construction	Kapton	~1.4 gm/cm ³ density, 3 equal & evenly spaced layers	
	Thickness, t (mil)	1		
	Thickness, t (cm)	0.00254		
Entry hole area	Experimental factor for Kapton (K)	1.1	Scaled as 1/density with 0.54 for Al	
	Smallest particle mass passing outer layer (gm)	1.3E-12	$m = t^* (K^{1/6} V^{7/8})^{-2.84}$ (NASA Std. PD-EC-1107)	
	Rounding down to smaller bin size	1.0E-12		
	# Larger particles	7.4E+05	From table to right	
	Cum larger particle area (cm ²)	1.3	From table to right	
	Ratio crater to particle diameter	1.25	By observation	
	Cum entry hole total area, both sides (cm ²)	2.0	Includes both sides of starshade	
	Cum entry hole area per side (ppm)	0.4	Each side of starshade	
Mid-Layer hole area	Separation between shield layers, S (cm)	0.8	3 evenly spaced layers, total thickness of 1.6 cm	
	Average path length, S' (cm)	1.128	Avg. 45° off-normal trajectory	
	Experimental factor for Kapton, C	100	50 for Al, assume K scales as 1/r (melt temp must be factor)	
	Smallest particle mass passing mid-layer (gm)	6.5E-07	$m = t / C * S'^2 / V$ (SMAD), if particle fully vaporized at entry	
	Rounding down to nearest bin size (gm)	6.3E-07		
	# Larger particles	1342	From table to right	
	Cum larger particle area (cm ²)	0.51	From table to right	
	Ratio Layer 2 hole to particle diameter	20	Penetrating portion of gas cloud expanding over 0.8 cm	
	Cum mid-layer hole area (cm ²)	204		
	Cum mid-layer hole area, both directions (ppm)	82	Allocation is 500 ppm in all 3-layers	
Exit hole area	Separation between shield layers, S (cm)	1.6	Neglect benefit of mid-layer	
	Average path length, S' (cm)	2.256	45° off-normal trajectory	
	Experimental factor for Kapton, C	100	Scaled as 1/density with 50 for Al	
	Smallest particle mass exiting outer layer (gm)	2.6E-06	$m = t/C * S'^2 / V$ (SMAD), if particle fully vaporized at entry	
	Rounding down to nearest bin (gm)	2.5E-06		
	# Larger particles	359	From table to right	
	Cum larger particle area (cm ²)	0.31	Integrate table to right	
	Ratio exit hole to particle diameter	40	Penetrating portion of gas cloud expanding over 1.6 cm	
	Cum exit hole total area (cm ²)	499	Includes both sides of starshade	
	Cum exit hole area per side (ppm)	100	Allocation is 500 ppm in all 3-layers	

Particle mass (gm)	Mean bin mass (gm)	# larger particles	# particles in mass bin	Mean in bin particle dia. (cm)	Mean cross-sectional area (cm ²)	Cum in bin particle area (cm ²)	Cum larger particle dia. (cm ²)	Cum fractional particle area
1.0E-03	8.2E-04	1	0	1.2E-01	1.1E-02	0.00		
6.3E-04	5.1E-04	1	1	9.9E-02	7.8E-03	0.01	0.01	0.005
4.0E-04	3.2E-04	2	1	8.5E-02	5.7E-03	0.01	0.01	0.011
2.5E-04	2.0E-04	3	2	7.3E-02	4.2E-03	0.01	0.02	0.019
1.6E-04	1.3E-04	6	4	6.3E-02	3.1E-03	0.01	0.04	0.029
1.0E-04	8.2E-05	10	7	5.4E-02	2.3E-03	0.02	0.05	0.042
6.3E-05	5.1E-05	17	12	4.6E-02	1.7E-03	0.02	0.07	0.058
4.0E-05	3.2E-05	29	20	4.0E-02	1.2E-03	0.02	0.10	0.077
2.5E-05	2.0E-05	50	33	3.4E-02	9.1E-04	0.03	0.13	0.101
1.6E-05	1.3E-05	83	54	2.9E-02	6.7E-04	0.04	0.17	0.129
1.0E-05	8.2E-06	137	86	2.5E-02	4.9E-04	0.04	0.21	0.162
6.3E-06	5.1E-06	223	135	2.1E-02	3.6E-04	0.05	0.26	0.200
4.0E-06	3.2E-06	359	208	1.8E-02	2.7E-04	0.06	0.31	0.242
2.5E-06	2.0E-06	567	314	1.6E-02	2.0E-04	0.06	0.37	0.290
1.6E-06	1.3E-06	880	462	1.4E-02	1.4E-04	0.07	0.44	0.342
1.0E-06	8.2E-07	1342	667	1.2E-02	1.1E-04	0.07	0.51	0.396
6.3E-07	5.1E-07	2009	939	9.9E-03	7.8E-05	0.07	0.58	0.453
4.0E-07	3.2E-07	2948	1290	8.5E-03	5.7E-05	0.07	0.66	0.510
2.5E-07	2.0E-07	4238	1730	7.3E-03	4.2E-05	0.07	0.73	0.567
1.6E-07	1.3E-07	5968	2266	6.3E-03	3.1E-05	0.07	0.80	0.621
1.0E-07	8.2E-08	8234	2900	5.4E-03	2.3E-05	0.07	0.86	0.673
6.3E-08	5.1E-08	11134	3634	4.6E-03	1.7E-05	0.06	0.93	0.720
4.0E-08	3.2E-08	14768	4465	4.0E-03	1.2E-05	0.05	0.98	0.763
2.5E-08	2.0E-08	19233	5376	3.4E-03	9.1E-06	0.05	1.03	0.801
1.6E-08	1.3E-08	24609	6316	2.9E-03	6.7E-06	0.04	1.07	0.833
1.0E-08	8.2E-09	3.1E+04	7.2E+03	2.5E-03	4.9E-06	0.04	1.11	0.861
6.3E-09	5.1E-09	3.8E+04	8.1E+03	2.1E-03	3.6E-06	0.03	1.14	0.884
4.0E-09	3.2E-09	4.6E+04	9.0E+03	1.8E-03	2.7E-06	0.02	1.16	0.902
2.5E-09	2.0E-09	5.5E+04	1.0E+04	1.6E-03	2.0E-06	0.02	1.18	0.917
1.6E-09	1.3E-09	6.5E+04	1.1E+04	1.4E-03	1.4E-06	0.02	1.19	0.930
1.0E-09	8.2E-10	7.6E+04	1.2E+04	1.1E-03	1.1E-06	0.01	1.21	0.940
6.3E-10	5.1E-10	8.9E+04	1.4E+04	9.9E-04	7.8E-07	0.01	1.22	0.949
4.0E-10	3.2E-10	1.0E+05	1.6E+04	8.5E-04	5.7E-07	0.01	1.23	0.956
2.5E-10	2.0E-10	1.2E+05	1.9E+04	7.3E-04	4.2E-07	0.01	1.24	0.962
1.6E-10	1.3E-10	1.4E+05	2.3E+04	6.3E-04	3.1E-07	0.01	1.24	0.968
1.0E-10	8.2E-11	1.6E+05	2.9E+04	5.4E-04	2.3E-07	0.007	1.250	0.973
6.3E-11	5.1E-11	1.9E+05	3.5E+04	4.6E-04	1.7E-07	0.006	1.256	0.978
4.0E-11	3.2E-11	2.3E+05	4.3E+04	4.0E-04	1.2E-07	0.005	1.262	0.982
2.5E-11	2.0E-11	2.7E+05	5.2E+04	3.4E-04	9.1E-08	0.005	1.266	0.985
1.6E-11	1.3E-11	3.2E+05	6.1E+04	2.9E-04	6.7E-08	0.004	1.270	0.988
1.0E-11	8.2E-12	3.8E+05	7.2E+04	2.5E-04	4.9E-08	0.004	1.274	0.991
6.3E-12	5.1E-12	4.5E+05	8.3E+04	2.1E-04	3.6E-08	0.003	1.277	0.994
4.0E-12	3.2E-12	5.4E+05	9.6E+04	1.8E-04	2.7E-08	0.003	1.279	0.995
2.5E-12	2.0E-12	6.3E+05	1.1E+05	1.6E-04	2.0E-08	0.002	1.282	0.997
1.6E-12	1.3E-12	7.4E+05	1.3E+05	1.4E-04	1.4E-08	0.002	1.283	0.999
1.0E-12	8.2E-13	8.8E+05	1.6E+05	1.2E-04	1.1E-08	0.002	1.285	1.000