

THE EXO-S PROBE-CLASS STARSHADE MISSION

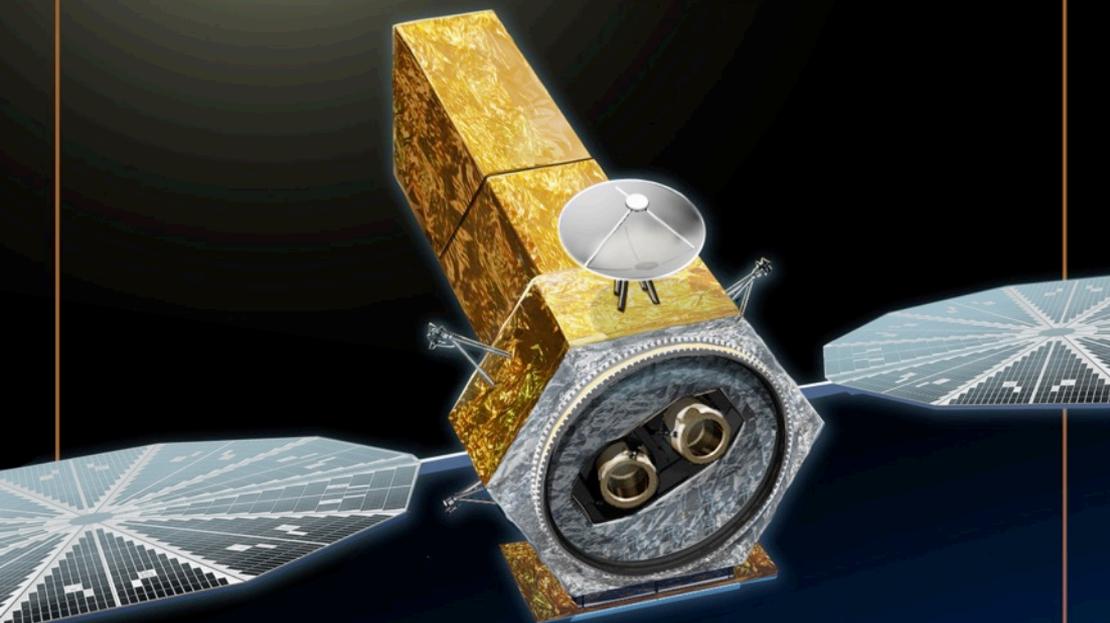
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Exoplanet Exploration Program

Jet Propulsion Laboratory, California Institute of Technology



Exo-S

Starshade Probe-Class

Exoplanet Direct Imaging Mission Concept

FINAL REPORT MARCH 2015

NASA's Recent Exoplanet Probe Studies

- ◎ The discovery of exoEarths, via a space-based direct imaging mission, is a long-term priority for US astrophysics
- ◎ Two studies of “probe sized” exoplanet direct imaging missions: one coronagraph (Exo-C), one starshade (Exo-S)
 - Total mission cost targeted at \$1B (FY15 dollars)
 - New start in 2017
 - Compelling science must be beyond the expected ground capability at the time of mission
- ◎ Studies also intended as a design input to the exoplanet community to help formulate ideas for the next Decadal Survey in 2020

Exo-S Team Members

STDT

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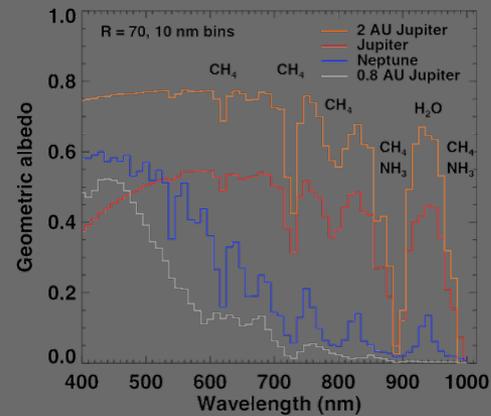
Key Areas of Advance

Science

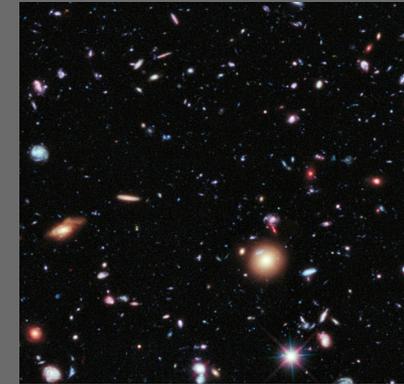
- Quantification of ability to find and identify exoEarths with a small-aperture space telescope
- Retrieval studies to understand what can be learned from giant planet atmosphere observations
- Direct imaging planet discovery challenges and solutions

Technology

- Created technology gap list for establishing TRL-5
- Developed testbed for new inner disk design that is customized to accommodate petals
- Developed concepts for optical shields, launch restraints and deployment control systems
- Focused formation flying challenge on sensing but not control



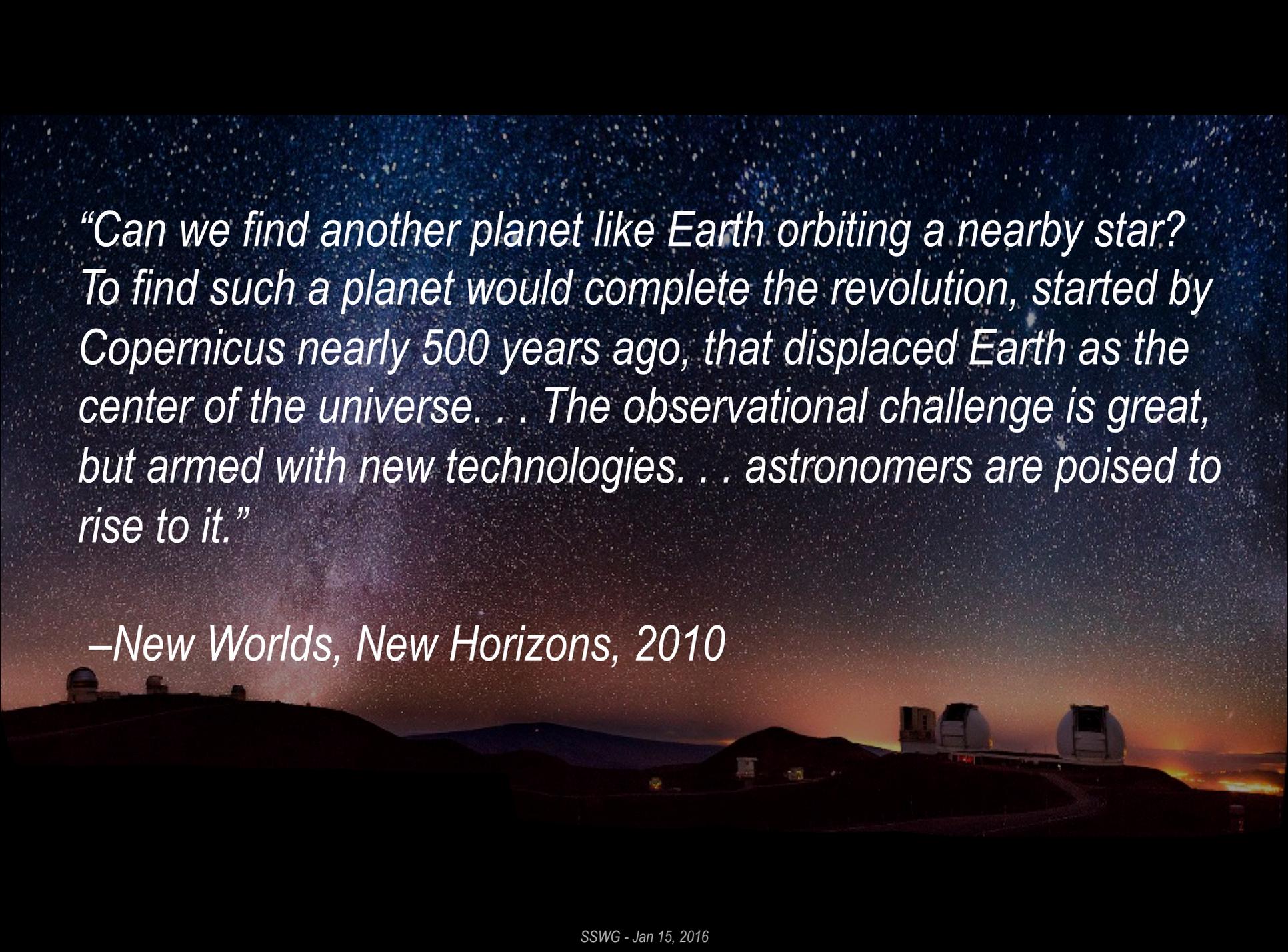
Giant planet R=70 spectra of solar system and modeled giant planets.



A one arcmin² FOV (corresponding to the planet detection field for Exo-S),



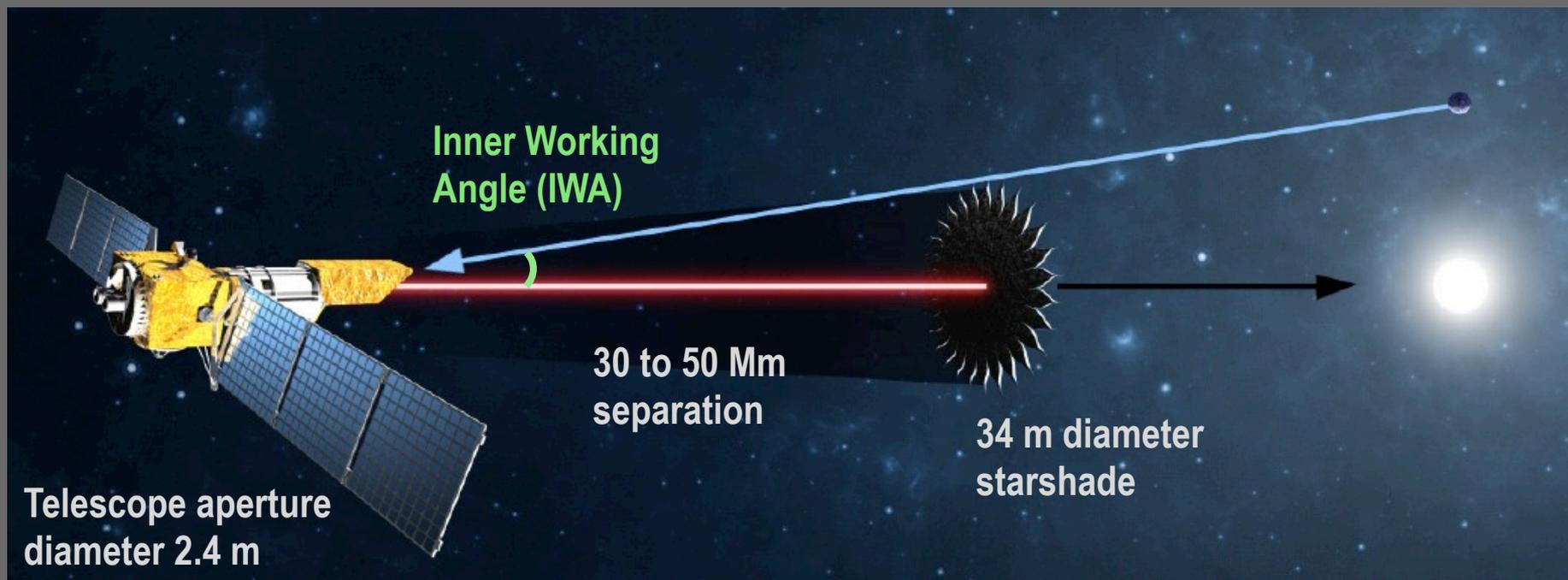
JPL half scale model starshade deployment test bed



“Can we find another planet like Earth orbiting a nearby star? To find such a planet would complete the revolution, started by Copernicus nearly 500 years ago, that displaced Earth as the center of the universe. . . The observational challenge is great, but armed with new technologies. . . astronomers are poised to rise to it.”

–New Worlds, New Horizons, 2010

Starshade Basics



- Contrast and IWA decoupled from telescope aperture size
- No outer working angle
- High throughput, broad wavelength bandpass
- High quality telescope not required
 - Wavefront correction unnecessary
- Retargeting requires long starshade slews (days to weeks)

Two Cost Constrained Exo-S Concepts

◎ Exo-S Dedicated Co-Launched Mission

- Starshade and telescope launch together to conserve cost
- Telescope: low-cost commercial Earth observer, 1.1 m diameter aperture
- Starshade: 30 m diameter
- Orbit: heliocentric, Earth-leading, Earth-drift away

◎ Exo-S **Rendezvous** Mission

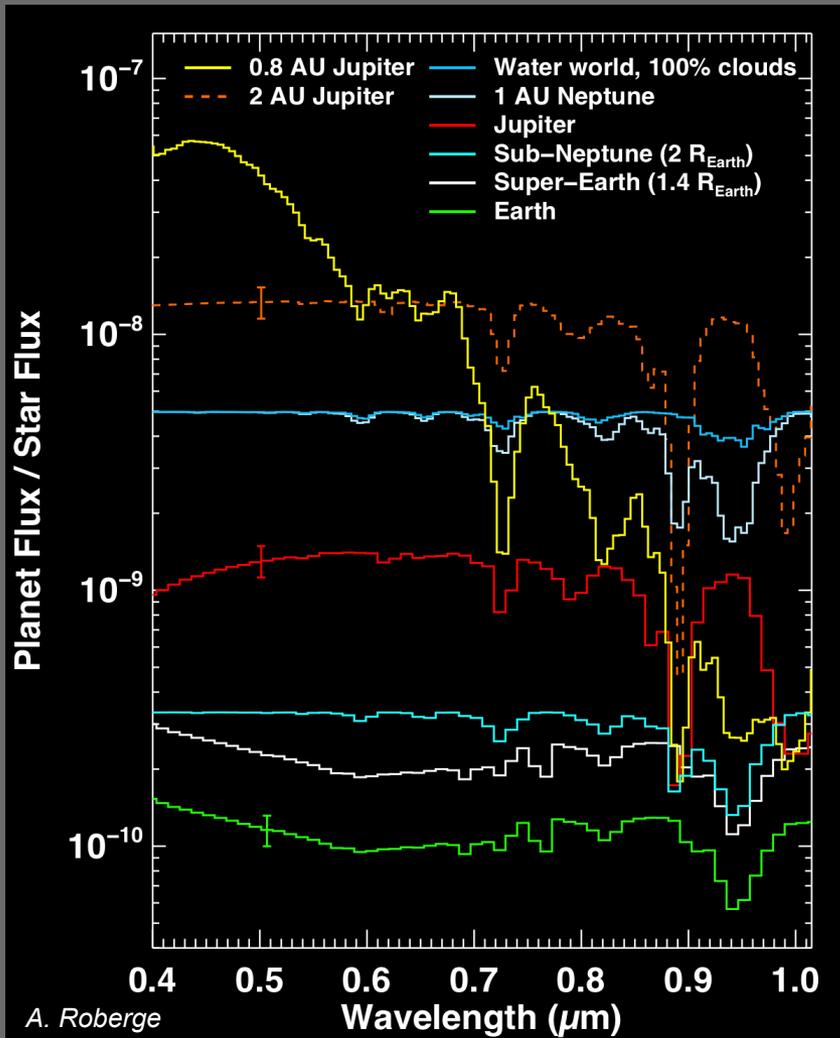
- Starshade launches for a rendezvous with an existing telescope
- Telescope: WFIRST/AFTA 2.4 m is adopted
- Starshade: 34 m diameter
- Orbit: Earth-Sun L2 (assumption for the purposes of the Exo-S study)
- Three year Class C mission
- Minimal impact to current mission design
 - No stringent requirements are imposed on the WFIRST/AFTA spacecraft
 - No new instrument, only modification to the existing coronagraph

**WFIRST/AFTA + Starshade
simulated image of
Beta Canum Venaticorum
plus solar system planets
(8.44 pc, G0V)**



Image credit: M. Kuchner

Exo-S Science Goals

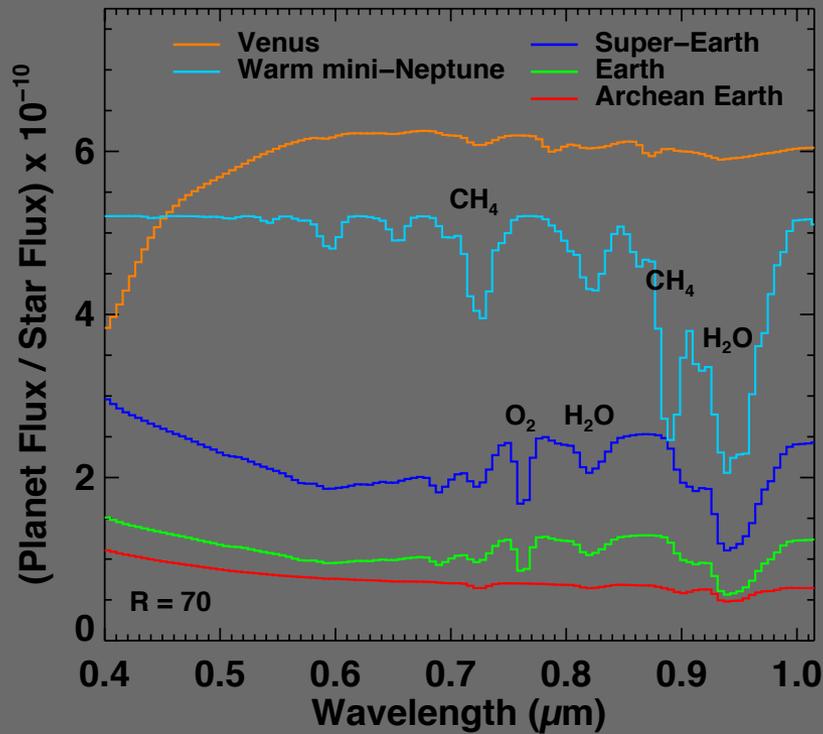


Simulated R=70 planet spectra for the 2.4-m mission, with three representative 10% error bars.

1.1-m mission cannot reach R=70 on small planets.

1. Discover new exoplanets from giants down to Earth size
2. Characterize new planets with R=10 to 70 optical spectra
3. Characterize known giant planets with R=70 spectra and constrain masses
4. Study planetary systems including circumstellar dust
 - Locate dust parent bodies
 - Evidence of unseen planets
 - Exozodi assessment for future missions

Exo-S Science Goals

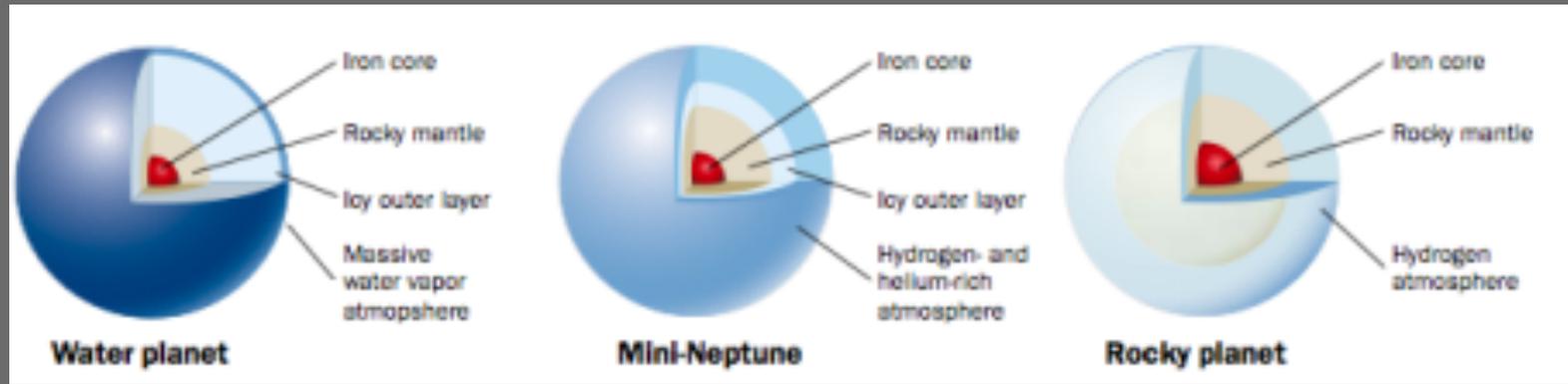


1. Discover new exoplanets from giants down to Earth size
2. Characterize new planets with $R=10$ to 70 optical spectra
3. Characterize known giant planets with $R=70$ spectra and constrain masses
4. Study planetary systems including circumstellar dust

Above: Simulated spectra of small planets. The Earth, Venus, and super Earth models are from the Virtual Planet Laboratory (VPL; <http://depts.washington.edu/naivpl/>). The sub-Neptune model is from Renyu Hu (personal communication). The spectra have been convolved to $R=70$ spectral resolution and re-binned onto a wavelength grid with 0.1 μm bins.

- Locate dust parent bodies
- Evidence of unseen planets
- Exozodi assessment for future missions

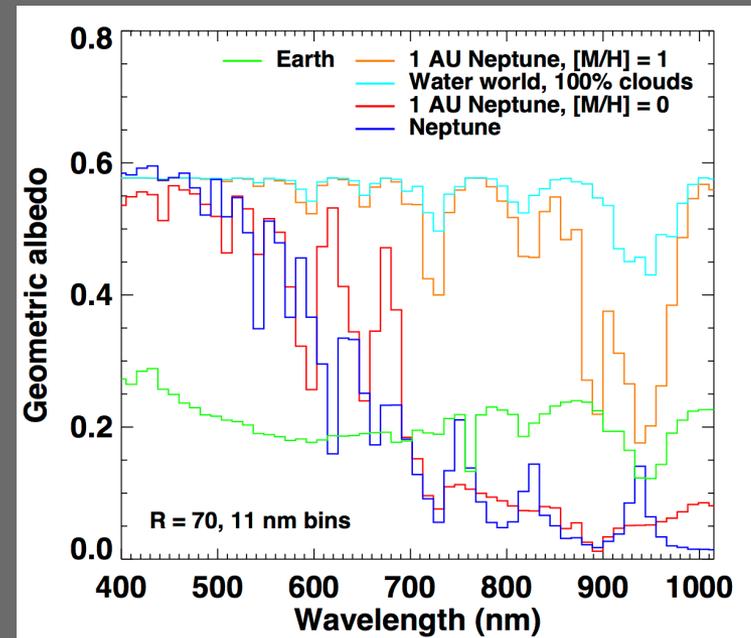
Exo-S Science Goals



Alternate views of the composition of a mini Neptune

Credit: Science News 2010

- Comparative exoplanetology for a range of planet types
- Spectra of sub Neptunes to ascertain the very nature of the low-density, extremely common, yet mysterious planets
- Potential for exoEarth discovery (and characterization) for up to 50 target stars (100 mas and $\text{lim}\Delta\text{mag} 26$)



Simulated reflected light spectra of exoplanets convolved to $R=70$

DRM Yield Summaries

| | Completeness | | |
|----------------|--------------------|--------------|--------------|
| | Case 1 | Case 2 | Case 3 |
| HZ Earth | 6.3 | 3.6 | 10.9 |
| Earth | 1.7 | 2.1 | 3.7 |
| Sup. Earth | 14.9 | 10.6 | 27.3 |
| Sub-Neptune | 30.3 | 26.8 | 52.3 |
| Neptune | 43.0 | 42.7 | 71.1 |
| Jupiter | 63.2 | 64.4 | 93.9 |
| Total | 159.5 | 150.2 | 259.2 |
| | Mean Planet Yields | | |
| | Case 1 | Case 2 | Case 3 |
| HZ Earth | 1.0 | 0.6 | 1.7 |
| Earth | 0.3 | 0.3 | 0.6 |
| Super Earth | 1.5 | 1.1 | 2.7 |
| SubNeptune | 3.0 | 2.7 | 5.2 |
| Neptune | 4.3 | 4.3 | 7.1 |
| Jupiter | 6.3 | 6.4 | 9.4 |
| Known Jupiters | 14 | 14 | 12 |
| Total | 30.4 | 29.4 | 38.8 |

Completeness is the probability of detecting planet if it's there, summed over all stars

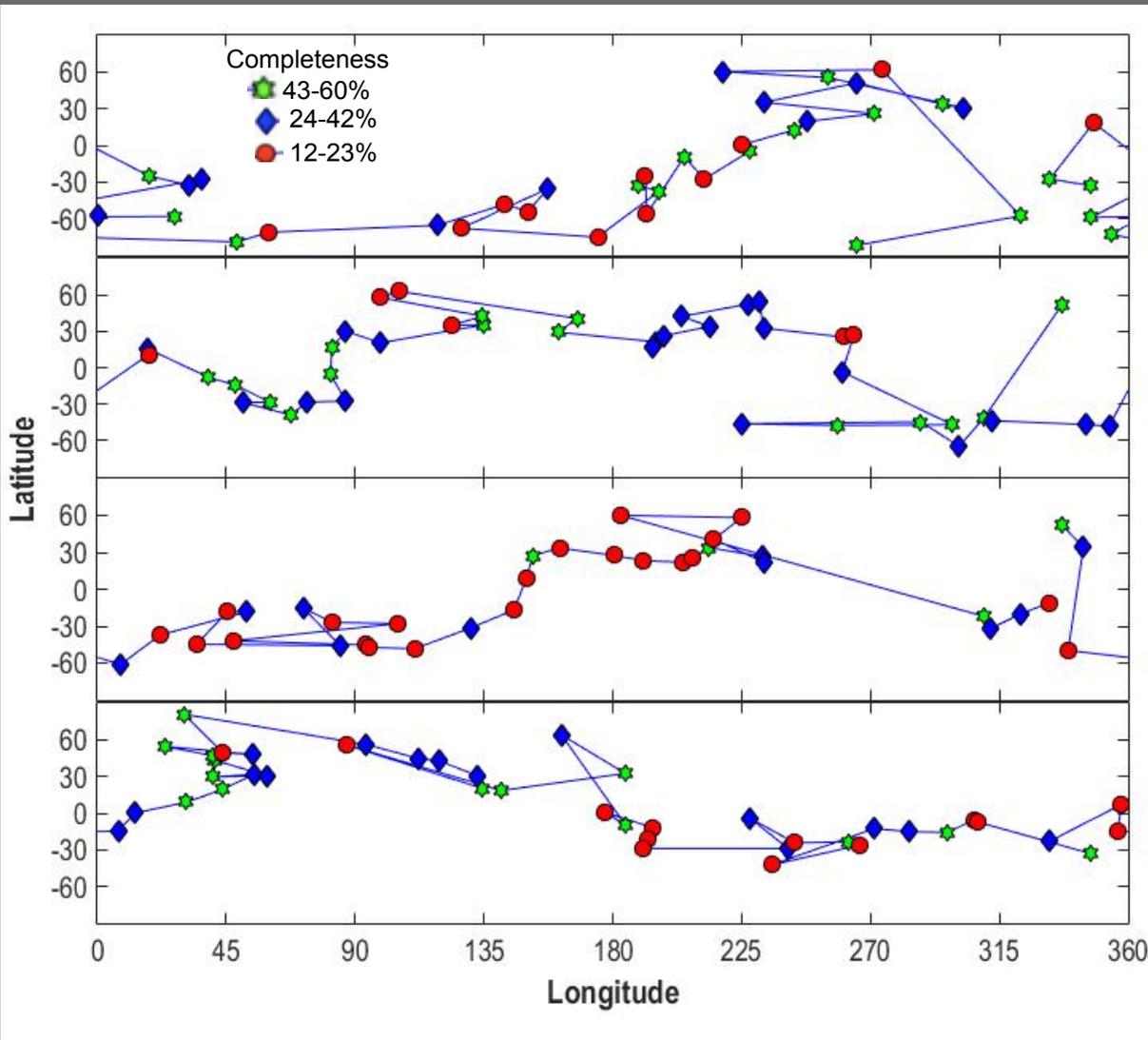
Multiply completeness by planet frequency (η) to get expected yield

Assumed $\eta = 16\%$ for Earths, $\eta = 10\%$ for all other planets

Starshade Rendezvous Earth Finder Mission

- ⦿ Enhanced concept compared to the Rendezvous Mission (3-year Class C mission)
- ⦿ A 5-year Class B mission
- ⦿ A combination of slightly larger starshade and/or increase in optical throughput with a dedicated instrument
- ⦿ Focus on habitable zones
- ⦿ Yields at least 10 exoEarth detections if η_{Earth} is $\geq 20\%$, plus a much larger number of other planet types.

Starshade Rendezvous Earth Finder Mission



- 158 Target stars
- Summed Completeness
 - Earth: 52
 - Sub-Neptune: 109
 - Jupiter: 135
- Time on Target = 701 days, 46% of mission time
- IWA 55 mas
- 28% throughput

Cost Estimates

- ◎ Cost and technology readiness analysis by Exo-S Team, JPL Team X, and Aerospace Corporation
- ◎ Estimated 2.4-m mission Phase A – F cost: **\$627M**
 - Includes launch and modifications to WFIRST
- ◎ Aerospace Corp. raised no issues with technology schedule

The cost information contained in this document is of a budgetary and planning nature and is intended for informational purposes only. It does not constitute a commitment on the part of JPL and Caltech.

Take-Away Message

WFIRST-AFTA can be leveraged for a unique and timely opportunity

- The Starshade Rendezvous Mission with 34-m starshade can access exoEarths in the habitable zone for up to 50 unique target stars
- JPL is now studying a range of mission options that fall under the “Rendezvous” mission umbrella
- Minimal modification to WFIRST needed for starshade readiness