

The NASA Exoplanet Exploration Program: The Search for Exoplanets, Habitability, and Life in our Galaxy

Gary H. Blackwood Manager, NASA Exoplanet Exploration Program February 18, 2015

> ATLAST Seminar Series Goddard Space Flight Center

NASA began the exploration of other worlds around our Sun...





To the rest of the Universe, the Earth... Looks like just another exoplanet!

Why Astrophysics?

Astrophysics is humankind's scientific endeavor to understand the universe and our place in it.







1. How did our universe begin and evolve?

2. How did galaxies, stars, and planets come to be?

3. Are We Alone?

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Our galaxy is teeming with exoplanets...

....At least one for every star in the sky

Yet our eyes can see only a few thousand stars....of the hundreds of billions within our galaxy alone

Exoplanets for \$1000, please!

May 2014



THE FIRST EARTH-SIZE EXOPLANETS WERE FOUND BY THE NASA SATELLITE LAUNCHED IN 2009 & NAMED FOR THIS 17th C. GERMAN



Even with the Kepler Mission, we have explored only a tiny corner of our galaxy



Total Exoplanets: 5013 (candidate and confirmed)

Ref.: http://exoplanetarchive.ipac.caltech.edu/docs/counts_detail.html - Updated 4 November 2014

...and we are particularly interested in those in the Habitable Zone



NASA Kepler's Hall of Fame: Small Habitable Zone Planets

As of January 2015



We've discovered an abundance of Super Earths...



...and Rogue planets not bound to any star at all!



Where will exploration take us in 100 years? Introducing the *Exoplanet Travel Bureau*







Let's call the *Exoplanet Travel Bureau*, and book a trip...

Let's visit *Kepler-186f* !

The Exoplanet Exploration Program

Astrophysics Division: Driving Documents



http://science.nasa.gov/astrophysics/documents

Here's how we are Organized

Within the NASA Science Mission Directorate



The Exoplanet Program Office is managed by the NASA Jet Propulsion Laboratory for the Astrophysics Division, NASA Science Mission Directorate

What is the Purpose of the Program?

Described in 2014 NASA Science Plan



Exoplanet Exploration Program

The Exoplanet Exploration Program aims at

- 1. Discovering planets around other stars
- 2. Characterizing their properties
- 3. Identifying candidates that could harbor life

Interdisciplinary Studies of **Exoplanets:**

Crosscutting Work Between the Astrophysics and Planetary Science Divisions

The Exoplanet Exploration Program



Extreme Precision Doppler Spectrometer

Exoplanet Exploration Program Organization Chart



Exoplanet Exploration Program

The Program relies on the Scientific Community

Active teams and committees:

- ExoTAC (Technology Assessment Committee) Chair: A. Boss, Carnegie Institute
- WFIRST/AFTA SDT (Science Definition Team) Chairs: D. Spergel, Princeton University N. Gehrels, Goddard Space Flight Center
- **STDT** (Science and Technology Definition Team). Two:
 - Exo-C (Probe Coronagraph) Chair: K. Stapelfelt, GSFC
 - Exo-S (Probe Starshade) Chair: S. Seager, MIT
- ExoPAG (Program Analysis Group) EC Chair: S. Gaudi, Ohio State University

Key Exoplanet Science Questions

1. Discovering Planets: How abundant are exoplanets in our Galaxy?

Radial Velocity

- <1 m/s
- Transit Photometry
 < 10 parts per million

2. Characterizing Planets: What are exoplanets like?

Transit Spectroscopy	< 25 parts per million	(large planets)
 Direct Imaging		
 High Contrast 	< 1E-9	(after post-processing)
Small Inner Working Angle	< 500 mas (<200 mas)	
 Spectroscopy 	R~40 in visible, near infrared	(water lines)

3. "Pale Blue Dots": Are the planets habitable? Are there signs of life?

Current Exoplanet Science Missions

Kepler Space Telescope



- **PI:** W. Borucki, NASA Ames Research Center
- Launch Date: March 6, 2009
- Science Data Collection through May 2013
- Final processing of full data set underway

Kepler Closeout

Harvesting the exoplanet yield from the mission

- Already available to Community: Q0-Q16
- Uniform Processing: Q0-Q17 (9.2)
 - Long cadence light curves
 - Short cadence light curves
 - Release notes

- Dec 2014 Mar 2015 Jul 2015
- Final Data processing: Q0-Q17 (9.3)
 - Light curves
 - Release notes

Jan 2016 Aug 2016

Kepler (K2) is now observing ~80-day windows in the ecliptic

Hubble – an Exoplanet Observatory

Advancing the art of coronagraphy and transit spectroscopy







JWST – another Exoplanet Observatory

- Transit spectroscopy and photometry (1-20 μm)
- Coronagraphic imaging at 3 23 μm of planets (young Jupiters to Saturns)
- Spectra of coolest brown dwarfs (free floating planets)





TESS

WFIRST / AFTA

Wide-Field Infrared Survey Telescope (WFIRST) Astrophysics Focused Telescope Assets (AFTA)

> Goddard Space Flight Center Jet Propulsion Laboratory STScl NExScl

Wide-field Instrument

- H4RG detectors (Qty 18)
- Wavelength: 0.6 to 2.0 micron
- FOV: 0.28 deg^2

Wide-field Instrument Science

- Dark Energy
- Infrared Survey
- Microlensing survey for exoplanets



WFIRST / AFTA Microlensing survey completes the census begun by Kepler

But wait, there's more! the WFIRST / AFTA Coronagraph Direct Imaging of Exoplanet Nearest Neighbors



Coronagraph Instrument

- Imaging and spectra channels
- 0.4 1 μ m bandpass
- $\leq 10^{-9}$ detection contrast
- 100 mas inner working angle at 0.4 μm
- R ~ 70

Coronagraph Science

- Imaging and spectroscopy of exoplanet atmospheres down to a few Earth masses
- Study populations of debris disks



Coronagraph will develop the technologies for a future exo-Earth mission

WFIRST Coronagraph images cool gas and ice giants



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WFIRST SDT Final Report



Wide-Field InfraRed Survey Telescope-Astrophysics Focused Telescope Assets WFIRST-AFTA 2015 Report by the Science Definition Team (SDT) and WFIRST Study Office January 30, 2015

- Final SDT Report Delivered to HQ on 1/30/2015
- Presentation to John Grunsfeld and Paul Hertz 2/13
- Public release afterwards, following incorporation of comments from HQ and presentation.

Science Definition Team

D. Spergel¹, N. Gehrels² C. Baltay³, D. Bennett⁴, J. Breckinridge⁵, M. Donahue⁶, A. Dressler⁷, B. S. Gaudi⁸, T. Greene⁹, O. Guyon¹⁰ C. Hirata⁸, J. Kalirai¹¹, N. J. Kasdin¹, B. Macintosh¹², W. Moos¹³, S. Perlmutter¹⁴, M. Postman¹¹, B. Rauscher² J. Rhodes¹⁵, Y. Wang^{16,17}, D. Weinberg⁸,

> Ex Officio D. Benford¹⁸, M. Hudson¹⁹, W. -S. Jeong²⁰, Y. Mellier²¹, W. Traub¹⁵, T. Yamada²²

Consultants P. Capak¹⁷, J. Colbert¹⁷, D. Masters¹⁷, M. Penny⁶, D. Savransky²³, D. Stern¹⁵

Study Team R. Barry², L. Bartusek², K. Carpenter², E. Cheng²⁴, D. Content², F. Dekens¹⁵, R. Demers¹⁵, K. Grady², C. Jackson²⁵ G. Kuan¹⁵, J. Kruk², M. Melton², B. Nemati¹⁵, B. Parvin¹⁵, I. Poberezhskiy¹⁵, C. Peddie², J. Ruffa² J.K. Wallace¹⁵, A. Whipple²⁴, E. Wollack², F. Zhao¹⁵



Exoplanet Exploration: A Decade Horizon

NASA and ESA efforts

ExoPlanet Exploration Program

ExEP



EXOPLANET Q&Alien


Enabling and Creating the Exo-Future: Science and Mission Studies

Key Exoplanet Science Questions

1. Discovering Planets: How abundant are exoplanets in our Galaxy?

- Radial Velocity
- Transit Photometry
- <1 m/s
- < 10 parts per million

2. Characterizing Planets: What are the (large) exoplanets like?

Transit Spectroscopy

< 100 parts per million

- **Direct Imaging** _
 - High Contrast
 - Small Inner Working Angle
 - Spectroscopy

- (after post-processing) < 500 mas (<200 mas)
- R~40 in visible, near infrared
 - (water lines)

3. "Pale Blue Dots": Are the planets habitable? Are there signs of life?

< 1E-9

Key Exoplanet Science Questions

1. Discovering Planets: How abundant are exoplanets in our Galaxy?

2. Characterizing Planets: What are the (large) exoplanets like?

3. "Pale Blue Dots": Are the planets habitable? Are there signs of life?

- Transit Spectroscopy
- Direct Imaging
 - High Contrast
 - Small Inner Working Angle
 - Spectroscopy
 - η_Earth
 - Exozodiacal Dust
 - Yield

- < 1 part per million
- < 1E-10
- < 100 mas (<40 mas)

R~70 in visible, near infrared Quantify, for mission design Quantify, for mission design Ideally: dozens of rocky planets (after post-processing)

(biosignature gases)

Large Binocular Telescope Interferometer

Measures exozodiacal dust in habitable zones

University of Arizona P. Hinz, PI



LBTI Performance

- 24 zodi (one sigma) achieved during November 2014 run
- LBTI will characterize the exo-zodiacal dust emissions of 50 target stars in mid-IR to a level of 3 - 6 zodi (one sigma)

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NASA/NSF Partnership for Exoplanet Research

Extreme Precision Doppler Spectrometer

- Scope:
 - Exoplanet-targeted Guest Observer program with existing instrumentation on WIYN using NOAO share (40%) of telescope time
 - Solicitation for facility-class extreme precision radial velocity spectrometer for WIYN telescope (commissioning goal: 2018)
- Motivation
 - Follow-up of current missions (K2, TESS, JWST)
 - Pathfinder observations inform design/operation of future missions
- Timeline:
 - Dec 2014: community announcement
 - Jan 2015: amendment to ROSES 2014 NRA



3.5m WIYN Telescope Kitt Peak National Observatory Arizona

The Exoplanet Direct-Imaging Challenge



Probe-Scale studies

High-Contrast Imaging



Exo-C:

Internal Occulter (Coronagraph)

K. Stapelfeldt, STDT Chair, GSFC



Exo-S:

External Occulter (Starshade)

S. Seager, STDT Chair, MIT





Enabling the Exo-Future: Technology Development

Technology Development for Coronagraphs



Xinetics

e2v Electron Multiplying CCD

Extracting Exoplanets from the Speckles

Effect of post-processing on raw coronagraph images

- JPL-simulated fields from observation sequence of 47 Uma using HLC coronagraph
- Post-processing done at STScl "blindly" using PCA-KLIP algorithm (Soummer et al. 2010)
- Two known planets correctly retrieved





Pueyo, Soummer et al. STScl

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The External Occulter: Starshade

The starshade could launch together with a telescope. Once in space, it would split off and move into position to block the starlight.

Starshade Rendezvous

34 meter diameter

At L2 with WFIRST/AFTA

Primary bandpass: 600 – 850 nm Raw contrast: 1 × 10⁻¹⁰ IWA: 100 milliarcsec

2.4 meter telescope



Starshade Rendezvous

Possible Science

- Observe 52 stars in 2 years
- 13 known exoplanets
- 19 HZ targets. Expect
 ~ 2 Earths or Super-Earths
- Can detect sub-Neptunes to Jupiters around all HZ targets and 20 additional stars



Technology Development for Starshades (External Occulters)



Deployment Testing at Northrop Grumman (Astro-Aerospace)

Demonstration of starshade development model

Desert Testing of Starshades



S. Warwick, Northrop Grumman

Desert Testing of Starshades



S. Warwick, Northrop Grumman

Exo-Earths require large telescopes



Stark et al, 2014 For Coronagraphs

- Yield most sensitive to (in order):
 - Telescope diameter
 - Coronagraph inner working angle
 - Coronagraph contrast
 - Coronagraph noise floor
- Also sensitive to η_earth (strong) and exozodical dust (relatively weak)

Formative Era: Large UV-Optical-IR Telescope



Gravitational Wave Surveyor
GRAV WAVE
CMB POL
CMB Polarization Surveyor FAR IR
UVOIR LUVOIR Surveyor
XRAY
Xray Surveyor

	Surveyor
Formation flying	
Interferometry: precision metrology	
X-ray interferometry	n - seda
High-contrast imaging techniques	
Optics deployment and assembly	
Broadband coatings	
X-ray optics	
Large-format detector arrays	
New detector capabilities	
Cryogenics	

LUVOIR

Formative Era: Large UV-Optical-IR Telescope (LUVOIR)

Optics Deployment and Assy



SiC Active Hybrid Mirror, Xinetics



MOIRE, BATC



Lightweight ULE, ITT

Starlight Suppression Systems



Visible Nuller, GSFC



Pupil Mapping, Univ. Arizona



Starshade NGAS, Princeton, JPL

Formation Flying

Broadband Mirror Coatings

Telescope Mechanical Isolation Systems

The Program Address the Key Questions

Through Science, Advanced Studies, and Technology Development

- 1. Discovering Planets: How abundant are exoplanets in our Galaxy?
 - Radial Velocity
 - Transit Photometry
- 2. Characterizing Planets: What are the (large) exoplanets like?
 - Transit Spectroscopy
 - Direct Imaging



- Transit Spectroscopy
- Direct Imaging
 - High Contrast
 - Small Inner Working Angle
 - Spectroscopy
 - η_Earth
 - Exozodiacal Dust
 - Yield









1/17mm





Engaging the Science Community

NASA Exoplanet Science Institute

Archives, Tools, and Professional Education



Sagan Fellowships and Workshops





Kepler Community Follow-up Program

Exoplanet Database Science Data Archives

Home	1	Nupres-10				- I -		
Introduction A	Plane	Detail	ed Informa	tion			_	
The Interactive Exoplanet Visualizer displays all exoplanet data currently		Trans exopt	it Ephemer anet.eu	*		Semi-Major Asia [AU]		
stored in the archive. Users		Gant	anata con					
download the data, and	WASP-19 .	8			008	0.0164 12 0000	1	
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Column Controls Update	WASP-12 .	b	tran	1.091423±0.0000	0.3	0.0229±0.0040	1	
Default Columna	OGLE-TR-56 +	b	tren	1.211909±0.0000	21	0.0225±0.0004	1	
First Name	HAT-P-23	b	tran	1.212884±0.0000	02	0.0232±0.0002	4	
Planet Letter	WASP-33 .	b	man	1.2196609±0.000	0012	0.02555±0.00017		
Discovery Method	TrES-3 .	b	Inan	1.30619		0.02272 *0.00017	1	
E Orbital Period [days]	HD 41004 8 🕈	b	N.	1.328300±0.0000	12	0.0177	1	
B Semi-Major Axis (ALI)	Qatar-2 🗰	D	tran	1.3371182±0.000	0037	0.02149±0.00036		
E Coentricity	WASP-4 +	b	tran	1.3382324 *0.0000	29	0.02255 *0.0008	1	
Panet Mass [Jupiter]	Qatar-1 🖷	b	tran	1.420033±0.0000	16	0.02343 *0.00028		
Planet Radius (Aprel)	WASP-46 ·	D	man	1.4303700±0.000	0023	0.02448±0.00028		
Ra (dec)	OGLE-TR-113	b	tran	1.4324757±0.000	0013	0.0229±0.0002	1	
Dec (dea)	TrES-5 #	b	Itran	1.4822446±0.000	0007	0.02446±0.00068	T	
E Z Distance (parsec)	Kepler-17 +	b	tran	1.4857108±0.000	0002	0.02591 *0.00037	-	
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NExScI hosts an archive of all data ever acquired on all Keck Instruments NExScI provides science center support for the WFIRST Coronagraph

Program Engages the Public

EYES ON **EXOPLANETS**

explore a visual database of new worlds >



http://planetquest.jpl.nasa.gov



Ways to Become Involved

ExoPAG (Program Analysis Group)

- Solicits and coordinates community input
- Provides analysis findings through Astrophysics subcommittee of the NASA Advisory Council

ExoPAG Executive Committee members

Scott Gaudi Chair	Ohio State University
Rus Belikov	NASA Ames Research CenterÂ
Nick Cowan	Northwestern University
Jonathan Fortney	University of California, Â Santa Cruz
Dave Latham	Harvard Smithsonian Center for Astrophysics
Amy Lo	Northrop Grumman Aerospace Systems
Peter Plavchan	Caltech/NASA Exoplanet Science Institute
Gene Serabyn	Jet Propulsion Laboratory
Remi Soummer	Space Telescope Science Institute
Maggie Turnbull	Global Science Institute
Lucianne Walkowicz	Princeton University

Active Science Analysis Groups

- Precision radial velocity
- Probe/Medium-scale direct imaging mission requirements
- Atmospheres / transit spectroscopy
- High-precision astrometry

Active Science Interest Group

 Toward a Near-Term Exoplanet Community Plan

Preparing for the Next Astrophysics Decadal Survey

Charge from Astrophysics Division Director, Jan 4 2015:

- Part A: Identify a small set (~3-4) of large missions to study
- Part B: Science and Technology Definition Teams conduct studies
 - Supported by engineering design team assigned to NASA Centers

From ADD Charge, Jan 4 2015



ExoPlanet Exploration Program



Preparing for the 2020 Decadal Survey Large Mission Concepts

The initial short list (in alphabetical order):

- FAR IR Surveyor The Astrophysics Visionary Roadmap identifies a Far IR Surveyor as contributing through improvements in sensitivity, spectroscopy, and angular resolution.
- Habitable-Exoplanet Imaging Mission The 2010 Decadal Survey recommends that a habitable-exoplanet imaging mission be studied in time for consideration by the 2020 decadal survey.
- UV/Optical/IR Surveyor –The Astrophysics Visionary Roadmap identifies a UV/Optical/IR Surveyor as contributing through improvements in sensitivity, spectroscopy, high contrast imaging, astrometry, angular resolution and/or wavelength coverage. The 2010 Decadal Survey recommends that NASA prepare for a UV mission to be considered by the 2020 Decadal Survey.
- X-ray Surveyor The Astrophysics Visionary Roadmap identifies an X-ray Surveyor as contributing through improvements in sensitivity, spectroscopy, and angular resolution.

Enabling Decadal Mission Readiness

Program Office Leadership for Exoplanet Science

- What can we do <u>now</u> for a solid story in <u>2020</u> to assure TRL5 readiness by <u>2024</u>?
- Approach:
 - Engineer a technical and programmatic solution
 - Start by assessing technology success criteria and capability today => technology gaps leads to technology plan
- ExEP will facilitate, plan, advocate for technology readiness to maximize likelihood that exoplanet missions are ranked highly by the next Decadal Committee

In Closing...

...And on those other worlds, are there beings who wonder as we do? C. Sagan

We dream about other Worlds...

Now we have the means to image our nearest neighbors, To search for Habitable Worlds, and for Life in our Galaxy



¹ NASA/ESA Partnership ² CNES/ESA Imagine your role in the discovery of Habitable Worlds and Search for Life in our Galaxy



National Aeronautics and Space Administration

Jet Propulsion Laboratory California Institute of Technology Pasadena, California

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Backup

Coronagraph Masks for WFIRST/AFTA



Reflective shaped pupil mask Jeremy Kasdin, Princeton





PIAA-CMC focal plane mask (backup) Olivier Guyon, U. of Arizona



Transmissive hybrid Lyot mask John Trauger, JPL






TESS Transiting Exoplanet Survey Satellite



Standard Explorer (EX) Mission

PI: G. Ricker (MIT)

Mission: All-Sky photometric exoplanet mapping mission.

Science goal: Search for transiting exoplanets around the closest and brightest stars in the sky.

Instruments: Four wide field of view (24x24 degrees) CCD cameras with overlapping field of view–operating in the Visible-IR spectrum (0.6-1 micron).

Operations: 2017 launch with a 2-year prime mission

CURRENT STATUS:

- Major partners:
 - PI and science lead: MIT
 - Project management: NASA GSFC
 - Instrument: Lincoln Laboratory
 - Spacecraft: Orbital Science Corp
- Development progressing on plan.
 - Preliminary Design Review (PDR) successfully completed Sept 9-12, 2014.
 - Confirmation Review, for approval to enter implementation phase, successfully completed October 31, 2014.

Characterizing the Spectrum of Exoplanets

Integral Field Spectrometer

- Low spectral crosstalk needed for spectral science data
- Extracts data cube
- Used in post-processing for speckle suppression



M. McElwain, Roman Fellow, GSFC



Starshade Concept

Inner Working Angle (IWA)

> Starshade diameter 34 m

±1 m lateral control

Separation distance 37,000 km ±250 km

Telescope diameter 1.1 m

Read more about technologies, studies, and the Exoplanet Exploration Program at

http://exep.jpl.nasa.gov

 Contrast and inner working angle are decoupled from the telescope aperture size
A simple space telescope can be used
No wavefront correction is needed

No outer working angle

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Ways to Become Involved

- ExoPAG: SAGs, and SIG
- EPDS initiative
- Program and decadal studies
- Competitive Funding:
 - Exoplanet Research Program (XRP)
 - Astrophysics Data Analysis Program (ADAP, supports archival Kepler/K2 research)
 - K2 Guest observer program
 - Astrophysics Theory Program (ATP)
 - Hubble Guest Observer program (supports exoplanet research).
 - SAT / ROSES / TDEM for exoplanet technology development

Read more at: <u>http://exep.jpl.nasa.gov</u>