

In-Space Assembled Telescope (iSAT)

Steering Committee Kickoff Telecon

May 3, 2018

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Today's Agenda

- 1. Background and Motivation
- 2. Plan Forward
- 3. Participants and Roles
- 4. Process
- 5. Next Steps

Thanks to our Steering Committee

Steering Committee (C=confirmed)

	_
Dave Redding (C)	NASA JPL
Joe Pitman (C)	consultant
Scott Knight (C)	Ball Aero
Bill Doggett (C)	NASA LaRC
Matthew Greenhouse (C)	NASA GSFC
Joanne Hill-Kittle (C)	NASA GSFC
Ron Polidan (C)	consultant
John Grunsfeld (C)	NASA (ret)
Keith Belvin (C)	NASA STMI
. Brad Peterson (C)	STScI/OSU
. Florence Tan (C)	NASA SMD
. Ray Bell (C)	Lockheed
. Nasser Barghouty (C)	NASA APD
. Eric Smith (C)	NASA APD
	Joe Pitman (C) Scott Knight (C) Bill Doggett (C) Matthew Greenhouse (C) Joanne Hill-Kittle (C) Ron Polidan (C) John Grunsfeld (C) Keith Belvin (C) Brad Peterson (C) . Florence Tan (C) . Ray Bell (C) . Nasser Barghouty (C)

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Our Sponsors



Dr. Paul Hertz Director Astrophysics Division NASA Headquarters



Mike Seablom Chief Technologist Science Mission Directorate NASA Headquarters

Background

Background

- SMD is interested in advancing exoplanet science which benefits from everincreasing large telescopes to detect theses faint objects and characterize them through spectroscopy.
 - Large telescopes also advances many topics in general astrophysics
- But why do these telescopes have to cost so much? Is there a way to break the cost model of telescopes that currently goes something like \$ ∝ D²⁻³?
- Last November, Harley Thronson (NASA GSFC) led a Workshop on in-Space Servicing & Assembly



70+ participants from government, industry, and academia https://exoplanets.nasa.gov/exep/technology/in-space-assembly/

Extracted from Nov TIM Summary Report TIM Suggestions (1 of 2)

Commission a design study to understand how large-aperture telescopes could be assembled and serviced in space

- Suggest joint SMD/STMD/HEOMD study with industry and academia participation
- Multi-disciplinary, multi-institutional
- Initiate the study in time for initial results to be available to Gateway and robotics designers within 2018, but certainly before end 2019.

1. Produce several iSA concepts and prioritize them

- 2. Select one implementation concept for a deeper engineering study
 - a) identify capability needs, SOA, and technology gaps and produce a list of technologies that could be demonstrated to close these gaps
 - b) assess opportunities for engineering demonstrations that may be deployed on the ISS within the next few years.
 - c) determine balance of human and robotic support
 - d) understand servicing options
 - e) produce an early list of preliminary interface consideration to the DSG

3. Estimate the cost and understand scaling laws to compare costs/risks to an autonomously deployed telescope

#8) Why Now?

Extracted from Nov TIM Summary Report https://exoplanets.nasa.gov/exep/technology/in-space There are large future space observatories being studied and designed today to be serviceable but the servicing capabilities do not currently exist.

- There are large future space observatories being studied and designed ٠ today that are limited by current and future launch vehicle fairing sizes. "We are now hitting a wall [towards what is possible]"
- Potential space telescope missions planned to be serviced and/or • assembled in the 2030s need to start their technology activities in the 2020s.
- A valuable venue for assembly demonstrations, the ISS, may be ٠ decommissioned in the mid-2020s.
- There is a near-term opportunity to inform the 2020 Decadal Survey about ٠ the potential benefits of iSSA as a potential implementation approach for future large apertures and the current SOA.
- There is at present a window of opportunity through 2019 to recommend ٠ augmentations to the DSG team before their designs are frozen.

Study Objective and Deliverables

Study Objective:

 "When is it advantageous to assemble space telescopes in space rather than to build them on the Earth and deploy them autonomously from individual launch vehicles?"

• Deliverables:

A whitepaper by May 2019 assessing:

- 1. the telescope size at which iSA is necessary (an enabling capability)
- 2. the telescope size at which iSA is cheaper or lower risk with respect to traditional launch vehicle deployment (*an enhancing capability*)
- 3. the important factors that impact the answers (e.g., existence of HEOfunded infrastructure, architecture of space telescope (segments or other), cryogenic or not, coronagraph capable (stability) or not, etc.)

The intention of the whitepaper is to inform NASA and the 2020 Decadal Survey of the total cost and risk benefits of the iSA of space telescopes.

#3) How does iSSA reduce cost and risk, both technical and programmatic? (2 of 4)

Extracted from Nov TIM Summary Report (/exonlanets.nasa.gov/exep/technology/in-space-assembly/

Potential cost saungs onered through iSSA:

- Eliminates engineering design work and testing required to (1) creatively fit large structures into existing fairings and (2) autonomously deploy
 - JWST invested a significant effort into designing and testing the telescope's folded wing design; even more for the observatory deployment with > 100 single point failures
- Reduces need for hardware redundancy
- Reduces system "ruggedization" to survive launch environment
- Reduces need for new and unique ground test facilities
 - JWST required new ground facilities to be built
- Reduces the need for a large standing army during I&T
- Leverages existing and less-costly medium-lift LVs
- New instruments can be swapped out over longer periods of time before new additional observatories are needed

#3) How does iSSA reduce cost and risk, both technical and programmatic? (4 of 4) Extracted from Nov TIM Summary Report

Risk reduction opportunities arising from iSSA

Reducing risk becomes increasingly more important as mission costs increase.

https://exoplanets.nasa.gov/exep/technology/in-space-assembly,

- Future larger observatories are likely to require more complex ٠ deployment schemes. iSSA can mitigate risk of failure by:
 - Modularizing the design enabling repair and replacement of faulty sections
 - Designing servicing capabilities (robotic and/or human) into the architecture
 - Minimizing single-point failures -
 - Enabling end-to-end testing (often not possible on ground)
- iSA does not require next-generation launch vehicles
 - Several future mission concepts under study rely on the SLS Block II
- Launch failure need not be equivalent to mission failure •

Plan Forward

Study Charter

- Draft Charter written and will be submitted for signature by the Sponsors.
- The iSAT Study Working Group is intended to represent expert knowledge in the area of telescope design and architecture, assembly and testing across academia, NASA, and industry.

In-Space Assembled Telescope (iSAT) Study 4/23/2018, v2

DRAFT Charter

A. Background

Large aperture telescopes benefit all astrophysics as well as planetary and Earth science. They provide unprecedented spatial resolution, spectral coverage, and signal to noise advancing all of these science areas. Envisioning the need for future large segmented telescopes to one day exceed the fairing size of existing or even planned launch vehicles, NASA will need to begin considering the in-space assembly (ISA) of these future assets. In addition, robotically assembling space telescopes in space rather than deploying them from single launch vehicles offers the possibility, in some circumstances, of reduced cost and risk for even smaller telescopes. This possibility, however, has not been proven. Therefore, following discussions within NASA's Science Mission Directorate (SMD) and Astrophysics Division (APD), the SMD Chief Technologist and APD Division Director have commissions study to assess the cost and risk benefits, if any, of the iSA of space telescopes. In particular, the study must answer the question: "When is it advantageous to assemble space telescopes in space rather than to build them on the Earth and deploy/them autonomously from individual launch vehicles?" This document charters the plan for the study deliverables, process, and membership. The goal for completion of the study is May 2019 culminating in a submitted whitepaper to the National Academies' 2020 Astronomy & Astrophysics Decadal Survey.

B. Deliverables

The in-Space Assembly Telescope (iSAT) Study Working Group is chartered by the NASA SMD Chief Technologist and APD Director to deliver by the goal of May 2019 a whitepaper assessing:

1. the telescope size at which iSA is necessary (an enabling capability)

2. the telescope size at which iSA is cheaper or lower risk with respect to traditional launch vehicle deployment (an enhancing capability)

* * *

Paul Hertz

Division Director NASA Astrophysics Division

Michael Seablom

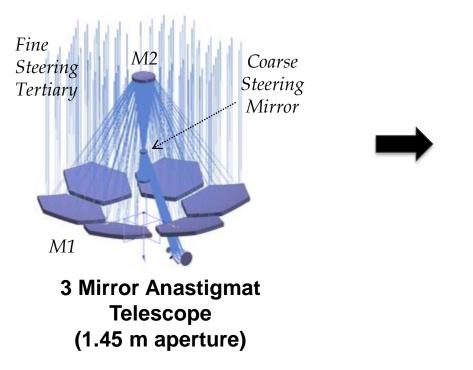
Directorate Chief Technologist NASA Science Mission Directorate

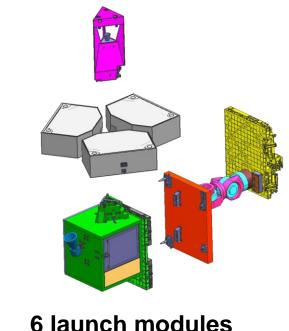
Activity 1a

Concept Design and Architecture for the iSAT

Select a reference <u>design and architecture</u> concept for a 20 m, filled aperture, non-cryogenic space telescope to be assembled and tested in space.

- Paradigm shift in architecture: Modularization
- An example, from the 2012 OpTIIX study (NASA JSC/GSFC/JPL/STScI):

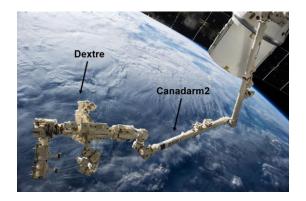


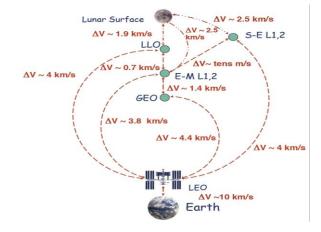


for assembly

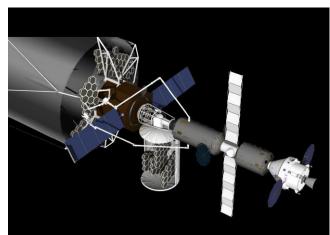
Activity 1b: Concept for Assembling and Testing the ISAT

Select a reference in-space <u>assembly and testing concept</u> for the "assemble-able" space telescope architecture, defining robotics, orbit, launch vehicle, and assembly platform.









Activities 2a and 2b Detailed Engineering Design and Costed

Activity 2a: Advance the engineering fidelity of the concepts sufficiently so that they can be costed.

- a) Inputs from Activity 1a and 1b
- b) Select a team of NASA engineers, academia, government labs, and commercial companies to conduct the work.
- c) Needs funding

Activity 2b: Estimate, through an independent body, the cost of designing, architecting, assembling, and testing the reference 20 m space telescope?

- a) Input design from Activity 2a
- b) Identify risks
- c) Parameterize the cost to smaller apertures

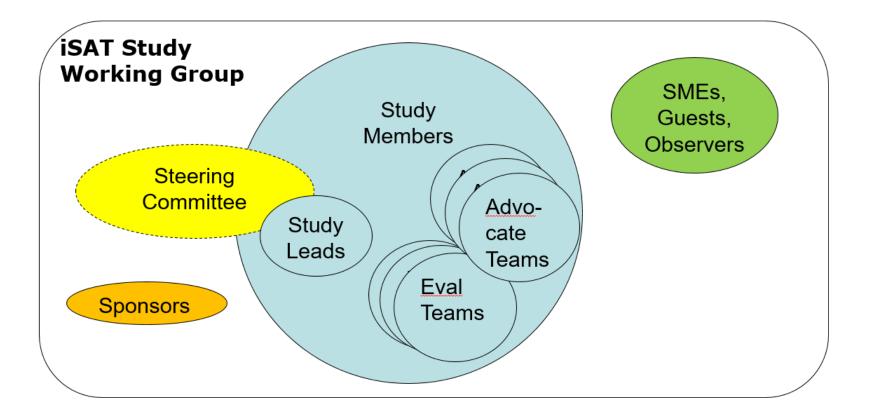
Activity 3 Deliver Final Whitepaper

Write and deliver the Final Whitepaper

a) Submit to APD Director who submits to 2020 Decadal Survey

Participants and Roles (US Persons Only)

The iSAT Study Working Group



Participants of the iSAT Study WG

Study Members (aim to come to consensus)

- Option Advocates, Technical Evaluation Team, Programmatic Evaluation Team will come out of this group.
- This first group of Study Members are preferentially selected and focused on Activity 1a (Telescope design and architecture concept; see below for the different activities)

<u>Name</u> Invited (C-confirmed: I-inv	Institution	<u>Expertise</u>
Invited (C=confirmed; I=invi 1. Joel Nissen (C)	NASA JPL	Matualagu
2. Gary Matthews (C)	Consultant	Metrology Mirror Segi
3. Fang Shi (C)	NASA JPL	
4. Larry Dewell (C)	Lockheed	WF Sensing
5. Oscar Salazar (C)	NASA JPL	Pointing, St
6. Phil Stahl (C)	NASA JPL NASA MSFC	Pointing, St Telescope A
7. Jon Arenberg (C)	Northrop	
	1	Telescope A
8. Doug McGuffey (C) 9. Kim Aaron (C)	NASA GSFC NASA JPL	Systems En
	NASA JPL NASA LaRC	Systems En
10. Sharon Jeffries (C) 11. Al <u>Tadros (</u> I)	SSL	Systems En Robotics
12. Joel Burdick (I)	Caltech	Robotics
13. Bob Hellekson (C)	Orbital-ATK	Telescope S
14. Gordon Roessler (C)		Robotics
		Optical Des
15. Michael Rodgers (C) 16. Hsiao Smith (C)	NASA JPL NASA GSFC	Robotics
17. Eric Mamajek (C)	NASA GSFC NASA ExEP	Astrophysic
, , ,	NASA EXEP NASA JPL	Optical Des
18. Shanti Rao (C)	NASAJPL	Optical Des
19. Ray Ohl (C)	NASA GSFC	Optical Alig
20. Sergio Pellegrino (C)	Caltech	Telescope S
21. Cal Ablanalp (I)	Harris	Telescope I
22. Tere Smith (C)	NASA JPL	I&T
23. Paul Backes (C)	NASA JPL	Robotics
24. Jim Breckenridge (C	Univ of Arizona	Optical Des
25. Alison Barto (C)	Ball	Optical SE/
26. Jeanette Domber (C)	Ball	SE/Structur
27. Joe Parrish (C)	DARPA	Robotic Sys
28. Acey Herrera (I)	NASA GSFC	I&T
29. Paul Dizon (I)	NASA GSFC	I&T
30. David Stubbs (C)	Lockheed	Telescope S
31. John Dorsey (C)	LaRC	Telescope S
32. David Yanatis (1)	Harris	Optical Syst
33. Rudra Mukherjee	NASA JPL	Robotics
,		

Aetrology Airror Segments VF Sensing/Control, Coronagraphy ointing, Stability, Control ointing, Stability, Control elescope Architecture elescope Architecture ystems Engineering systems Engineering/Structures ystems Engineering Robotics Robotics Celescope Systems Robotics Optical Design Robotics Astrophysicist Optical Design

ptical Alignment/Test elescope Structures 'elescope Design &Т lobotics ptical Design ptical SE/testing E/Structures/Instruments lobotic Systems &Т &Т 'elescope Structures/Design 'elescope Structures ptical Systems lobotics

Study Members – those who will be making recommendations for the Designing and Architecture concepts (Activity 1a)

Expect changes for the Assembling and Testing concepts (Activity 1b)

Participants of the iSAT Study WG

Steering Committee (C=confirmed)

1. E	Dave Redding (C)	NASA JPL	Study Member (mirror segments, WFSC)
2. J	oe Pitman (C)	consultant	Study Member (opto-mech structures)
3. S	Scott Knight (C)	Ball Aero	Study Member (optical design)
4. E	Bill Doggett (C)	NASA LaRC	Study Member (telescope structures)
5. N	Matthew Greenhouse (C)	NASA GSFC	Study Member (astrophysicist)
6. J	oanne Hill-Kittle (C)	NASA GSFC	
7. F	Ron Polidan (C)	consultant	Study Member (telescopes)
8. J	ohn Grunsfeld (C)	NASA (ret)	
9. k	Keith Belvin (C)	NASA STMD	
10. E	Brad Peterson (C)	STScI/OSU	Study Member (astrophysicist)
11. F	Florence Tan (C)	NASA SMD	
12. F	Ray Bell (C)	Lockheed	Study Member (telescope systems)
13. N	Nasser Barghouty (C)	NASA APD	-
14. E	Eric Smith (C)	NASA APD	

Some are also Study Members because of their technical expertise

Subject Matter Experts, Observers, and Guests:

40. Lynn Bowman (C)

- 41. Keith Warfield (I)
- 42. Rich Rynders (C)
- 43. Howard MacEwen (C)
- 44. Brendan Crill (C)

NASA LaRC NASA ExEP

Orbital-ATK

Reviresco NASA JPL (Organizing Committee) Can add more people here; not "consensus members"

Consensus

Drawn from NASA Policy

Consensus decisions

- May produce more durable decisions than those by votes or decree.
- However, convergence time can be a factor.
- We adopt a Constrained Consensus method defined as: Strive for consensus in the reasonable time available, else, the leaders make a decision. Dissent (if any) is captured and the groups moves on with full support of the decision.

• Follow 7120.5E, Ch 3.4, "Process for Handling Dissenting Opinion"

- Three options:
 - (1) Agree,
 - (2) Disagree but fully support the decision,
 - (3) Disagree and raise a dissenting opinion
- Treat (1) and (2) as consensus for iSAT Study Working Group
- Dissents (3) if any will be documented and delivered to the Study Leads and to the Sponsors

Role of the Steering Committee

- 1. Recommend membership in the Study Working Group
 - Ensuring they are well represented in terms of expertise.
- 2. Advise the Study Leads, providing feedback at key junctions of the Study regarding its progress and direction in moving the work forward.
- 3. Provide input regarding the Study's Assumptions and Initial Conditions:

Initial Conditions

- a) Filled aperture, non-cryogenic UV/O/NIR telescope
- b) Coronagraph-related requirements on telescope structures and optics

Assumptions

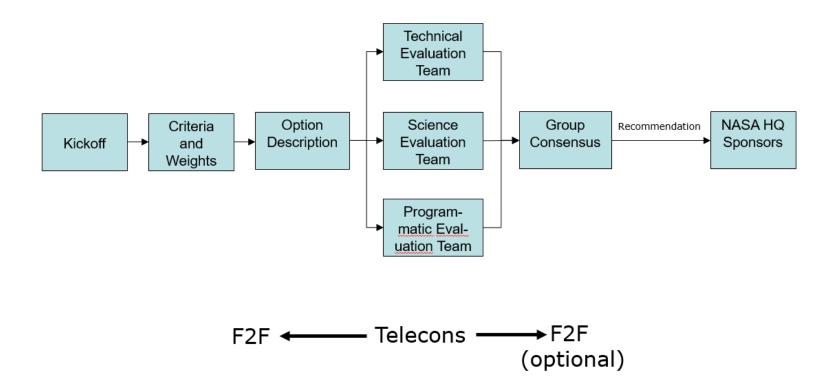
a) Operational destination is SEL2

4. Telecons as needed but before all milestones and critical junctures

1. No surprises

Process

How will iSAT Study WG Produce a Recommendation?



• Examples of Recommendations following this Trade Process:

- WFIRST Coronagraph: http://wfirst.gsfc.nasa.gov/science/AFTA_Coronagraph_Arch_Selection/Coronagraph_Downselect_Rec_Dec13_2013.pdf
- Starshade Readiness Working Group: https://exoplanets.nasa.gov/system/internal_resources/details/original/339_SSWG_APD_briefing_final.pdf
- Starshade Mechanical Deployment Trade Study

Features of Kepner-Tregoe Decision Process

Systematic Decision Making

Decision Statement											
u					Opti	on 1	Opti	ion 2	Opti	ion 3	
Description		Featu	re 1								
Scri		Featu	re 2								
De		Featu	re 3								
	Musts										
		M1				•		/		/	
_		M2				•		?		?	
Evaluation		М3			✓		~		×		
lua	Wants	;	Weights								
Eva		W1	w1%		Rel s	core	Rel s	core	Rel s	core	
		W2	w2%		Rel s	core	Rel s	core	Rel s	core	
	_W3 w3%					Rel score		Rel score		Rel score	
			100%	Wt sum =>	Sco	re 1	Score 2		Score 3		
	Risks				С	L	С	L	С	L	
		Risk 1			М	L	М	L			
		Risk 2			Н	Н	М	М			
Final	Final Decision, Accounting for Risks										
					C = Con	sequend	e, L = Lil	kelihood	ł		

A little consensus at a time

Process Overview

٠

- Agree on Evaluation Criteria and Weights
- Document Options and Description
- Evaluate Options vs Criteria
- Reach Consensus on Evaluation
- Document Risks,
 Opportunities
- Recommendation accounting for Risks, Opportunities

Example of a Completed Trade Matrix

				Option	n1	Option 2	Option 3	Option 4	Option 5	Option 6	Notes
	Name			SPC		PIAACMC	HLC	VVC	VNC - DA	VNC - PO	
Mu	usts	Programmatic									
	M1 - T	Science: Meet Threshold requirements? (1.6, x10)			Yes	Yes	Yes	No	No	U	
	M2	Interfaces: Meets the DCIL**?			Yes	Yes	Yes	Yes	Yes	U	
		TRL Gates: For baseline science is there a credible									 yes, or expected likely unknown
	M3	plan to meet TRL5 at start of FY17 and TRL6 at start			Yes	Yes	Yes	U	No	U	x no, or expected showstopper
_		of FY19 within available resources?									
	M4	Ready for 11/21 TAC briefing			Yes	Yes	Yes	Yes	Yes	No	
	M5	Architecture applicable to future earth- characterization missions			Yes	Yes	Yes	Yes	Yes	U	
_		characterization missions									
Wa	ants		Weights	SPC		PIAACMC	HLC	vvc	VNC-DA	VNC - PO	
	W1	Science	40								
											Range of opinions between "significant and small". For S
	а	Relative Science yield (1.6, x10) beyond M1-T		Sr	m/Sig	Best	Sm/Sig	VL	VL		and VNC2 the search area is ~3 times less than 360deg, an
											that was taken into acct in comparisons
	W2	Technical	30								
	а	Relative demands on observatory (DCIL), except		E	Best	Best	Best	Best	Small		
		for jitter and thermal stability Relative sensitivities of post-processing to low									For n-lambda over D or different amplitudes the designs
	b	order aberrations		E	Best	Sig	Sig	VL	U		have the same relative ranking
	с	Demonstrated Performance in 10% Light		s	Small	Sig	Best	Sig	VL		Demonstrated Performance (10%) and Prediction
	d	Relative complexity of design			Best	Small	Best	Small	Sig		
	e	Relative difficulty in alignment, calibration, ops		E	Best	Small	Best	Small	Sig/Sm		Identify "Best" and others are: -Wash
	W3	Programmatic	30								-Small Difference
	а	Relative Cost of plans to meet TRL gates		E	Best	Small	Best	Sig	Sig		-Significant Difference
		Wt. sum =>	100%								-Very Large Difference
		Wa Sull ->	10070								
Ris	ks	(all judged to be Hgh consequence)		SPC		PIAACMC	HLC	vvc	VNC-DA	VNC - PO	
				С	L	C L	C L	C L	C L	C L	
	Risk 1	Technical risk in meeting TRL5 gate]			м	M/L	м/н	н		PIAA trend over the last three working days lower, but
	NISK 1				L.	IVI	IVI/ L				recommendation to keep M
	Risk 2	Schedule or Cost risk in meeting TRL5 Gate			L.	м	M/L	м/н	н		
	Risk 3	Schedule or Cost risk in meeting TRL6 Gate			L	L	L.	м	M		
	Diele 4	Diele of a stars still satisfy a star stars						н	н		
	Risk 4	Risk of not meeting at least threshold science			L.			-			
	Risk 5	Risk of mnfr tolerances not meeting BL science			L.	L	L	M/L	н		One dissent, previous TDEM performance track record an
		Risk that wrong architecture is chosen due to									Bala's assessment should be taken into account.
	Risk 6	assumption that all jitter >2Hz is only tip/tilt			L.	M/H	м	M/H	м		
		Risk that wrong architecture is chosen due to any									
	Risk 7	assumption made for practicality/simplicity		0	pen en	ided question, s	pawned evaluati	ons on Risk 5, Ri	sk 6, Risk 8, and C	Oppty 1	
		Risk that ACWG simulations (by JK and BM)									Model validation is a risk that needs to be evaluated in t
	Risk 8	overestimate the science yield due to model			discu	ussed; not enoug	gh understanding	at this time to n	nake an evaluatio	on.	future
		fidelity									
	unities	(judged to be High benefit)		SPC		PIAACMC	HLC	vvc	VNC-DA	VNC - PO	
	unities	Gadeo to se tilgi belienti				B L	B L				
port				В	L	B L	B L	B L	B L	B L	1
port	Oppty 1	Possibility of Science gain for 0.2marcsec jitter, x30			L	М/Н	м	L	н		
port											
port											
-	cicion - A	ecounting for Picks and Opportunit			_						
-	ecision, A	Accounting for Risks and Opportunit	ies:								
-	ecision, A	accounting for Risks and Opportunit	ies:			C = Conseque	nce, L = Likeliho	od, B=Benefit			indicates those few areas where consensus was not ach

Next Steps

Next Steps

- Kick-Off Telecon with the entire Working Group
 - Monday and Tuesday (5/7 and 5/8)
- Subsequent Telecons with the entire Working Group
 - Bi-weekly cadence
 - Advancing work

• First Face-to-Face Workshop for the Working Group

- June 5-7 at Caltech
- Focus is on Activity 1a: Designing and Architecting a Modularized Telescope
- Draft Agenda completed being sent out today
- Breakout sessions

Draft Schedule Under Review

